

19 **Abstract**

20 A new automated method to retrieve charge layer polarity from flashes, named Chargepol, is
21 presented in this paper. Using the RELAMPAGO field campaign NASA Lightning Mapping
22 Array (LMA) data deployed in Cordoba, Argentina, from November 2018 to April 2019, this
23 method estimates vertical charge layer polarity, altitude, and depth from VHF-based
24 observations of lightning flashes and, when extended for long periods of time, infers charge
25 structure for thunderstorms' entire life cycle. This method provided reliable charge retrievals as
26 demonstrated in validation when assigning VHF lightning source polarity manually. Examples of
27 Chargepol applied to normal and anomalous charge structure storms in Central Argentina during
28 RELAMPAGO are presented for the first time. Application of the algorithm to months of LMA
29 data in Central Argentina and several locations in the United States allowed for the
30 characterization of the charge structure in these regions and for a reliable comparison using the
31 same methodology. About 13.3% of Cordoba thunderstorms presented an anomalous charge
32 structure, slightly higher than in Oklahoma (12.5%) and West Texas (11.1%), higher than
33 Alabama (7.3%) and considerably lower than in Colorado (82.6%). Some of the Cordoba
34 anomalous thunderstorms presented enhanced low-level positive charge, a feature rarely if ever
35 observed in Colorado thunderstorms.

36 **1 Introduction**

37 This study aims to characterize the charge structure in Central Argentina region for the
38 first time, utilizing a large dataset, as it is a key science goal of the RELAMPAGO (Remote
39 sensing of Electrification, Lightning, And Mesoscale/microscale Processes with Adaptive
40 Ground Observations) field campaign (Nesbitt et al., 2021). This is achieved through a new
41 automated method to retrieve thunderstorm charge layer polarity using Lightning Mapping Array
42 (LMA) source and flash data, which is described in this paper. Past studies have associated the
43 severity of thunderstorms with patterns in charge distribution (Wiens et al., 2005; Fuchs et al.,
44 2018). The dominant meteorological environment provides initial conditions that would
45 influence the kinematics and microphysics within thunderstorms, which in turn affects its charge
46 structure and dominant cloud-to-ground lightning (CG) polarity. Relatively few studies have
47 documented the charge structure over continents other than North America (Lopez et al., 2019;
48 Pawar & Kamra, 2004; Pineda et al., 2016; Qie et al., 2005). Furthermore, documenting charge
49 structure in regions such as Argentina, which has perhaps some of the highest flash rate
50 thunderstorms in the world (Zipser et al., 2006), is of crucial importance.

51 Due to the nature of lightning processes and their characteristic emission in the VHF,
52 thunderstorm charge structure associated with flashes can be inferred from LMA observations
53 (Lang & Rutledge, 2011; Rust et al., 2005; Wiens et al., 2005). An intra-cloud lightning (IC)
54 flash initiates in a region with strong electric field, in between regions of charge with opposite
55 polarities. Initially, bi-directional leaders form and move in opposite directions: a positive leader
56 moves to a region of net negative charge, and a negative leader moves to a region of positive
57 charge in the cloud (Kasemir, 1960). Often when a leader reaches a charge layer, it propagates
58 horizontally through the charge layer away from the flash initiation location (Shao & Krehbiel,
59 1996). Flashes propagating through charge regions that constitute a vertical dipole with positive
60 charge located above negative charge, is referred to as positive cloud flash (+IC), while flashes
61 that propagate through negative over positive dipoles are named negative cloud flash (-IC,
62 Bruning et al., 2014). K-processes may also occur, transporting charge to the base of the initial

63 channel (Shao & Krehbiel, 1996). Based on knowledge of radiation propagation by lightning,
64 VHF-based sensors primarily detect radiation from negative breakdown of lightning that
65 propagates through regions of positive charge (Mazur & Ruhnke, 1993; Rison et al., 1999).
66 Then, mapping of VHF sources is used to manually determine the location of positive and
67 negative charge layers (Bruning et al., 2007; Lang and Rutledge, 2008; Rust et al., 2005; Wiens
68 et al., 2005). Observations of the VHF source distribution in height for long periods of time are
69 used to infer the location of charge regions, as the altitude with most sources are often associated
70 with positive charge layers (Fuchs et al., 2018; Fuchs & Rutledge 2018; Lang et al., 2020; Lang
71 & Rutledge, 2011). Tessendorf et al. (2007b) infer charge layer polarity automatically by using
72 the first LMA source altitude for a flash, and the number of sources above and below that
73 altitude. Stough and Carey (2020) utilized the DBSCAN (Density-Based Spatial Clustering of
74 Applications with Noise, Ester et al., 1996) algorithm to identify regions of dense sources and
75 infer charge region polarity. Electric field soundings have been deployed to infer polarity of
76 charge regions within and nearby thunderstorms (Marshall et al., 1995; Rust & MacGorman,
77 2002; Stolzenburg et al., 1998), and have been compared to LMA-inferred charge regions (Rust
78 et al., 2005). In order to interpret electric field dataset with altitude, the Gauss' Law
79 approximation is assumed, where the charge density is proportional to the electric field variation
80 with height (Stolzenburg et al., 1998).

81 In order for clouds to build regions of net charge polarity and become electrified, the non-
82 inductive charging (NIC) mechanism is thought to dominate, which does not require a pre-
83 existing electric field to polarize cloud and precipitation size particles. In the NIC mechanism,
84 the polarity that graupel particles acquire when colliding with ice crystals in the presence of
85 supercooled liquid water (Saunders et al., 1991; Takahashi, 1978) depends on the temperature,
86 and the effective liquid water content (EWC, the accreted fraction of the liquid water content).
87 High (low) temperature and large (low) EWC are associated with graupel charging positively
88 (negatively), and crystals charging negatively (positively) (Berdeklis & List, 2001; Pereyra et al.,
89 2000; Saunders et al., 1991, 2001; Saunders & Peck, 1998; Takahashi, 1978). Particle
90 differential fall speeds and updrafts lead to storm-scale charge separation, with crystals being
91 transported upward to cloud tops, and graupel residing in the mid-level, mixed-phase layer,
92 forming two main regions of charge during the developing-to-mature stage of thunderstorms
93 (Williams, 1985).

94 A thunderstorm with an upper level net negative charge, and a mid-level, mixed-phase
95 layer with net positive charge characterizes an anomalous charge structure, as found in
96 thunderstorms during the STEPS field campaign conducted in Kansas, Colorado and Nebraska
97 (MacGorman et al., 2005; Rust et al., 2005; Rust & MacGorman, 2002; Tessendorf et al., 2007a;
98 Tessendorf et al., 2007b; Weiss et al., 2008; Wiens et al., 2005). They have also been observed in
99 thunderstorms in Oklahoma by Marshall et al. (1995) and Emersic et al. (2011), during the
100 TELEX field campaign (MacGorman et al., 2008), in Texas (Chmielewski et al., 2018), Alabama
101 (Stough & Carey, 2020), and Spain (Pineda et al., 2016). Storms with a normal charge structure
102 would have dominant net negative charge in the mixed-phase layer, and net positive above, as
103 demonstrated in early foundational studies reviewed by Williams (1985), in the in-situ aircraft
104 studies by Dye et al. (1986, 1988, 1989), during TELEX (Bruning et al., 2007) and STEPS
105 (Weiss et al., 2008) field campaigns, among others. A low-level charge layer with opposite
106 polarity to the nearest charge region is occasionally present (Lopez et al., 2019; Pawar & Kamra,
107 2004; Williams, 1989) and, if positive and abnormally large, may also be termed anomalous
108 (Bruning et al., 2014; Fuchs et al., 2015; Qie et al., 2005). Some events can have multiple charge

109 regions, such as mesoscale convective systems (MCSs) (Lang & Rutledge, 2008; Lund et al.,
110 2009; Stolzenburg et al., 1998), multicell storms (Bruning et al., 2007), and supercells (Bruning
111 et al., 2010; Calhoun et al., 2013; Wiens et al., 2005).

112 Fuchs and Rutledge (2018) analyzed a large lightning flash dataset for isolated cells in
113 four different regions in the United States, and observed that Colorado storms have a prevalence
114 of anomalous charge structure compared to other regions, as Colorado's highest flash rate mode
115 was observed at lower levels (warmer temperatures and higher radar reflectivity values) than in
116 other regions. In addition, they suggested that Colorado is followed by Oklahoma in terms of
117 anomalous storm frequency, followed by Alabama and Washington D.C. with rare anomalous
118 observations. A large occurrence of positive cloud-to-ground lightning (+CG) is often associated
119 with anomalous charge structure storms, as a main net positive charge region is at the middle or
120 low levels of a storm instead of near its top, facilitating the propagation of positive leaders
121 toward the ground, especially if a small opposite (negative) charge region is present at lower
122 levels. Orville and Huffines (2001) found that the percentage of +CGs in the United States varies
123 from 2% in Florida to 10-20% in a region extending from the High Plains of Eastern Colorado to
124 the Upper Midwest. In the central and north Great Plains, a high percentage (>50%) of severe
125 storm reports were found to be associated with predominantly +CG lightning (>50% of CGs
126 being positive), when compared to southern Great Plains and eastern United States (Carey et al.,
127 2003).

128 Southeast South America has among the most severe thunderstorms in the world in terms
129 of high flash rate (Zipser et al., 2006), hail size (Cecil & Blankenship, 2012), heavy
130 precipitation, and flash floods (Rasmussen et al., 2014). Lightning characteristics have only been
131 documented using LMA data recently in this region (Lang et al., 2020), and the distribution of
132 charge within Argentina thunderstorms is explored for the first time in great detail in this study.
133 The general charge structure is estimated for a large dataset with a new algorithm, allowing for
134 the inference of the likelihood of normal and anomalous charge structure. Similar to Tessendorf
135 et al. (2007b) and Stough and Carey (2020), this method automatically infers charge polarity
136 from flashes, more closely resembling Tessendorf et al. (2007b) method but with improved
137 procedures, better emulating the steps that a human expert would perform when assigning
138 polarity to LMA sources for a flash by detecting the negative leader in a bi-directional model and
139 assigning polarity to sources of a flash (e.g., Rust et al. 2005). In this study, if a given lightning
140 flash passes a series of conditions, an algorithm analyzes its source location and time in order to
141 produce an estimate of charge layer polarity for that flash. This method has the capability to be
142 quickly applied to a large number of lightning flashes in a large LMA dataset (e.g., a few
143 minutes to process 24 hours of LMA flash level data within 100 km of the network center),
144 which allows for the inference of the general charge structure and its evolution in time for a
145 thunderstorm or for a large area of interest, as demonstrated by examples in this paper. The new
146 algorithm infers three-dimensional charge distribution on the flash level but its output is
147 simplified to vertical charge layer profiles for the science applications highlighted in this study.
148 Hence, output of this method is similar to manual assignment of polarity, providing positive and
149 negative layer altitude and depth, but it is much less labor extensive. This algorithm provides a
150 detailed inference of the charge layer distribution in the vertical, including altitude and depth of
151 negative charge layers, which is often not possible to be analyzed from the VHF source
152 distribution analysis, a method in which positive charge altitude is inferred from its peak
153 distribution. Lastly, this paper will present a detailed application of the new charge layer polarity
154 algorithm by characterizing the charge structure of Central Argentinian thunderstorms by

155 processing a large multi-month sample of LMA observations for the first time. The algorithm
156 performance is then further demonstrated through its application to multi-month LMA datasets
157 from several locations in the United States in which charge structure has already been
158 documented using the LMA-based charge layer retrieval techniques discussed above. The
159 additional application herein allows the charge structure of Central Argentinian thunderstorms to
160 be compared for the first time to several well-studied locations in the United States such as
161 Colorado, Oklahoma, West Texas and Alabama using the same algorithm.

162 This paper is organized as follows: Section 2 shows the lightning network datasets used
163 for method development and validation, section 3 describes the charge layer polarity
164 identification method and the performed validation, section 4 shows examples of the method
165 applied to Argentinian thunderstorms, section 5 shows a charge structure climatology during
166 RELAMPAGO over Central Argentina and compares it to four additional locations in the United
167 States representing a variety of charge structure climates (Fuchs and Rutledge, 2018), and
168 section 6 presents summary and discussion.

169 **2 Lightning Networks Deployed During RELAMPAGO and DC3**

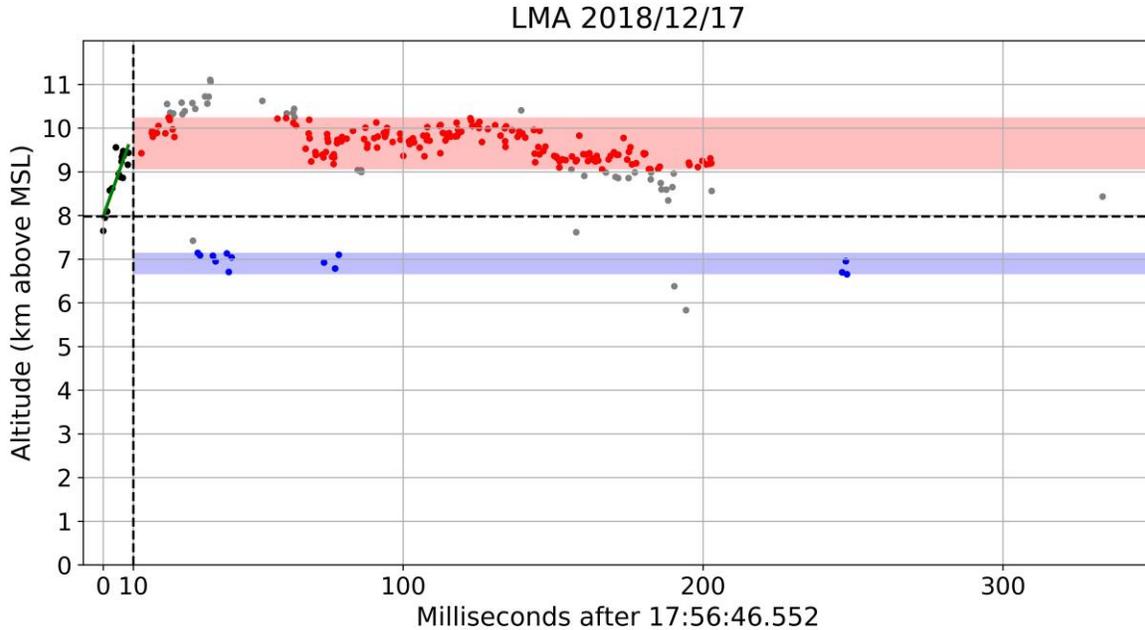
170 The Lightning Mapping Array (LMA) is a GPS-based network (Goodman et al., 2005;
171 Koshak et al., 2004; Krehbiel et al., 2000; Rison et al., 1999) that operates in the VHF
172 electromagnetic spectrum (Krehbiel et al., 2000), in which radiation events detected are often
173 associated with lightning breakdown processes (Rison et al., 1999). A VHF source location and
174 time is found by calculating the time-of-arrival equation (Koshak et al., 2004; Koshak &
175 Solakiewicz, 1996; Lhermitte & Krehbiel, 1979) best solution using a χ^2 goodness-of-fit function
176 iteration lower than 5, and minimum number of sensors detecting a source equal to 6. Lightning
177 flash datasets were obtained using the lmatools Python package (Bruning et al., 2015) that is
178 based on the DBSCAN (Ester et al., 1996) algorithm, a machine-learning procedure used for
179 clustering sources into flashes. DBSCAN searches for random sources, and groups them
180 considering space and time dimensions simultaneously. Distance and time thresholds between
181 sources are 3000 meters and 150 ms respectively, with a maximum flash duration of 3 seconds
182 (Fuchs et al., 2016).

183 As part of the RELAMPAGO field campaign (Nesbitt et al., 2021), an LMA of 11
184 sensors was deployed by NASA Marshall Space Flight Center to the eastern side of the Sierra de
185 Cordoba mountains in the province of Cordoba, Central Argentina, from November 2018 to
186 April 2019 (Lang et al., 2020). In this study, LMA datasets from the DC3 (Deep Convective
187 Clouds and Chemistry, Barth et al., 2015) field campaign are used to independently estimate the
188 charge structure in a variety of climatological regimes of the United States and compare the
189 algorithm results with other studies using different retrieval algorithms and with the Cordoba
190 region using the same algorithm. During the DC3 field campaign, LMA networks were deployed
191 simultaneously in Alabama, West Texas, Oklahoma, and Colorado in May and June 2012 (Barth
192 et al., 2015; DiGangi et al., 2016; Mecikalski et al., 2015). Only flashes with centroid location
193 within the 100 km range distance from the LMA network center are being considered in this
194 study, as altitude errors are expected to be smaller and the flash detection efficiency to be higher
195 (Chmielewski & Bruning 2016; Koshak et al. 2004; Lang et al., 2020; Thomas et al. 2004)
196 within this range. Also, flashes with less than 20 sources are not considered in this study (more
197 in section 3).

198 **3 Description of the Charge Layer Polarity Identification Method**

199 The charge layer polarity identification method (hereafter Chargepol) consists of an
200 automated algorithm that applies a series of procedures to each lightning flash retrieved by the
201 Imatools, in order to infer charge layer polarity from a flash. For reference, Figure 1 shows a
202 flash example with the procedures illustrated. Firstly, flashes with less than 20 sources are
203 disregarded because those flashes would not allow a sufficient number of sources to characterize
204 the initial negative leader breakdown, negative leader propagation through a positive charge
205 region, and sources associated with a negative charge region. Then, all sources contained in the
206 first 10 ms of a flash, referred to here as the Preliminary Breakdown sources (PB sources) for
207 simplicity, are analyzed. A minimum of 3 PB sources is required, and the time interval between
208 the first and last PB source has to be at least 4 ms, in order to better characterize the initial
209 vertical motion of the negative leader. Duration between 4 and 10 ms follow typical duration
210 periods for PB (Zheng et al., 2019). We make the assumption that PB sources are associated with
211 negative breakdown having a predominant vertical motion toward a region of positive charge
212 (Shao & Krehbiel, 1996). Hence, linear regression is applied to the PB sources time-height
213 dimension. The linear regression slope is used as a proxy for the vertical speed of the leader, and
214 has to be greater than a threshold of absolute value of 0.05 (0.5 km height variation in 10 ms),
215 which is equivalent to a vertical speed of $5 \times 10^4 \text{ ms}^{-1}$, or half the typical order of magnitude speed
216 of a negative leader (Shao & Krehbiel, 1996; van der Velde & Montanya, 2013). By applying
217 that slope threshold, flashes with no clear initial vertical motion are discarded, facilitating further
218 a correct depiction of charge region polarity. In addition, the linear regression performed on the
219 PB sources needs to have a mean squared error (MSE) lower than a threshold of 0.25, in order to
220 avoid linear regression that included noise sources in its solution.

221 Only flashes that satisfy all the aforementioned conditions are used for charge layer
222 depiction, which is typically about 16% of all flashes (more on this in section 5). The fact that
223 not all flashes are analyzed does not interfere with the objective of this study, because estimating
224 charge polarity for some flashes is sufficient to determine the charge structure evolution over
225 long periods of many hours, as demonstrated in the next section. Next, non-PB sources are used
226 to infer charge layer polarity, altitude and depth. The PB linear regression intercept altitude is
227 used as a threshold, referred to here as the Charge Height Threshold (CHT), in order to separate
228 positive and negative charge layers candidate sources. For a positive PB linear regression slope
229 (i.e., a flash with initial negative leader moving upward), all non-PB sources above (below) the
230 CHT are candidate sources to define a positive (negative) charge layer. A flash with initial
231 downward motion (negative PB linear regression slope) would have all non-PB sources below
232 (above) the CHT as candidate sources for positive (negative) charge layer. Then, among the
233 candidate sources for each layer polarity, the interval between the 10th and the 90th percentile
234 source heights is used to define a charge layer. For some flashes, it is possible that only one layer
235 polarity is estimated, which leads to the total number of estimated positive layers from flashes
236 for a large period of time being slightly larger than the number of estimated negative layers from
237 flashes.



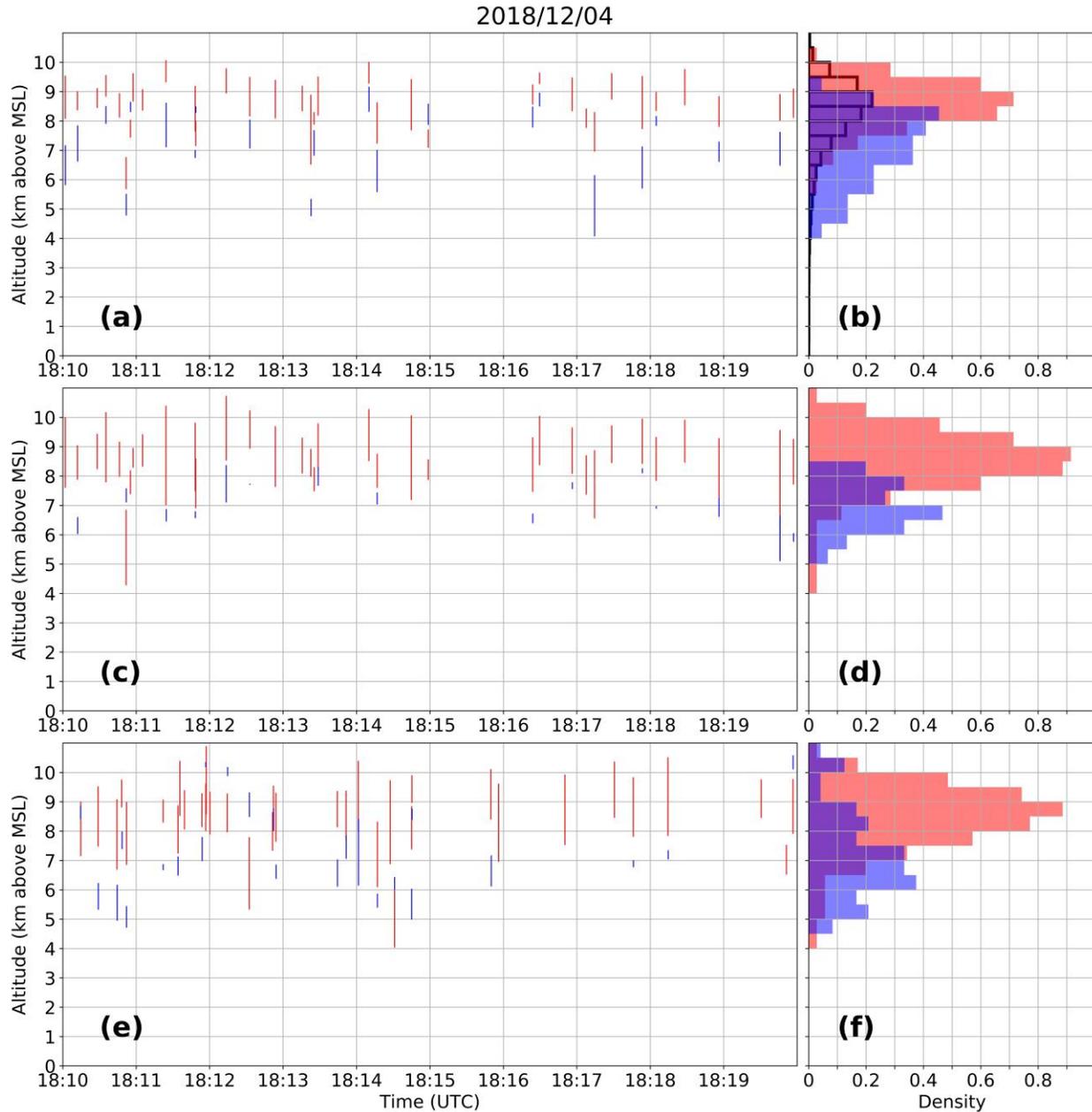
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239 **Figure 1.** A time-height plot for a positive intracloud flash. The vertical dashed line marks the 10
 240 ms time limit that defines the PB sources (black dots). The green line is the linear regression fit
 241 on the PB sources. The horizontal dashed line is the CHT (Charge Height Threshold), that
 242 separates candidate sources for positive and negative charge layers. Red and blue dots (and
 243 shaded areas) define the positive and negative charge layers altitudes and width for this flash,
 244 found by applying the interval between the 10th and 90th percentile source altitudes for each
 245 polarity candidate sources. Gray dots are candidate sources outside the 10th-90th percentile
 246 interval, which were not used to define charge layers.

247 3.1 Validation using Manual Analysis of LMA

248 In order to validate the automated Chargepol identification method, manual polarity
 249 inference (Rust et al., 2005; Wiens et al., 2005) was performed on some lightning flashes, and
 250 compared with the Chargepol algorithm output. First, a 10-minute period with a predominance of
 251 normal charge structure (i.e., normal dipole with positive charge over negative charge) was
 252 chosen from the RELAMPAGO LMA dataset. Among the 168 flashes that occurred in this
 253 period, the algorithm estimated charge layers from 35 of them (21%) (Figure 2a). Figures 2b
 254 shows a histogram density with the altitude where each charge layer polarity was detected (a
 255 peak of positive polarity of 0.7 between 8.5 and 9 km height means that 70% of all positive
 256 charge occurred at that level). Hence, source polarities were manually assigned for the same 35
 257 flashes, shown in Figures 2c and 2d. The positive charge altitude was estimated to be between
 258 about 8 km and 9.5 km from both Chargepol (Figure 2b) and the manual method (Figure 2d).
 259 Manual depiction of negative charge (Figure 2d) proved to be challenging, as it could not be
 260 estimated from all lightning flashes. Even so, it is notable that the negative charge layer is
 261 located at altitudes generally below the altitude of positive charge, with peak occurrence between
 262 6.5 and 7 km height (Figure 2d). Additional validation was performed by assigning polarity for
 263 another 35 randomly chosen flashes among the 133 flashes during the same 10-minute period
 264 that were not considered by Chargepol (Figure 2e). Most of these flashes did not have a clear

265 vertical trend of the initial leader (not shown). However, as shown in Figure 2f, most positive
 266 charge layer detection from flashes were estimated to be between 8 and 9.5 km, consistent with
 267 the automated method (Figure 2b), while negative charge is located at lower altitudes. The
 268 analysis on an independent subset of flashes from Figures 2e-f demonstrates that Chargepol
 269 analysis on a fraction of total flashes is sufficient for charge structure analysis.

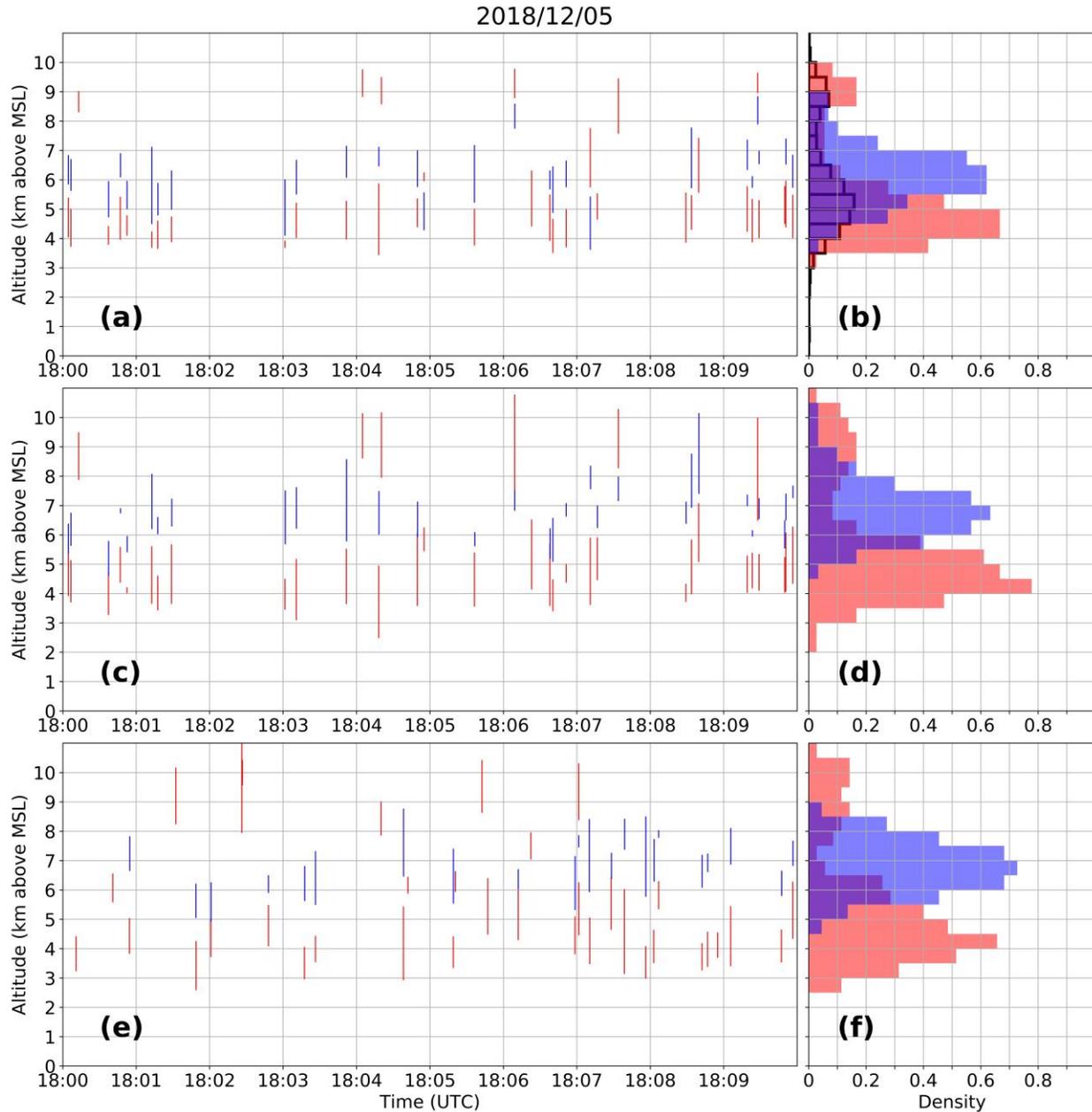


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271 **Figure 2.** (a) Charge layers estimated from 35 flashes using the Chargepol automated method,
 272 (c) polarity assigned manually for the same 35 flashes considered by Chargepol, and (e) polarity
 273 assigned manually for other 35 flashes not considered by Chargepol during the same time period.
 274 Each red (blue) vertical line represents a positive (negative) charge layer estimated from a flash.
 275 (b), (d), and (f) shows histograms with the density of retrieved positive and negative charge
 276 layers with height for (a), (c) and (e), respectively. Black histogram in (b) shows the source

277 height distribution. The 10-minute period chosen had a predominance of normal charge structure
278 as clearly shown by the Chargepol algorithm, manual analysis, and source distribution.

279 This procedure was repeated for a 10-minute period with a predominance of anomalous
280 charge structure (dipole with positive charge located below negative charge), shown in Figure 3.
281 During this period, a high flash rate storm produced mostly negative ICs propagating through a
282 lower positive charge. Another storm with low flash rate and upper positive charge layer was
283 active at the same time. A total of 107 flashes occurred during this period, in which Chargepol
284 estimated charge layers for 36 of them (Figure 3a-b). Manual depiction of charge polarity for
285 these same 36 flashes (Figure 3c) show that altitude of positive and negative charge (Figure 3d)
286 is in agreement with Chargepol, although manual inference of negative charge is at slightly
287 higher altitude. From Figure 3d, more than 50% of negative charge layers occurred in altitudes
288 from 6 to 7.5 km, while Chargepol estimated negative charge layers from 5.5 to 7 km height
289 (Figure 3b). The small differences in charge layer altitudes between the manual and automated
290 method demonstrate the small uncertainty of the method. Manual inference for a different set of
291 36 flashes during the same time period that was not considered by Chargepol is shown in Figures
292 3e and 3f, and it is consistent with other flashes (Figures 3a and 3c) in locating lower positive
293 charge and mid-level negative charge. The altitude distance between positive and negative
294 charge layers centers (from histogram plots) for all methods is about 2 km. The few upper
295 positive charge layers located above 8.5 km by all methods are from the normal charge structure
296 storm aforementioned.



297

298 **Figure 3.** Same as in Figure 2, but for a 10-minute period with predominance of anomalous
 299 charge structure.

300 The manual depiction of charge layers polarity agrees qualitatively well with the
 301 automated depiction. In general, the vertical distance between maximums for each polarity, as
 302 seen in the histograms plots of Figures 2 and 3, is sufficiently large and well separated by more
 303 than 1 km, leading to charge layers being well identified in the vertical dimension.

304 **3.2 Validation using Vertical Distribution of VHF Sources**

305 An additional method to validate the Chargepol algorithm is the estimate of the positive
 306 charge layer altitude from the peak in the VHF source histogram (Fuchs et al., 2018; Fuchs &

307 Rutledge 2018; Lang et al., 2020; Lang & Rutledge, 2011). Figures 2b and 3b display an
308 additional histogram with the vertical source density. The histogram for the normal case (Figure
309 2b) presents the peak at the same altitude the Chargepol method shows a peak with the most
310 occurrences of positive charge. A comparison of these two methods shows that the Chargepol
311 method has the advantage of inferring negative charge layer altitude, which is not possible to
312 estimate from the LMA VHF source distribution. For the anomalous case (Figure 3b), the main
313 low-level peak from the anomalous charge structure storm and the secondary peak from the
314 normal storm are depicted. The peak from the source histogram is at a slightly higher altitude, 5
315 to 5.5 km, compared to Chargepol's positive inference at 4 to 5 km. However, both methods
316 generally agree and the depiction of the negative layer by Chargepol is notable.

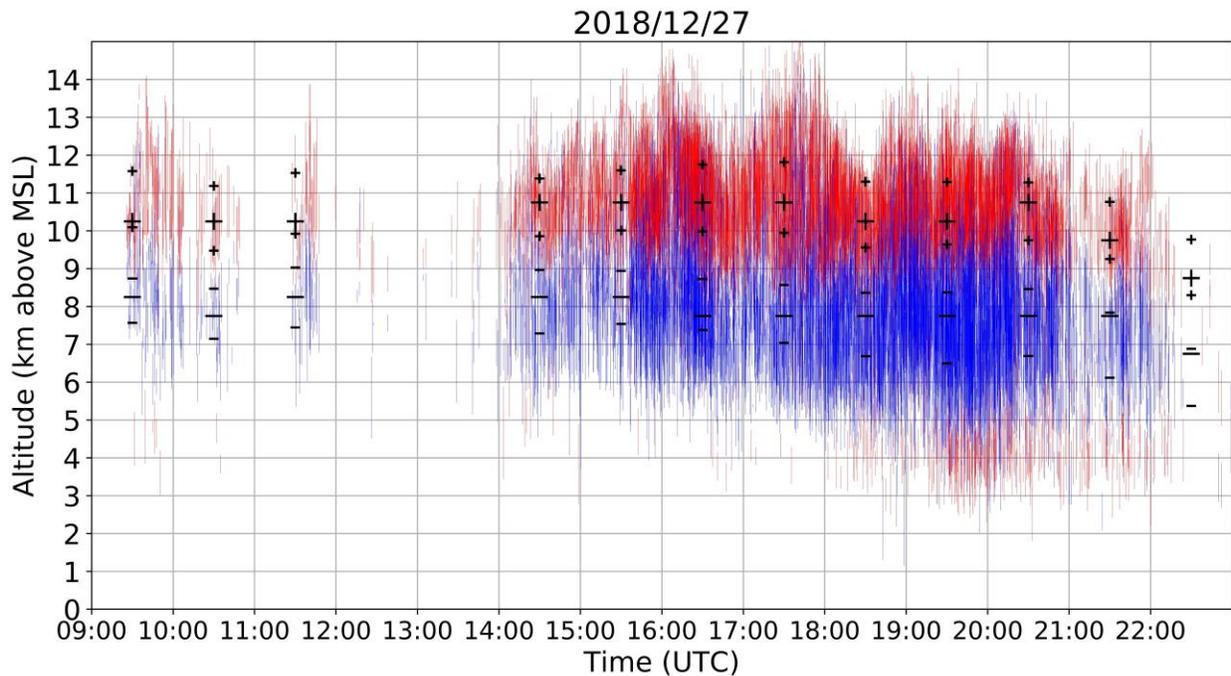
317 **4 Chargepol Method Applied to RELAMPAGO Thunderstorms**

318 As thunderstorm charge structure in Argentina is generally unknown, we illustrate in this
319 section events in which the Chargepol method described in the last section was applied. During
320 the five-month period the RELAMPAGO LMA network was operating in Cordoba, different
321 thunderstorm modes were observed, such as isolated convection, multicellular storms, supercells,
322 and mesoscale convective systems (Nesbitt et al. 2021). In order to demonstrate the capability of
323 the algorithm to depict gross charge structures, examples of distinct Cordoba cases and their
324 evolution in time are presented. Examples of thunderstorms with differing charge structure in
325 Cordoba are shown, such as normal, anomalous, a case with an enhanced lower positive, and a
326 case that presented a transition from one archetype to another. The altitude with most
327 occurrences of a charge layer polarity for every hour, and the mean altitude of the top and bottom
328 of each layer polarity for every hour is also presented in this section to demonstrate variation of
329 charge layer altitude in time for these cases. For a lower polarity of a dipole, the mean altitude of
330 the top and bottom of charge layers estimated from flashes were only calculated for charge layers
331 in which the top was at a lower altitude than the upper dipole polarity altitude. Similarly, for the
332 upper polarity dipole, mean altitude was obtained from charge layers with its bottom above the
333 altitude of the lower polarity dipole. These restrictions were put in place to focus analysis on the
334 top and bottom altitudes of the dominant positive and negative charge layers in the main dipole.
335 To further demonstrate the algorithm's capabilities over regions of the United States that have
336 been studied and well characterized with other charge retrieval methods (e.g., Bruning et al.,
337 2010; Chmielewski et al., 2016; MacGorman et al., 2005; Mecikalski et al., 2015; Wiens et al.,
338 2005), an example from each of the LMA networks deployed during DC3 are shown in the
339 supporting information, including a normal tripole case in Alabama, anomalous storms in
340 Colorado, a case with a transition from anomalous to normal charge structure in Oklahoma, and
341 a normal dipole in West Texas at typical altitudes (negative in mid-levels, positive in the upper
342 levels) but with a very high altitude negative charge layer observed above the upper positive. In
343 this section, the RELAMPAGO dataset is emphasized, as the charge structures in Central
344 Argentinian thunderstorms have yet to be documented in detail.

345 4.1 27 December 2018 Case: Normal Charge Structure

346 Figure 4 shows the estimate of charge layer polarity for all convective storms that
347 occurred in the RELAMPAGO LMA domain for a 14-hour period on 27 December 2018. Most
348 thunderstorms that occurred on this day presented an upper-level positive charge layer above 9-
349 10 km height, and a mid-level negative charge layer between about 5 and 9 km height. Altitude
350 variation in charge layers is speculated to be due to different thunderstorms having varying

351 updraft strength and cloud-top heights. As the number of estimated charge layers is correlated to
 352 flash rate (not shown), high flash rates are inferred from periods with charge layers compacted in
 353 short periods of time. For most of the period between 1500 and 2100 UTC, the total storm flash
 354 rate was higher than 50 flashes/minute, considering flashes with more than 10 sources and all
 355 active thunderstorms in the domain. The total flash rate of storms in the domain peaked at 195
 356 flashes/minute at 1609 UTC. The dominance of positive over negative charge structure means
 357 that most flashes depicted by Chargepol were +ICs, with a typical initial upward motion of a
 358 negative leader and further propagation through the upper positive charge layer. This general
 359 dipole structure characterizes a typical normal charge structure, as it is common in many regions
 360 of the United States as shown in similar LMA-based charge retrieval studies, such as in Alabama
 361 (Mecikalski et al., 2015) and Oklahoma (Bruning et al., 2007). Some flashes propagated through
 362 a lower positive charge layer below 5 km height, principally after 1920 UTC. That was caused
 363 by -IC flashes with initial negative leaders with downward motion and further propagation
 364 through the low-level positive charge region. Hence, from 1900 to 2200 UTC, a typical tripole
 365 charge structure (Williams, 1989) was present, though the upper positive region is considerably
 366 more active than the lower positive due to more flashes contributing to the upper positive
 367 depiction.



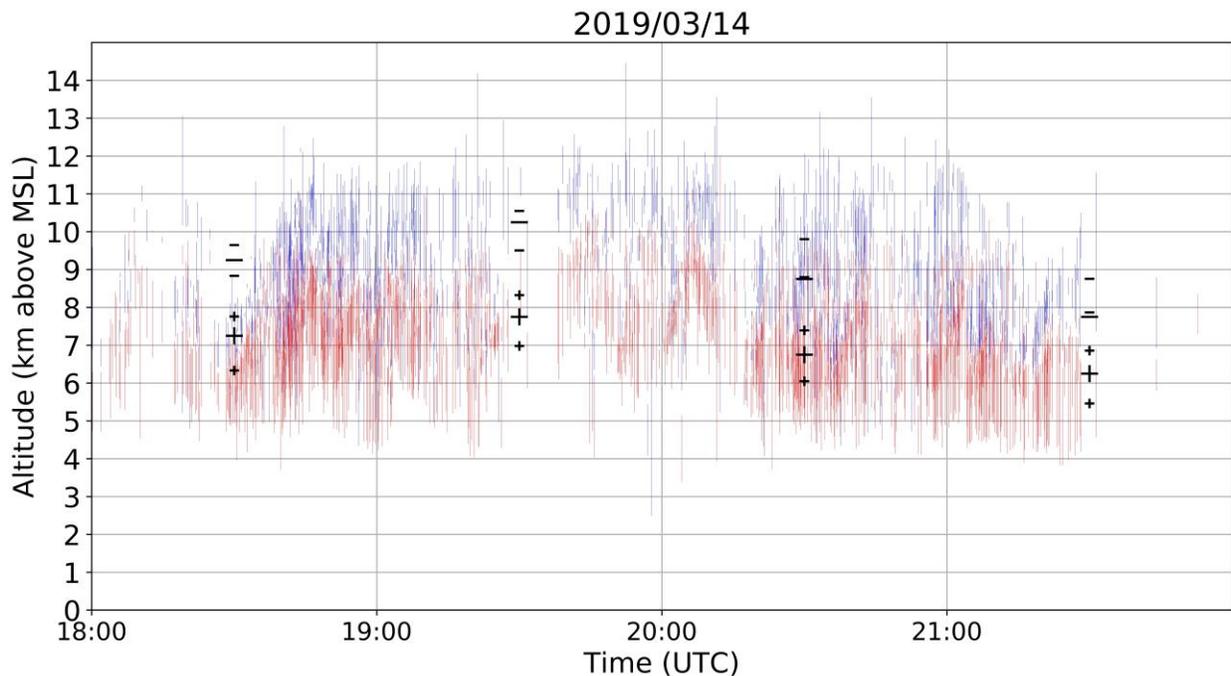
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369 **Figure 4.** Charge layers estimated from flashes using the Chargepol automated method for all
 370 RELAMPAGO thunderstorms on 27 December 2018 from 0900 to 2300 UTC. Each red (blue)
 371 vertical line represents a positive (negative) charge layer estimated from a flash. Large black
 372 symbols represent the altitudes in which most charge layers of a certain polarity were estimated
 373 for each hour period, as long as more than 30 layers with that polarity were present in that hour.

374 Small black symbols represent the mean altitudes of the top and bottom of charge layers for each
 375 polarity and hour.

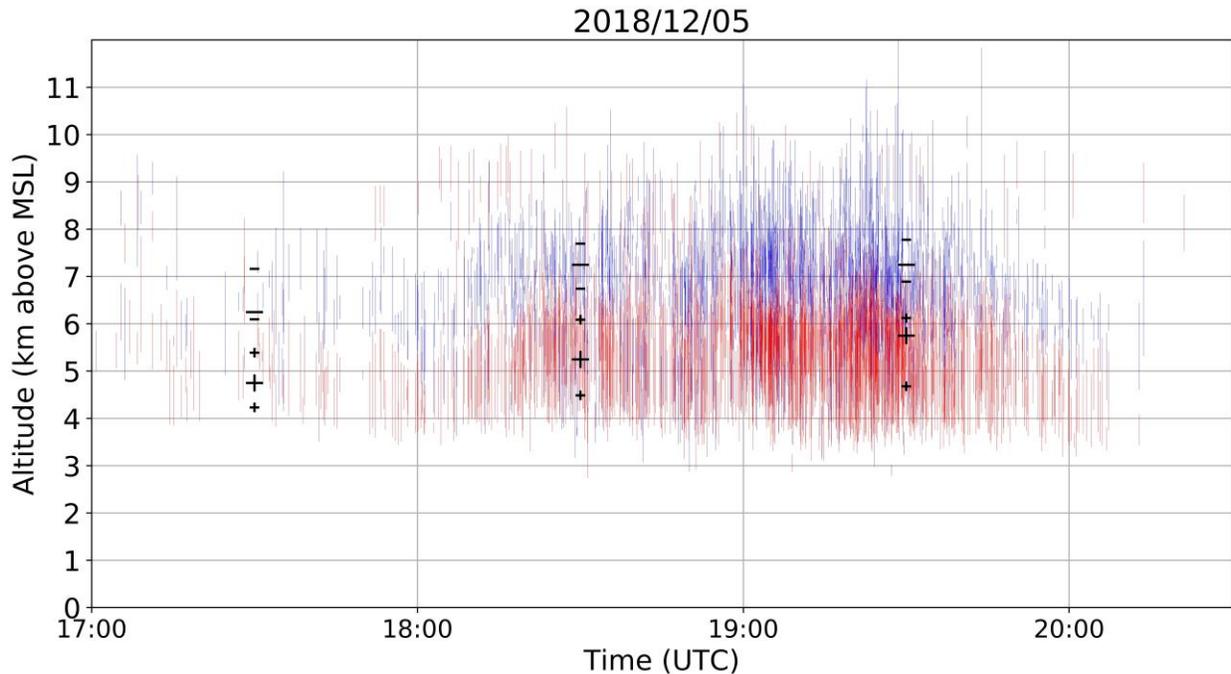
376 4.2 14 March 2019 and 5 December 2018 Cases: Anomalous Charge Structure

377 A cluster of RELAMPAGO storms on 14 March 2019, all with an anomalous dipole
 378 charge structure (e.g., Bruning et al., 2014; Marshall et al., 1995), is shown in Figure 5. These
 379 storms had a dominant mid-level positive charge layer and upper-level negative charge layer,
 380 similar to some anomalous storms over Colorado (Fuchs et al. 2015). As multiple storms are
 381 shown in Figure 5, a large altitude variation is noticeable for the charge layers, which is possibly
 382 dependent on individual storm intensity. Storms with stronger updrafts are speculated to have
 383 higher flashes and higher charge layers (Stolzenburg et al., 1998). Most flashes in these storms
 384 presented -IC lightning, which means that negative breakdown had an initial downward
 385 propagation, hence negative charge is estimated at higher levels than positive charge. A similar
 386 anomalous dipole case occurred in an isolated thunderstorm on 5 December 2018 (Figure 6).
 387 This storm had a flash rate higher than 30 flashes/minute for most of the period between 1815
 388 and 1945 UTC, with a peak flash rate of 128 flashes/minute at 1902 UTC. This anomalous case
 389 is different from the 14 March 2019 anomalous case because estimated charge layers are located
 390 at lower levels: negative charge is located in the mid-levels, while positive charge is in the low-
 391 levels. Also, this was a relatively shallow storm system exhibiting a radar echo top at about 10-
 392 km height (not shown), hence no upper positive charge layer had developed. Upper positive
 393 charge at about 9 km height from 1800 to 1900 UTC was from another storm in the domain (see
 394 discussion in section 3.1).



395

396 **Figure 5.** Same as in figure 4, but for 14 March 2019 from 1800 to 2200 UTC.

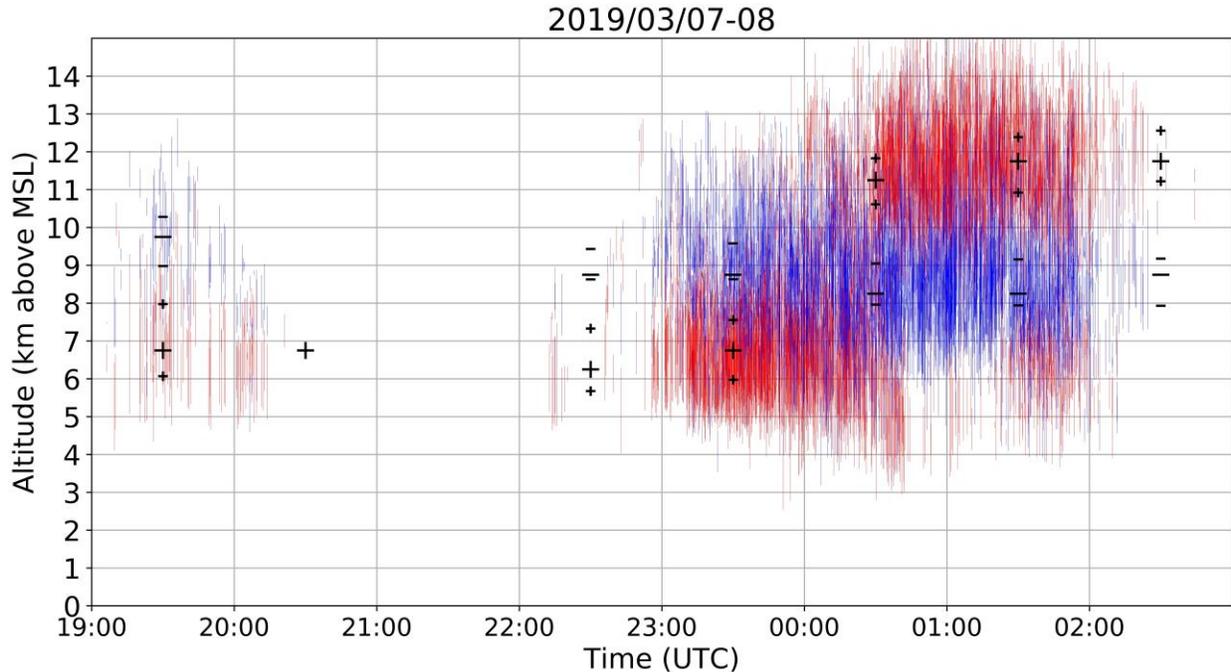


397

398 **Figure 6.** Same as in Figure 4, but for 5 December 2018 from 1700 to 2030 UTC.

399 4.3 7-8 March 2019 Case: Transition from Anomalous to Normal Charge Structure

400 Thunderstorms on 7 March 2019 (Figure 7) during RELAMPAGO presented an
 401 anomalous charge structure with mid-level positive charge and upper level negative charge.
 402 From 1900 to 2300 UTC, storms that occurred in the LMA domain had a low flash rate (less than
 403 30 flashes/minute considering all thunderstorms in the domain), then few charge layers were
 404 depicted by Chargepol, but an anomalous dipole is clearly present, similar to the storm studied
 405 by Fuchs et al. (2018) over Colorado. After 2300 UTC, a MCS formed with a dominant
 406 anomalous charge structure, with its flash rate rapidly increasing to more than 100 flashes/minute
 407 in the LMA domain. On the following UTC day, high flash rates remained, reaching a peak of
 408 496 flashes/minute at 0124 UTC in the domain, and an upper positive charge layer formed above
 409 10 km height. This upper positive layer became visible because flashes started propagating
 410 through that layer. After 0045 UTC, fewer flashes propagated through the lower positive charge
 411 layer. Hence, this case characterizes a transition from anomalous to normal charge structure. This
 412 case demonstrates how complex charge structure evolution can be estimated by the Chargepol
 413 method, such as the presence of anomalous and normal main dipoles, tripoles, and their
 414 evolution in time.



415

416 **Figure 7.** Same as in Figure 4, but for 7 March 2019 at 1900 UTC to 8 March 0300 UTC.

417 **5 Frequency of Anomalous Charge Structure in Central Argentina Compared to the U.S.**

418 As the described Chargepol method allows for a relatively fast processing time for large
 419 datasets of months of LMA data, one can obtain the general charge structure evolution in time
 420 for a domain area, as shown in the previous section. Hence, in order to characterize the
 421 likelihood of normal and anomalous charge structure for the five months in which the LMA was
 422 deployed in the Cordoba, Argentina region for the first time, the Chargepol layer polarity output
 423 was summarized for a better understanding and interpretation of the general dominant charge
 424 structure.

425 In order to achieve a summary of the main charge structure for such long periods, the
 426 charge layer polarity information was initially subdivided into time periods of one hour to obtain
 427 the dominant dipole for every hour period. Then, the number of charge layers of a given polarity
 428 were counted for every altitude in 0.5-km bins for every hour. The altitude with most positive
 429 charge layers estimated from flashes, and the altitude with most negative charge layers, define a
 430 single altitude bin for each layer polarity, which characterizes the dominant dipole for that hour,
 431 as long as both maximum polarities occur at different heights. A minimum threshold of 30
 432 charge layers from each polarity occurring in one hour was applied, in order to avoid
 433 thunderstorms with low flash rate contributing to the charge structure estimation. The large black
 434 symbols present in Figures 4-7 represent the altitude with most occurrences of a charge layer
 435 polarity for each hour, and the resultant main dominant dipole for an hour period.

436 In this study, an estimated dipole structure for a one hour period is referred to as a
 437 “sample”. Samples in which dipoles had positive located at a higher altitude than negative are
 438 referred in this study as normal charge structure (Dye et al., 1986; Williams, 1985). A normal
 439 charge structure sample could have few flashes that estimated the presence of a low-level
 440 positive charge layer, however if more flashes contributed to the maximum height occurrence of

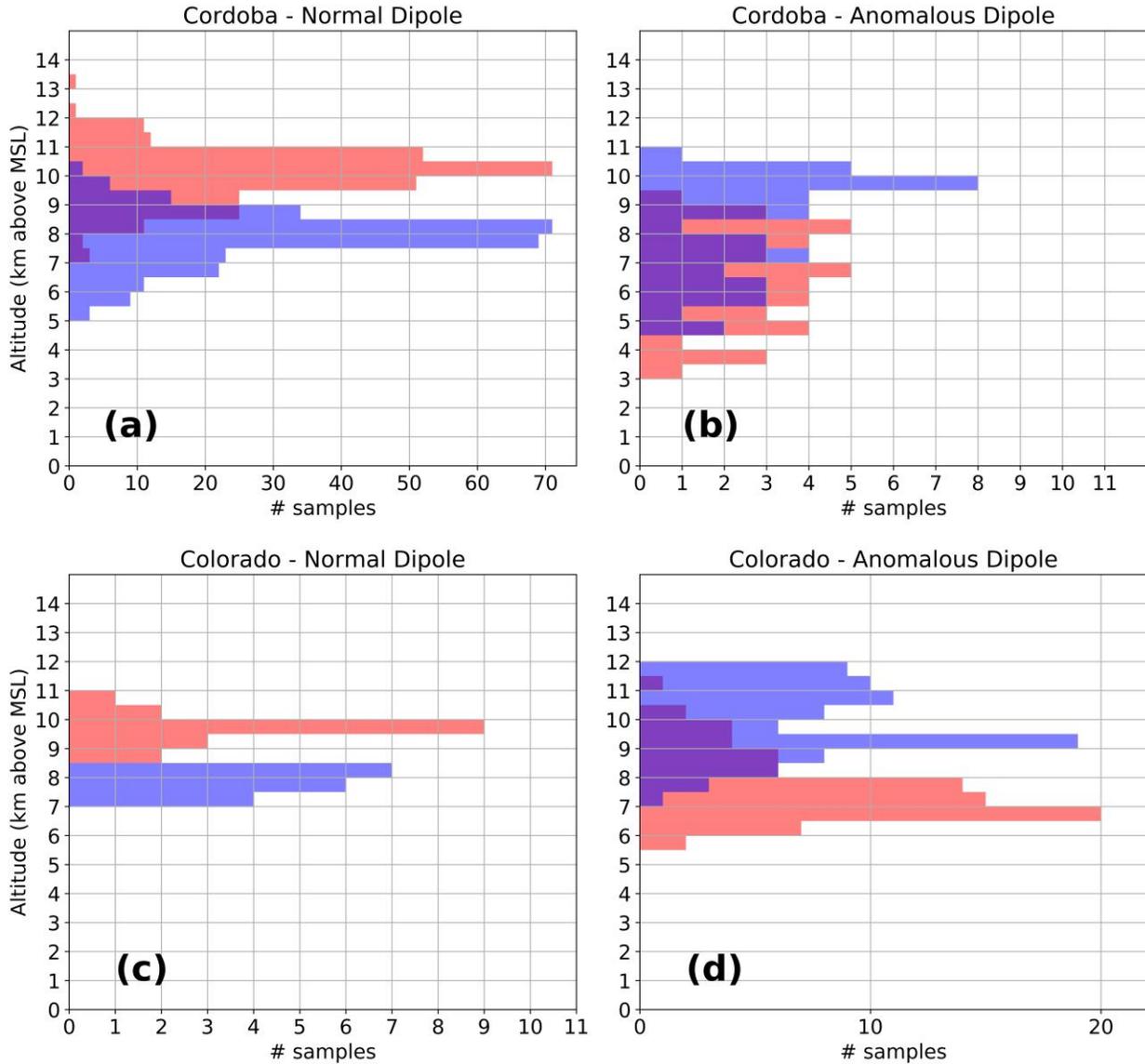
441 positive being at high levels, it would be considered a normal charge structure sample. Figure 4
442 shows an example of a normal tripole charge structure (Williams, 1989) with more positive
443 layers estimated at high levels, leading to a normal dipole estimation. Samples with negative
444 charge over positive charge are considered to have a dominant anomalous charge structure. The
445 most common type of anomalous dipole sample is the type with positive charge at the mid-levels
446 or mixed-phase layer, and negative in the upper levels of a storm (Figure 5). Another structure
447 that could lead to a negative-over-positive dipole is when enhanced positive charge is at low
448 levels of a storm, while negative charge is at the mid-levels (Figure 6, Bruning et al., 2014;
449 Fuchs et al., 2015). In this scenario, an upper positive charge could be present, which could lead
450 to an interpretation of a normal tripole charge structure, an uncommon characteristic during
451 RELAMPAGO as the enhanced low-level positive charge layer is not typically accompanied by
452 an upper-level positive charge layer (Figure 6). However, in this study and others (e.g., Fuchs et
453 al. 2015), a normal tripole scenario with more flashes propagating through the lower positive
454 charge layer than through the upper positive would imply the characterization of an anomalous
455 dipole. Both anomalous scenarios (positive in the mid-levels, and in the low levels) imply that
456 most flashes consisted of -ICs with negative leaders having an initial downward motion, rather
457 than +ICs that would initially move upward. Hence, in this study, scenarios with a dominance of
458 -ICs, or negative-over-positive dipoles, are considered anomalous.

459 During the five months that the LMA was operating in Cordoba, Argentina, 306 samples
460 were observed, which means 306 hours with lightning activity in which the aforementioned
461 methodology estimated a dominant dipole structure. Among the 306 Cordoba samples, 265
462 consisted of normal dipole charge structure, while the other 41 were anomalous (Table 1). That
463 means that 13.3% of samples had a dominant anomalous charge structure, which can be
464 interpreted as an approximate frequency of occurrence of anomalous storms in Cordoba,
465 Argentina. Table 1 shows the number of normal and anomalous samples for the Cordoba LMA
466 deployed during RELAMPAGO, as well as for the four LMA networks deployed during DC3 in
467 several locations across the United States (e.g., Colorado, West Texas, Oklahoma and Alabama)
468 for comparison, all sampled in the warm season. The comparison of Cordoba dipole samples
469 with DC3 networks is shown to demonstrate the application and capabilities of Chargepol, as it
470 resembles climatologies from other studies. Even though the sample numbers vary for the
471 different locations, consistent with Carey et al. (2003) and Fuchs and Rutledge (2018), Alabama
472 showed the lowest percentage of anomalous storms (7.3%), and Colorado anomalous frequency
473 was much higher than any other region (82.6%). Oklahoma and West Texas fell in between these
474 two regions, and with similar anomalous frequencies to Cordoba (12.9% for Oklahoma and
475 11.1% for West Texas). From the flash centroid altitude distribution for the entire
476 RELAMPAGO LMA dataset, Lang et al. (2020), observed a peak at 10 km height and a
477 secondary peak at 6 km height, the latter attributed to anomalous storms and stratiform lightning.
478 For the normal and anomalous Cordoba events shown in Lang et al. (2020), normal and
479 anomalous samples were consistently depicted by the Chargepol algorithm. Table 1 also shows
480 the total number of flashes with more than 20 sources, and the fraction of flashes that were
481 considered by the algorithm, being 16.7% overall for all LMA networks.

482 **Table 1.** Number of normal and anomalous samples for Cordoba, Alabama, West Texas,
 483 Oklahoma, and Colorado.

	Cordoba	Alabama	West Texas	Oklahoma	Colorado
Number of days	157	32	48	41	61
Number of flashes (>20 sources)	808416	39046	261713	497139	545005
Number of flashes considered by Chargepol	165767 (20.5%)	7653 (19.5%)	65309 (24.9%)	58900 (11.8%)	62556 (11.4%)
Total number of samples	306	41	99	80	98
Normal samples	265	38	88	70	17
Anomalous samples	41	3	11	10	81
% Anomalous	13.3	7.3	11.1	12.5	82.6

484 The distribution of normal samples with altitude demonstrated that most normal dipoles
 485 were present in the mid-to-upper levels; i.e., with mid-level negative and upper-level positive
 486 charge. Figure 8a shows the distribution of normal dipoles for Cordoba, Argentina. The altitude
 487 distribution of anomalous samples (i.e., negative over positive dipoles) in Argentina shows that
 488 there were cases in which negative charge was present in the upper levels with positive in the
 489 mid-levels, and cases of negative in the mid-levels, with enhanced positive in the low levels
 490 (Figure 8b). In Colorado, few normal samples were observed, but their altitude distribution is
 491 similar to Cordoba (Figure 8c). The distribution of anomalous samples with altitude in Colorado
 492 showed that most dipoles had upper level negative and mid-level positive. No apparent presence
 493 of an anomalous dipole located in the low-mid-levels occurred in the Colorado DC3 LMA
 494 dataset. Therefore, the Cordoba 5 December 2018 case (Figure 6) demonstrates a singular
 495 thunderstorm charge structure that is either rare or completely absent in Colorado. The normal
 496 sample distributions in height for the other 3 U.S. locations (not shown) were similar to Cordoba
 497 and Colorado, while the anomalous sample distribution for these 3 locations (not shown) proved
 498 inconclusive due to the low sample number.



499

500 **Figure 8.** Distribution of normal and anomalous samples with altitude for Cordoba (a, b) and
 501 Colorado (c, d).

502 **6 Summary and Discussion**

503 This paper presented charge structures for the warm season thunderstorms in Cordoba,
 504 Argentina for the first time through thunderstorm examples and long-term statistics utilizing a
 505 new method that identifies charge layer polarity at a flash level from LMA VHF data. This
 506 method is able to estimate general charge structures such as normal and anomalous dipoles,
 507 tripoles, altitude and depth of charge layers. Chargepol was applied to months of LMA data,
 508 allowing for the inference of the frequency of anomalous and normal charge structure
 509 thunderstorms in Cordoba, Argentina, and comparison to four well-studied U.S. regions using
 510 the same methodology.

511 This method was developed from a meteorological standpoint, which means that the
512 objective was to obtain the general charge structure evolution through the entire thunderstorm
513 life cycle, or for many hours of data. In order to achieve that, there was no need to retrieve
514 charge polarity from every flash as demonstrated in the comparison of Chargepol relative to
515 manual charge structure analysis and the VHF source distribution peak. Instead, only flashes
516 with less doubtful characteristics were used to provide an accurate charge polarity retrieval.
517 Hence, when considering such long periods of time, the frequency of anomalous and normal
518 charge structures can be estimated. Also, we found that it is sufficient to summarize the data into
519 the main dominant dipoles for every hour in order to characterize the charge structure for a
520 region. It is important to emphasize that, once charge layers are retrieved from individual
521 lightning flashes, one can organize this same dataset in any other manner depending on the user's
522 purpose. Examples include considering the algorithm output as a database to be organized into
523 shorter or longer time periods, obtaining the density of charge layers polarity over the time-
524 altitude domain, calculating statistics for comparison with observations from other
525 instrumentation such as radar, etc.

526 The complexities of a three-dimensional charge structure that may be present at sub-
527 storm scale, with charge layers extending through different altitudes depending on distance to an
528 updraft core, are not being fully accounted for. For a flash analysis, we consider the charge
529 distribution over the vertical dimension only, which proved to be sufficient for this study's
530 objectives. For a given flash, the Chargepol method can estimate no more than two charge layers
531 with opposite polarities. However, when observing charge layers output for numerous flashes, it
532 is possible to infer the presence of dipoles, their altitude and time evolution, the presence of
533 tripoles and even multiple charge layers if flashes propagate through it. Only charge layers that
534 had flashes moving through them can be inferred. In the case of a positive charge layer without a
535 lightning flash moving through it, the charge layer cannot be visualized as a product of the
536 algorithm, which is a fundamental limitation of all LMA-based charge retrieval methods (Rust et
537 al., 2005). The fact that Chargepol neglects small flashes for charge layer estimation, as it
538 discards flashes with less than 20 sources, makes it hard to locate small pockets of charge within
539 thunderstorms. Even if these charge regions were located, it could be hard to visualize and
540 interpret their evolution over minutes. However, estimating the general dipole and tripole charge
541 structures is feasible with this algorithm, satisfying this study's purpose.

542 The Chargepol method proved capable for analyzing large LMA datasets in a reasonable
543 processing time of minutes, allowing for efficient interpretation of charge structures over
544 Cordoba, Argentina during the recent RELAMPAGO field campaign and a consistent
545 comparison of these novel results with thunderstorms from different regions of the United States
546 whose charge structures have been sampled with LMA and are more well understood. A high
547 frequency of anomalous storms were found for Colorado consistent other studies (Fuchs et al.,
548 2018). Examples of Cordoba anomalous thunderstorms with altitude distributions of positive
549 charge layers that are uncommon in Colorado were presented. Interestingly, Cordoba showed
550 slightly higher anomalous charge structure frequency compared to Oklahoma and West Texas,
551 while Alabama presented the lowest anomalous frequency among all studied regions consistent
552 with prior work (Fuchs et al., 2018). Reasonings for these results were not explored in this study.
553 The meteorological, environmental, kinematic and microphysical conditions in Central
554 Argentina are speculated to be important contributors to the observed charge structures
555 documented herein during RELAMPAGO, and they will be explored in future studies and
556 compared to past work from other regions throughout the world. The charge polarity outputs

557 presented in this study have the potential to be useful for numerous applications in lightning
558 research, and Chargepol will be made available as an open-source algorithm in the near future.

559 **Acknowledgments and Data**

560 We thank the NOAA GOES-R Program and the NASA Lightning Imaging Sensor (LIS) for the
561 RELAMPAGO LMA funding. The first author and coauthors Carey and Bitzer wish to
562 acknowledge support from the National Science Foundation Award AGS 1661785 and NASA
563 MSFC Grant NNM11AA01A. We thank RELAMPAGO collaborators: Eldo E. Ávila, Rodolfo
564 G. Pereyra, Richard J. Blakeslee, Jeff Burchfield, Matthew Wingo, Steven J. Goodman, Michael
565 Solomon, and Joy Marich. RELAMPAGO LMA data are available on
566 <https://doi.org/10.5067/RELAMPAGO/LMA/DATA101>. NSF DC3 LMA data are available on
567 https://data.eol.ucar.edu/master_lists/generated/dc3/.

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