

1 **Lithosphere structure and seismic anisotropy offshore eastern**
2 **North America: Implications for continental breakup and**
3 **ultra-slow spreading dynamics**

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8 Key Points:

- 9 • A 15–20 km thick low-velocity lid extends ~200 km from the margin and is interpreted as stretched
10 continental mantle lithosphere
- 11 • Complete continental breakup and onset of normal seafloor spreading occurred ~170 Ma, ~25 Ma later
12 than previously thought
- 13 • Observed margin-parallel lithospheric anisotropy resulted from plate-motion induced shear during
14 ultra-slow spreading 170–200 Ma

Abstract

16 The breakup of supercontinent Pangea occurred ~ 200 Ma forming the Eastern North American Margin
17 (ENAM). Yet, the precise timing and mechanics of breakup and onset of seafloor spreading remain poorly con-
18 strained. We investigate the relic lithosphere offshore eastern North America using ambient-noise Rayleigh-
19 wave phase velocity (12–32 s) and azimuthal anisotropy (17–32 s) at the ENAM Community Seismic Experi-
20 ment (CSE). Incorporating previous constraints on crustal structure, we construct a shear velocity model for
21 the crust and upper ~ 60 km of the mantle beneath the ENAM-CSE. A low-velocity lid (V_S of 4.4–4.55 km/s)
22 is revealed in the upper 15–20 km of the mantle that extends ~ 200 km from the margin, terminating at the
23 Blake Spur Magnetic Anomaly (BSMA). East of the BSMA, velocities are fast (>4.6 km/s) and character-
24 istic of typical oceanic mantle lithosphere. We interpret the low-velocity lid as stretched continental mantle
25 lithosphere embedded with up to $\sim 15\%$ retained gabbro. This implies that the BSMA marks successful
26 breakup and onset of seafloor spreading ~ 170 Ma, consistent with ENAM-CSE active-source studies that
27 argue for breakup ~ 25 Myr later than previously thought. We observe margin-parallel Rayleigh-wave az-
28 imuthal anisotropy (2–4% peak-to-peak) in the lithosphere that approximately correlates with absolute plate
29 motion (APM) at the time of spreading. We hypothesize that lithosphere formed during ultra-slow seafloor
30 spreading records APM-modified olivine fabric rather than spreading-parallel fabric typical of higher spread-
31 ing rates. This work highlights the importance of present-day passive margins for improving understanding
32 of the fundamental rift-to-drift transition.

Plain Language Summary

33 The Eastern North American Margin (ENAM) formed during the breakup of supercontinent
34 Pangea, marking the opening of the Atlantic Ocean. However, details of the timing and mechan-
35 ics of the breakup are not well understood. The ENAM region provides a fossilized record of this
36 transition from continental rifting to seafloor spreading, informing understanding of the funda-
37 mental “rift-to-drift” plate-tectonic process. Using Rayleigh-waves, we solve the the 3-D shear
38 velocity structure of the lithosphere offshore North Carolina, revealing a 15–20 km thick low-
39 velocity “lid” that extends ~ 200 km from the margin, terminating at the Blake Spur Magnetic
40 Anomaly. We interpret this as stretched continental lithosphere, which implies that complete
41 breakup of Pangea did not occur directly at the margin but rather ~ 200 km seaward. This cor-
42 responds to a breakup age of ~ 170 Ma, ~ 25 Myr later than previously thought. We also observe
43 a directional dependence of Rayleigh-wave velocities, where waves traveling parallel to the mar-
44 gin propagate 2–4% faster than waves traveling perpendicular, opposite of what is expected for
45 oceanic lithosphere. This provides evidence for margin-parallel deformation in the mantle during
46 breakup ~ 170 – 200 Ma. We propose that relative motion of the overriding plate drives mantle
47 deformation in ultra-slow seafloor spreading environments.

48 1 Introduction

49 The Eastern North American Margin (ENAM) is a passive volcanic margin that formed during the breakup
50 of supercontinent Pangea ~ 200 Ma (*Withjack et al.*, 1998). The breakup initiated at approximately 235 Ma
51 with crustal extension that is recorded in a sequence of rift basins along the length of the margin (*Withjack*
52 *et al.*, 2012), and it occurred alongside an outburst of volcanism known as the Central Atlantic Magmatic
53 Province (CAMP) that is dated to 200 ± 4 Ma (*Marzoli et al.*, 2011) and characterized by a large volume
54 ($\sim 1 \times 10^6$ km³) of flood basalts emplaced during a short period (< 1 Myr) (*McHone*, 2003, *Olsen et al.*,
55 2003). Though the precise temporal and tectonic relationship between CAMP magmatism and rifting is
56 debated (*McHone*, 2000), it is thought that normal seafloor spreading and opening of the Atlantic basin
57 began sometime between ~ 200 – 170 Ma.

58 Insights into the transition from continental rifting to seafloor spreading are contained in the crust and mantle
59 signature offshore ENAM. The offshore region is characterized by two positive-polarity magnetic anomalies
60 separated by the Inner Magnetic Quiet Zone (IMQZ) that lacks well defined magnetic lineations (Figure 1).

61 The higher amplitude East Coast Magnetic Anomaly (ECMA) occurs just seaward of the continental shelf
62 and has generally been interpreted as marking the transition to oceanic crust (e.g. *Kelemen and Holbrook,*
63 *1995, Klitgord et al., 1988, Lynner and Porritt, 2017*). ECMA emplacement ages range from 175–200 Ma
64 (*Benson, 2003, Klitgord and Schouten, 1986, Labails et al., 2010*), though recent revised estimates of ~ 195 Ma
65 have been proposed based on the African conjugate to the ECMA as well as salt deposits off Nova Scotia
66 and Morocco (*Labails et al., 2010, Sahabi et al., 2004*). Approximately 200 km seaward of the ECMA is the
67 lower amplitude Blake Spur Magnetic Anomaly (BSMA). The age of the BSMA is estimated at ~ 170 Ma,
68 but its precise origin and significance are still debated (*Greene et al., 2017*).

69 Recent work offshore ENAM has challenged the notion that the ECMA marks the completion of continental
70 breakup and onset of seafloor spreading, and instead, it has been proposed that the BSMA marks this impor-
71 tant transition (*Bécel et al., 2020, Shuck et al., 2019*). Using data collected during The ENAM Community
72 Seismic Experiment (ENAM-CSE) that extends farther offshore than ever before, detailed crustal imaging
73 shows thin proto-oceanic crust with higher lower-crust velocities and rougher basement topography west of
74 the BSMA compared to the east, consistent with deeper melting and ultra-slow opening rates (~ 0.65 cm/yr
75 half-rate) (*Bécel et al., 2020, Shuck et al., 2019*). This interpretation predicts a 15–20 km thick continental
76 mantle lithosphere underlying the oceanic crust west of the BSMA and normal oceanic lithosphere east of
77 the BSMA; however, previous shear-velocity imaging in the region shows little evidence for a distinct change
78 in mantle velocities across the BSMA (*Lynner and Porritt, 2017*).

79 Additional insights into dynamics associated with continental rifting and seafloor spreading may be preserved
80 in the olivine lattice-preferred orientation (LPO) in the lithospheric mantle, which acts as a record of past
81 mantle flow. The fast [100] axis of olivine crystals tend to align parallel to the direction of shearing in the
82 mantle, forming an LPO (*Karato et al., 2008, Zhang and Karato, 1995*). In models of mid-ocean ridges
83 (MOR), corner flow near the ridge aligns olivine LPO parallel to the spreading direction and is frozen-in as
84 the lithosphere cools (e.g. *Blackman and Kendall, 2002, Blackman et al., 1996, Kaminski and Ribe, 2001,*
85 *2002, Ribe, 1989*). This frozen-in LPO leads to the azimuthal anisotropy of seismic waves routinely observed
86 in the Pacific lithosphere, with a fast azimuth parallel to the fossil-spreading direction (FSD) (e.g. *Forsyth,*
87 *1975, Hess, 1964, Mark et al., 2019, Morris et al., 1969, Raitt et al., 1969, Russell et al., 2019*). Seismic
88 anisotropy of the deeper asthenosphere reflects present-day mantle deformation and at large length scales
89 broadly aligns sub-parallel to absolute plate motion (APM) beneath the ocean basins (*Beghein et al., 2014,*
90 *Nishimura and Forsyth, 1989, Schaeffer et al., 2016*), with deviations associated with smaller-scale convective
91 processes (e.g. *Becker et al., 2014, Lin et al., 2016*).

92 Previous seismic anisotropy observations at the ENAM-CSE from shear-wave splitting show margin-parallel
93 fast axes, significantly rotated from FSD and current-day APM, that was interpreted as present-day astheno-
94 spheric flow along the margin (*Lynner and Bodmer, 2017*). In addition, preliminary comparison of sub-Moho
95 V_P along crossing margin-parallel and margin-perpendicular refraction lines in the region suggests a margin-
96 parallel fast direction in the shallow lithosphere that is approximately perpendicular to the FSD (*Shuck and*
97 *Van Avendonk, 2016*). However, these observations are limited to two locations in the ENAM-CSE footprint
98 where the refraction lines intersect.

99 In this study, we use ambient-noise Rayleigh waves to construct a shear velocity model of the offshore ENAM-
100 CSE region that incorporates recent crustal constraints from refraction tomography (*Shuck et al., 2019*). Our
101 model reveals relatively low-velocity mantle lithosphere extending ~ 200 km seaward that we interpret in
102 the context of the detailed crustal architecture, providing further evidence for a prolonged breakup prior to
103 seafloor spreading. We also report margin-parallel Rayleigh-wave anisotropy in the lithosphere, perpendicular
104 to typical expectations for seafloor spreading, and offer an alternative perspective on lithosphere fabric formed
105 at slow-spreading ridges.

106 **2 Data and Methods**

107 The Eastern North American Margin Community Seismic Experiment (ENAM-CSE) consisted of onshore-
108 offshore active-source reflection and refraction as well as a one year broadband ocean-bottom seismometer
109 (OBS) deployment (*Lynner et al., 2020*). We use continuous seismic data from 28 broadband OBS deployed
110 during the ENAM-CSE from April 2014–May 2015. Water depth in the study region ranges from ~ 1300 m
111 near the shelf to ~ 5200 m on the eastern-most edge of the array. Instrument response is deconvolved to
112 displacement, and seismograms are downsampled to 1 Hz prior to processing.

113 **2.1 Daily OBS tilt noise removal**

114 We observe exceptionally strong horizontal noise on the vertical channels (i.e., tilt noise) at ENAM at periods
115 > 10 s (Figure 2a), presumably due to the strong Gulf Stream current that flows northeastward along the
116 coast. We remove this coherent horizontal energy from the vertical channels for each 24 hour segment of
117 the continuous data using the Automated Tilt and Compliance Removal (ATaCR) software (*Janiszewski*
118 *et al., 2019*), which implements the techniques developed by *Crawford and Webb (2000)*. We do not remove

119 pressure coherence from the vertical channel, as this has been shown to degrade the desired fundamental-
120 mode primary microseism (Bowden *et al.*, 2016).

121 It is not common practice to remove daily tilt noise prior to performing ambient-noise tomography at
122 OBS experiments, but we find that tilt removal improves the overall signal-to-noise ratio (SNR) of vertical
123 component empirical Green’s functions (EGFs) (See Section 2.2) by a factor of ~ 2 on average, and by an
124 order of magnitude for some station pairs (Figure 2b). The largest SNR improvements occur for station pairs
125 with shallower average water depth (Figure 2d). Similar ambient noise SNR improvements were reported
126 after tilt and compliance corrections at the shallow water stations at the Cascadia Initiative (Tian and
127 Ritzwoller, 2017).

128 2.2 Ambient noise processing

129 Ambient noise EGFs are constructed from tilt-removed seismograms following the general procedure of
130 Bensen *et al.* (2007) (Figure 3); however, we do not apply time-domain normalization or spectral whitening.
131 Daily displacement seismograms are split into 15 3-hour segments with 50% overlap. Normalized coherence
132 cross-spectra are calculated between the vertical channels for stations i and j and time window k :

$$\rho_{ijk}(\omega) = \frac{U_{ik}(\omega)U_{jk}^*(\omega)}{\sqrt{U_{ik}(\omega)U_{ik}^*(\omega)U_{jk}(\omega)U_{jk}^*(\omega)}} \quad (1)$$

133 where $U(\omega)$ is the vertical component displacement spectra at frequency ω and $U^*(\omega)$ is its complex conju-
134 gate. These coherence spectra are summed over the entire duration of the deployment for each station pair,
135 resulting in a final stacked spectrum, $\rho_{ij}(\omega) = \sum_k \rho_{ijk}(\omega)$. By utilizing the coherence spectrum rather than
136 the unnormalized cross-correlation spectrum and relatively short time windows, any windows containing
137 anomalous signals such as large earthquakes have little influence on the final stacked spectrum, precluding
138 the need for time-domain normalization or other alteration of the original high-quality waveforms. We found
139 that a typical one-bit normalization and spectral whitening procedure (e.g. Bensen *et al.*, 2007) degrades
140 signal-to-noise ratio in the 15–40 s period band by more than an order of magnitude.

141 **2.3 Interstation phase velocities**

142 Interstation phase velocities are estimated from the stacked coherence spectra using Aki’s spectral formula-
 143 tion, whereby the real part of the cross-spectra takes the form (Aki, 1957):

$$\bar{\rho}(\omega, r) = J_0\left(\frac{\omega r}{c(\omega)}\right) \quad (2)$$

144 where r is station separation, c is phase velocity, and J_0 is the Bessel function of order zero. Phase velocity
 145 dispersion is estimated at each zero crossing following *Ekström et al. (2009)* and then interpolated to a
 146 uniform frequency axis (Figure 4). In theory, this process identifies an infinite number of possible dispersion
 147 curves, and we select the one with 25–30 s velocities that are closest to a nominal mantle velocity of 3.9 km/s
 148 (Figure 4c). We then discard dispersion curves that are not smooth or do not decrease with increasing
 149 frequency. In order to minimize noise in the cross-spectrum prior to the zero-crossing analysis, we apply a
 150 cosine-taper window in the time domain defined by a minimum group velocity threshold of 0.3 km/s (that
 151 is, energy corresponding to group velocities < 0.3 km/s is zeroed). Additional smoothing is applied to $\bar{\rho}$ for
 152 periods > 17 s to eliminate the occurrence of spurious zero crossings, particularly for large r . Dispersion
 153 measurements for r less than one wavelength are discarded.

154 **2.4 Phase velocity inversion**

155 Interstation Rayleigh-wave phase velocities are inverted for 2-D phase velocity maps from 12–32 s period and
 156 azimuthal anisotropy from 17–32 s for the offshore ENAM region. We solve first for phase slowness maps
 157 $s(\mathbf{x})$ on a $0.1^\circ \times 0.1^\circ$ grid and take the reciprocal to obtain maps of phase velocity $c(\mathbf{x})$. Anisotropy terms
 158 are solved on a coarser $0.5^\circ \times 0.5^\circ$ grid. A perturbation in phase delay time $\delta\tau_{ij}$ between stations i and j
 159 due to a perturbation in phase slowness $\delta s(\mathbf{x})$ is given by

$$\delta\tau_{ij} = \iint_{\Omega} K_{ij}(\mathbf{x}, \omega) \delta s(\mathbf{x}) d\mathbf{x} - \int_i^j [A_c(\mathbf{x}, \omega) \cos(2\theta_{ij}) + A_s(\mathbf{x}, \omega) \sin(2\theta_{ij})] d\mathbf{r} \quad (3)$$

160 where θ_{ij} is propagation azimuth and r_{ij} is the great-circle distance between the stations. Coefficients A_c
 161 and A_s describe the frequency-dependent azimuthal anisotropy within the array footprint with magnitude
 162 $A = \sqrt{A_c^2 + A_s^2}$ and fast azimuth $\psi = 0.5 \tan^{-1}(A_s/A_c)$. The 2-D finite-frequency sensitivity kernel $K_{ij} =$

163 $\partial\tau/\partial s$ is given by (*Lin and Ritzwoller, 2010*)

$$K_{ij}(\mathbf{x}, \omega) = \frac{2c_0}{\omega} \sqrt{\frac{r_{ij}}{8\pi |\mathbf{x} - \mathbf{x}_j| |\mathbf{x} - \mathbf{x}_i|}} \cos\left(\omega \left[\tau_i(\mathbf{x}_j) - \tau_i(\mathbf{x}) - \tau_j^\dagger(\mathbf{x})\right] + \frac{\pi}{2}\right) \quad (4)$$

164 where $\tau_i(\mathbf{x})$ is the phase delay surface due to an impulse at station i , and $\tau_j^\dagger(\mathbf{x})$ is the adjoint phase delay
 165 field due to an impulse at station j . Reference velocity c_0 is taken as the average of all interstation dispersion
 166 measurements at frequency ω . To ensure that the kernel density matches the ray-theoretic value, the kernel
 167 is normalized such that

$$r_{ij} \iint_{\Omega} K_{ij}(\mathbf{x}, \omega) d\mathbf{x} = \frac{r_{ij}}{c_0} = \tau_0 \quad (5)$$

168 This formulation of the kernel, termed the “empirical” kernel by *Lin and Ritzwoller (2010)*, accounts for
 169 both finite-frequency effects and off-great-circle propagation caused by lateral velocity gradients along the
 170 ray path. In theory, the quantities $\tau_i(\mathbf{x})$ and $\tau_j^\dagger(\mathbf{x})$ can be determined empirically for each station by
 171 fitting a smooth surface to the interstation phase delays measured at all other stations across the array. In
 172 practice however, this is challenging due to uneven data distribution and presence of noise in the phase delay
 173 measurements.

174 We take a “semi-empirical” approach and approximate $\tau(\mathbf{x})$ numerically via spectral-element method (SEM)
 175 simulations by propagating S-waves through a realistic phase velocity map at each frequency of interest for
 176 an impulse centered at each station. The input synthetic phase velocity maps are constructed from the
 177 2-D V_P model along Line 1 from *Shuck et al. (2019)* after converting to V_S assuming $V_P/V_S = 1.8$ for the
 178 crust and mantle and 2.22 for the sediments (Figure 5) and extrapolating phase velocities to the full ENAM
 179 footprint via constant depth contours. Example finite-frequency kernels are shown in Figure 6 for 13 s and
 180 20 s period.

181 Equation (4) gives the sensitivity kernel at an instantaneous frequency ω and contains all Fresnel zones,
 182 but in practice each phase velocity measurement is made over a finite bandwidth $[\omega - \Delta\omega/2, \omega + \Delta\omega/2]$.
 183 We approximate finite-bandwidth kernels by forming a Gaussian weighted average of instantaneous kernels
 184 centered on ω with a half width of 10%, effectively limiting the spatial extent of the kernel to the first few
 185 Fresnel zones.

186 Phase slowness maps and anisotropy are inverted via eq. (3) using a linearized iterative least squares ap-

187 proach (Menke, 2012), and the final model is obtained after two iterations. Perturbations to the starting
188 homogeneous, isotropic phase slowness model are regularized using wavelength-weighted second derivative
189 smoothing and norm damping towards the starting model. At periods less than 17 s, anisotropy terms
190 are damped to zero, as variations in phases velocity associated with anisotropy are swamped by the large
191 variations associated with water depth. Sensitivity kernels are updated upon the second iteration to account
192 for off-great-circle propagation associated with lateral gradients in slowness.

193 2.4.1 Off-great-circle propagation

194 Water depth across ENAM increases by nearly 4000 m from west to east, leading to a drastic difference in
195 short period Rayleigh-wave sensitivity across the array that is strongly controlled by water depth (Figure 7).
196 A 13 s Rayleigh wave that primarily samples the eastern-most edge of the array where water is deep (4000–
197 5000 m) will travel slower than the same wave sampling shallow water near the shelf. This strong lateral
198 velocity gradient at short periods steers energy towards the faster shallow-water regions leading to off-
199 great-circle propagation (Figure 6). If great-circle paths are assumed, as is routinely done in surface-wave
200 tomography studies, an apparent 2θ azimuthal anisotropy signal is produced, where waves that travel parallel
201 to the margin appear faster on average than waves that travel perpendicular to the margin (Figure 6e). Left
202 unaccounted for, this off-great-circle propagation can result in (1) apparent 2θ azimuthal anisotropy with a
203 fast axis parallel to the margin and (2) biased fast isotropic velocities, on average. We avoid these biases
204 by using semi-empirical finite-frequency kernels that account for off-great-circle propagation and by limiting
205 anisotropy measurements to periods greater than 17 s, where the anisotropy bias is small (Figure 6e).

206 2.5 Inversion for shear velocity, V_S

207 The single-frequency phase velocity maps are resampled to a $0.25^\circ \times 0.25^\circ$ grid and combined to produce a
208 dispersion curve at each grid point in the model. Each dispersion curve is individually inverted for a 1-D
209 depth dependent shear velocity (V_S) profile, and all profiles are ultimately combined to produce the final
210 3-D V_S model. The starting 3-D V_S model for the inversion is the same as that used for the semi-empirical
211 kernels in Section 2.4, built up from the 2-D crustal V_P model along line 1 from Shuck *et al.* (2019) assuming
212 $V_P/V_S = 1.8$ in the crust and mantle and 2.22 in the sediments. Estimates of V_P/V_S are taken from the V_P
213 and V_S models along line 3 of Shuck *et al.* (2019). The mantle is not well resolved by the P_n modeling, so we
214 assume a nominal mantle V_P of 8.0 km/s. As seismic structure mostly varies perpendicular to the margin

215 (*Shuck et al.*, 2019) (i.e., with water depth), the 2-D line is extrapolated to the entire 3-D ENAM region
216 along contours of constant water depth.

217 Each 1-D inversion aims to solve for V_S while minimizing dispersion misfit via a standard linearized iterative
218 least-squares approach (*Menke*, 2012). The perturbation kernels and forward estimates of phase velocity
219 are calculated for a layered Earth model extending to 250 km using the SURF96 software (*Herrmann*,
220 2013) (Figure 7). The inversion is regularized with norm damping towards the starting model, second
221 derivative smoothing, and a constraint that seeks to preserve layer gradients in the crust and mantle (*Russell*
222 *et al.*, 2019). These constraints are broken across prescribed sediment and Moho boundaries to allow for
223 discontinuities. Compressional velocities (V_P) and density are held fixed. A linear increase in damping
224 toward the reference model is applied below 50 km such that the reference model is exactly maintained
225 by ~ 100 km depth. A total of 10 iterations are performed: iterations 1–9 maintain the same kernel, and
226 dispersion is predicted via perturbation theory using the kernels; upon iteration 10, the kernels and predicted
227 dispersion are recalculated using SURF96.

228 In order to evaluate starting model dependence, a Monte Carlo approach is used. At each grid point, the
229 starting model sediment, crust, and mantle V_S and layer thicknesses are randomly perturbed by drawing from
230 a zero-mean normal distribution with standard deviation of 10% of the reference values for the sediment,
231 crust, and mantle. This process is repeated 100 times at each grid point. Next, we repeat the entire inversion
232 processes for a 3-D starting model constructed from line 2, which is parallel to line 1 but located to the south
233 (*Shuck et al.*, 2019). This results in an ensemble of 200 models from which the median and middle 68% of
234 models with $\chi^2 \leq 1.75$ are taken as the preferred V_S and associated $1-\sigma$ uncertainties.

235 **3 Results**

236 **3.1 Phase velocity maps**

237 Figure 8 shows isotropic phase velocity maps for periods ranging from 13–32 s. The dashed grey boundary
238 shows the resolving limit defined by a value of 0.1 on the diagonals of the frequency-averaged resolution
239 matrix. At 13 s period (Figure 8a), Rayleigh wave velocities decrease from west to east due to the increasing
240 water depth. At longer periods this general trend reverses, and velocities increase eastward as a result
241 of the transition from thick continental crust in the west to thinner oceanic crust (and increased mantle
242 sensitivity) in the east. This same character is seen in the synthetic phase velocity maps in Figure 6a,c (see

243 also Figure S5). Similar to the synthetic maps, variations parallel to the margin are smooth and relatively
244 minor, suggesting that the measurements are largely 2-D and primarily controlled by structure associated
245 with the continent-to-ocean transition. At the longest periods (27–32 s), the highest velocities (>4.1 km/s)
246 are observed east of the BSMA, although resolution is limited by modest station distribution east of the
247 BSMA.

248 3.2 Azimuthal anisotropy

249 Azimuthal anisotropy of phase velocity is estimated from 17–32 s, sensitive primarily to mantle lithosphere
250 depths. Figure 9 shows 2-D maps of azimuthal anisotropy for two different choices of smoothing: (1) A
251 smoothly parameterized inversion that seeks to minimize the second spatial derivatives $\nabla^2 A_c$ and $\nabla^2 A_s$. (2)
252 An inversion that effectively solves for a single anisotropy fast azimuth and magnitude on either side of the
253 BSMA by strictly enforcing $\nabla A_c = \nabla A_s = 0$ but breaking this constraint at the BSMA.

254 Both inversions indicate dominantly margin-parallel anisotropy across the array. In detail, the smooth model
255 (Figure 9a) shows a gradual west-to-east clockwise rotation from margin-parallel. East of the BSMA and for
256 17–25 s period, both models indicate a ~ 25 – 45° rotation from margin-parallel towards the FSD (Figure 10).
257 Neither inversion suggests FSD-parallel anisotropy anywhere in the model. At periods >27 s, both models
258 indicate a reduction in anisotropy east of the BSMA. The smooth model shows evidence of a southeastward
259 reduction in anisotropy magnitude west of the BSMA, ranging from ~ 3 – 4% in the northwest to ~ 1 – 2% in
260 the southeast; however, these magnitude variations may not be well resolved (see Section 3.2.1; Figure S2a)

261 As the 2-D anisotropy maps suggest a relatively simple pattern dominated by margin-parallel anisotropy,
262 we also solve for a 1-D array-averaged fast azimuth and magnitude for the region. Phase velocity residuals
263 calculated relative to the 2-D path-integrated isotropic velocities are shown as a function of azimuth in
264 Figure 11a–f. Azimuthal patterns are clearly dominated by a 2θ sinusoid with fast azimuth parallel to
265 the margin and perpendicular to the FSD (Figure 11g,h). Anisotropy peak-to-peak magnitudes decrease
266 slightly with increasing period from 3–4% at 17 s to 2–2.5% at 32 s. To address potential bias caused by
267 uneven azimuthal sampling, we also estimate anisotropy strength and fast azimuth for data averaged in 20°
268 azimuthal bins. The fast azimuths and magnitudes obtained agree with the unbinned estimates but have
269 uncertainties of a factor ~ 2 – 3 larger.

270 We explore whether the simpler 1-D anisotropy model can explain the data as well as the 2-D models. Phase
271 delay residuals are compared for the three anisotropy inversions with varying degrees of freedom – the 1-D

272 array-averaged inversion, 2-D smooth inversion, and 2-D break at the BSMA (Figure 12). Overall, the 2-
273 D smooth inversion yields slightly lower root-mean-square (RMS) phase delay residuals on average, which
274 is unsurprising given that it has the most degrees of freedom. However, for most frequencies the data fits
275 produced by each of the three inversions are nearly indistinguishable (Figure 12g), and therefore we favor the
276 simplest model with 1-D margin-parallel anisotropy, though we cannot rule out a small ($25\text{--}45^\circ$) clockwise
277 rotation east of the BSMA.

278 **3.2.1 Synthetic recovery tests: resolving anisotropy east of the BSMA**

279 We perform synthetic tests to determine whether our dataset can resolve a 90° change in anisotropy fast
280 direction from margin-parallel west of the BSMA to FSD-parallel east of the BSMA (Figure S1–S4). At each
281 frequency, the synthetic dataset is calculated using the synthetic isotropic phase velocity maps extrapolated
282 from Line 1 and the anisotropic model shown in Figure S1. The input anisotropy model is characterized by a
283 constant 2% zero-to-peak anisotropy magnitude across the array with fast azimuthal parallel to the margin
284 west of the BSMA and parallel to the FSD east of the BSMA. Gaussian noise with standard deviation of 0.6 s
285 is added to the synthetic phase delay times. We test the two different 2-D inversion strategies introduced
286 above to evaluate their ability to recover the change in anisotropy direction across the BSMA.

287 We find that the smooth anisotropy inversion (Figure S2) successfully resolves the margin-parallel anisotropy
288 west of the BSMA and FSD-parallel anisotropy east of the BSMA but has trouble capturing the sharp change
289 directly at the BSMA, where recovered anisotropy magnitudes are small. Anisotropy strength is generally
290 poorly resolved in the smooth model; it is overestimated at the westernmost edge of the array and under-
291 estimated east of the BSMA. The inversion containing the break at the BSMA (Figure S3) more accurately
292 captures both the anisotropy magnitude and change in direction across the BSMA. This is expected given
293 that this parameterization closely resembles the character of the input anisotropic model.

294 Overall, both parameterizations are able to capture a 90° change in anisotropy across the BSMA, but the
295 parameterization that explicitly allows a break at the BSMA more accurately recovers the magnitude of
296 anisotropy (Figure S4). In both cases, the array-averaged anisotropy is dominated by the margin-parallel
297 anisotropy structure west of the BSMA, due to its larger footprint and the greater number of ray paths on the
298 western side of the array. We conclude that if a change in anisotropy direction occurs from margin-parallel
299 west of the BSMA to FSD-parallel east of the BSMA, the 2-D anisotropy inversions should recover it. It
300 follows that we can confidently state that FSD-parallel fabric is not present within the ENAM-CSE region.

301 **3.3 3-D shear velocity model**

302 **3.3.1 V_S uncertainty**

303 Example 1-D V_S profiles from the west and east are shown in Figure 13 with uncertainties estimated via the
304 Monte Carlo inversion described in Section 2.5. Uncertainties are larger in the crust and sediments than in
305 the mantle, particularly for the western profile (Figure 13a,b) where the rapid change in water depth makes
306 resolving crustal structure more difficult. In the western profile, crustal uncertainties generally range from
307 0.5–0.55 km/s compared to 0.25–0.3 km/s in the east. Mantle velocities for both the western and eastern
308 profiles appear well resolved with uncertainties ranging from 0.1–0.3 km/s in the upper 10 km of the mantle
309 and 0.03–0.1 km/s below (see also Figure S7).

310 **3.3.2 Crustal variations**

311 Horizontal slices through the 3-D V_S model are shown in Figure 14. In the upper 4 km of the crust, velocities
312 are slow (2–2.5 km/s) in the west corresponding to sediments and transition sharply to faster velocities (3–
313 3.6 km/s) as the sediments thin to <4 km in the east (*Shuck et al.*, 2019). The sharpness of this basement
314 transition is a product of the layer discontinuities built into the starting reference model (Figure S6), which
315 are preserved but otherwise poorly resolved by Rayleigh waves. In the lower crust (8 km), a slow anomaly
316 is observed just east of the ECMA.

317 In the 4–15 km depth slices, a band of anomalously slow velocities coincides with the shelf, correlating with
318 large gradients in seafloor depth (>2 km/°). This rapid decrease in water depth west of the shelf break
319 is just outside the array and poorly resolved by our relatively smooth phase velocity maps. This results
320 in slower than predicted phase velocities along the shelf that map directly into the crustal velocities. The
321 region of the model most affected by this bias is indicated by a semi-transparent mask in Figure 14, and we
322 do not interpret structure within that part of the model.

323 **3.3.3 Mantle variations**

324 Velocities in the mantle are generally well resolved in comparison to the crust as the mantle-sensitive Rayleigh
325 waves are not sensitive to changes in water depth. Within the array footprint, mantle velocities range from
326 ~ 4.4 km/s in the west to ~ 4.7 km/s in the east (Figure 14d–e). East of the BSMA and within the array

327 footprint, velocities peak at 4.6–4.7 km/s in the upper 10–25 km of the mantle and decrease with depth,
328 suggestive of thermal ocean lithosphere. At 55–60 km depth, velocities are nearly homogeneous (~ 4.5 km/s)
329 across the region.

330 4 Discussion

331 4.1 2-D margin-perpendicular structure

332 The average 2-D margin-perpendicular structure of our model shows more clearly the contrast in shallow
333 mantle velocities that coincides approximately with the BSMA (Figure 15). In the upper 15–20 km of the
334 mantle, relatively slow shear velocities (4.4–4.55 km/s) extend east of the margin ~ 200 km terminating
335 at the BSMA. East of the BSMA, velocities are elevated (>4.6 km/s) and characteristic of depleted, cold
336 oceanic lithosphere (e.g. *Lin et al.*, 2016, *Nishimura and Forsyth*, 1989). Below 40–50 km depth, velocities
337 are mostly homogeneous across the region, and therefore we interpret that the low-velocity lid within the
338 IMQZ is underlain by mantle lithosphere with temperature and composition consistent with the adjacent
339 oceanic lithosphere.

340 The average margin-perpendicular V_S structure in our model is broadly similar to the offshore structure in
341 the previous shear velocity model of *Lynner and Porritt* (2017). However, upon closer inspection important
342 differences do exist. Their inferred offshore crust is ~ 10 km thicker than what we observe. They also do
343 not observe low velocities in the uppermost mantle west of the BSMA, nor do they observe a change in
344 mantle velocities across the BSMA. These differences are likely attributed primarily to our differing starting
345 model assumptions, and in particular, our direct incorporation of the crustal constraints from *Shuck et al.*
346 (2019). In addition, our phase velocity maps explicitly account for azimuthal anisotropy and therefore should
347 more accurately capture the true variations in isotropic phase velocity. Finally, that our modeling is focused
348 exclusively on the offshore region rather than the full margin-crossing ENAM region both simplifies the
349 modeling and avoids potential smearing of continental structure into the oceanic domain.

350 4.2 Low velocity mantle lid west of the BSMA

351 We explore two hypotheses for explaining the 15–20 km thick low velocity mantle lid west of the BSMA.
352 First, velocities of the lid are similar to continental mantle velocities previously imaged just onshore (*Shen*

353 and Ritzwoller, 2016) (Figure 16a) and therefore may correspond to stretched and/or thermo-chemically
 354 modified continental lithosphere associated with rifting (e.g. Hopper *et al.*, 2020). This interpretation of
 355 the mantle structure is consistent with the detailed crustal structure recently imaged at the ENAM-CSE
 356 (Bécel *et al.*, 2020, Shuck *et al.*, 2019). Shuck *et al.* (2019) observed relatively thin crust (6–8 km) with
 357 fast lower crustal velocities ($V_P \sim 7.5$ km/s) on Lines 1 and 2 between the ECMA and BSMA that is best
 358 explained by deeper than usual mantle melting. This led them to predict the presence of a 15–20 km thick
 359 continental lithosphere that truncated the upper melting regime producing less voluminous, more mafic
 360 melts. The slower shear velocities ($V_S < 4.55$ km/s) that we observe in the upper mantle west of the BSMA
 361 are consistent with this prediction.

362 Second, the slow lithospheric velocities could be explained by gabbroic inclusions trapped within oceanic
 363 mantle lithosphere during ultra-slow spreading (e.g. Lizarralde *et al.*, 2004). Efficient conductive cooling at
 364 the MOR during periods of very slow spreading leads to a thicker mantle lid that could crystallize gabbroic
 365 melts and trap them in the mantle, reducing shear velocity and producing thinner crust overall. West of the
 366 BSMA, basement roughness is characteristic of typical ultra-slow spreading crust with a half-spreading rate
 367 of ~ 0.65 cm/yr (Bécel *et al.*, 2020). This offers an alternative mechanism for truncating the upper melting
 368 regime without invoking the emplacement of proto-oceanic crust on top of continental lithosphere. However,
 369 one possible caveat is that the notion of a thicker thermal lid may be difficult to reconcile with the relatively
 370 hot mantle potential temperatures of 1395–1420°C that have been inferred based on the crustal structure in
 371 the region (Shuck *et al.*, 2019), but detailed modeling of thermal evolution of ultra-slow spreading centers in
 372 the presence of elevated mantle temperatures is needed to evaluate temperature tradeoffs.

373 We evaluate whether the addition of gabbro to nominal oceanic mantle (second hypothesis) can explain the
 374 slow V_S west of the BSMA (Figure 16). We start by defining a reference oceanic V_S profile as the average
 375 east of the BSMA. A nominal gabbro V_P profile is estimated following Carlson and Miller (2004):

$$V_P^{\text{gabbro}}(z) = V_{P_0}^{\text{gabbro}} + \left(\left. \frac{\partial V_P}{\partial z} \right|_P + \left. \frac{\partial V_P}{\partial z} \right|_T \right) (z - z_{\text{moho}}) \quad (6)$$

376 where $V_{P_0}^{\text{gabbro}} = 6.85$ km/s is the reference velocity of gabbro at the Moho (z_{moho}) and the change in V_P due
 377 to pressure and temperature effects at depth are $(\partial V_P / \partial z)|_T = 0.02$ s⁻¹ and $(\partial V_P / \partial z)|_P = -0.01$ s⁻¹. This
 378 profile is converted to V_S assuming $V_P / V_S = 1.8$. Next, using the Voigt average we solve for the proportion
 379 of gabbro required to reduce velocities of the oceanic reference model to match velocities observed west of the
 380 BSMA. The resulting depth distribution of gabbro in Figure 16b shows that up to $\sim 35\%$ gabbro by volume

381 would be required to explain velocities west of the BSMA for an oceanic reference model. Integrating this
382 gabbro distribution implies crustal thinning west of the BSMA of ~ 4 km (assuming crust that is pure gabbro),
383 which is more thinning than observed (0.7–1.8 km) in the reflection and refraction imaging (*Bécel et al.*, 2020,
384 *Shuck et al.*, 2019) (Figure 16c). In contrast, for a continental reference profile calculated from the onshore
385 structure of *Shen and Ritzwoller* (2016), only up to $\sim 15\%$ gabbro and ~ 1 km of thinning are required, which
386 is consistent with the crustal observation. Though simple, this back-of-the-envelope calculation suggests a
387 combination of the two proposed hypotheses: slow mantle velocities west of the BSMA are due to extended
388 continental lithosphere that contains up to $\sim 15\%$ gabbro retained in the mantle during the volcanic process
389 that built the overlying basaltic crust.

390 4.3 Implications for the onset of normal seafloor spreading

391 The 15–20 km thick slow lithospheric lid west of the BSMA that we attribute to continental mantle has
392 implications for the timing and mechanics of breakup and onset of normal seafloor spreading in the North
393 Atlantic. In keeping with *Bécel et al.* (2020), *Shuck et al.* (2019), our results suggests that complete breakup
394 and the onset of normal seafloor spreading occurred at ~ 170 Ma, approximately 10–25 Myr after the initia-
395 tion of rifting and formation of the ECMA. Crust that was formed following emplacement of the BSMA is
396 consistent with normal seafloor spreading conditions (*Shuck et al.*, 2019), and mantle structure is character-
397 istic of typical oceanic lithosphere (Figure 16a). The BSMA also marks an increase in spreading rate from
398 ultra slow (~ 0.65 cm/yr half-rate) to slow (~ 1.3 cm/yr half-rate) as inferred by the decrease in basement
399 roughness stepping eastward across the BSMA (*Bécel et al.*, 2020).

400 Previously, complete breakup and the onset of normal seafloor spreading was thought to have taken place over
401 a relatively short-lived period that immediately followed formation of the ECMA (*Holbrook and Kelemen*,
402 1993, *Holbrook et al.*, 1994, *Kelemen and Holbrook*, 1995). Rather, the detailed lithospheric architecture
403 at the ENAM-CSE indicates that after formation of the ECMA and prior to the BSMA, a proto-seafloor
404 spreading mode was active with mantle-derived melts migrating vertically through the continental lid forming
405 thin proto-oceanic crust (*Bécel et al.*, 2020, *Shuck et al.*, 2019). Mantle velocities suggest that up to $\sim 15\%$
406 gabbro was crystallized and retained in the lid during melt migration, likely in channelized structures (e.g.
407 *Katz and Weatherley*, 2012). The potential for melts to remain trapped in the mantle has implications
408 for our understanding of melt extraction, crustal accretion, and thermal evolution at ultra-slow spreading
409 centers.

410 Numerical models have shown that mantle can remain intact even after the crust has separated at rifted
411 margins if the crust fails through brittle faulting while the underlying mantle lithosphere deforms through
412 ductile necking (*Huismans and Beaumont, 2011*). Such ductile deformation of the lithosphere is conceivable
413 considering the elevated mantle potential temperatures (1430–1480°C) inferred from CAMP lavas (*Callegaro*
414 *et al., 2013*) and the inferred ultra-slow extension rates (<2 cm/yr half-rate) (*Bécel et al., 2020, Davis et al.,*
415 *2018*).

416 **4.4 Comparison to previous anisotropy observations**

417 The Rayleigh-wave anisotropy that we observe in the lithosphere at the ENAM-CSE has a fast direction
418 sub-parallel to the margin. Previous shear-wave splitting observations at the ENAM-CSE also showed
419 margin-parallel anisotropy, but it was attributed to deeper present-day asthenospheric flow along the margin
420 associated either large-scale density or pressure driven flow or edge-driven convection due to the large litho-
421 spheric root of the continent (*Lynner and Bodmer, 2017*). Our observations do not preclude such flow in
422 the asthenosphere but show that strong margin-parallel anisotropy is present in the lithosphere, potentially
423 contributing at least in part to the previous shear-wave splitting observations. A preliminary comparison
424 of shallow mantle V_P along the margin-parallel and -perpendicular refraction lines at ENAM also indicated
425 ($\sim 8\%$) faster velocities parallel to the margin (*Shuck and Van Avendonk, 2016*) in agreement with our
426 observations, though spatially limited to the crossing points of the profiles.

427 At the Far-offset Active-source Imaging of the Mantle (FAIM) seismic refraction experiment (115–130 Ma)
428 ~ 800 km southeast of the ENAM-CSE, *Gaherty et al. (2004)* inferred a lithospheric olivine LPO sub-parallel
429 to the FSD (to within $\sim 15^\circ$). This implies that a rotation in lithosphere LPO from spreading-perpendicular to
430 spreading-parallel occurred between 165–130 Ma. Observations of Love-Rayleigh scattering offshore ENAM
431 offers evidence for such lateral gradients in anisotropy (*Servali et al., 2020*), though the frequencies considered
432 (~ 100 s) have significant asthenospheric sensitivity, and the inferred scattering points are widely distributed
433 throughout the western Atlantic including within the ENAM-CSE footprint. Additionally, global models
434 show variable anisotropy at lithospheric depths in the North Atlantic that is often highly rotated from
435 the FSD (e.g. *Becker et al., 2014, Debayle and Ricard, 2013, Schaeffer et al., 2016, Yuan and Beghein,*
436 *2013*). As structure in the Atlantic varies significantly between different global models, our high-resolution
437 estimate provides a useful benchmark for the northwest Atlantic and confirms the previous notion of variable
438 anisotropy in the Atlantic lithosphere.

439 4.5 Interpretation of fossilized margin-parallel anisotropy

440 4.5.1 Plate-motion controlled LPO

441 The surface-wave anisotropy that we observe is similar in strength and overall character to that observed in
442 other oceanic regions (e.g. *Eddy et al.*, 2019, *Forsyth*, 1975, *Lin et al.*, 2016, *Nishimura and Forsyth*, 1989,
443 *Russell et al.*, 2019, *Takeo et al.*, 2018), suggesting that olivine LPO is a likely mechanism for explaining
444 the anisotropy. However, the margin-parallel orientation is $\sim 90^\circ$ rotated from what is expected for a typical
445 seafloor spreading environment, where corner flow near the ridge generates spreading-parallel olivine LPO
446 that is locked into the lithosphere (e.g. *Blackman and Kendall*, 2002). Instead, our results suggest margin-
447 parallel shear deformation during continental breakup and initial seafloor spreading. Based on recent seismic
448 anisotropy observations in Pacific and Juan de Fuca lithosphere that deviate from the FSD by 10–70° (e.g.
449 *Shinohara et al.*, 2008, *Shintaku et al.*, 2014, *Takeo et al.*, 2016, 2018, *Vanderbeek and Toomey*, 2017),
450 it has been suggested that shear deformation associated with APM at the time of plate formation may
451 dominate the spreading-related deformation if the absolute plate velocities outpace the spreading rate at the
452 MOR (*Vanderbeek and Toomey*, 2017). This may explain the overall poor correlation between lithosphere
453 anisotropy and FSD in the slow-spreading Atlantic (*Becker et al.*, 2014). At ENAM, estimates of spreading
454 rate prior to the BSMA formation are ultra-slow (~ 0.65 cm/yr half-rate), and increase to ~ 1.3 cm/yr
455 half-rate just after BSMA formation (*Bécel et al.*, 2020).

456 We explore whether our observations of margin-parallel anisotropy can be explained by fast margin-parallel
457 plate velocities at the time of ENAM formation (165–200 Ma) using four recent plate reconstruction models
458 via the GPlates software (*Boyden et al.*, 2011): S12-ESR (*Seton et al.*, 2012); M16-AREPS (*Müller et al.*,
459 2016); M16-GPC (*Matthews et al.*, 2016); M19-T (*Müller et al.*, 2019) (Figure 17). These studies utilize
460 continuously closing topological plate polygon networks that account for inception and cessation of plate
461 boundaries in order to reconstruct plate motions from present day back ~ 200 Ma. Relative plate motions
462 in these models are well constrained primarily by seafloor magnetic anomaly picks and fracture zones,
463 particularly in the Atlantic where seafloor is preserved on both conjugate flanks of the MOR. However,
464 absolute plate motions are less well constrained and depend strongly on the choice of absolute reference
465 frame (e.g. *Shephard et al.*, 2012, *Torsvik et al.*, 2008), which varies between studies. Evidence for this is
466 seen in Figure 17, where North America plate motion varies strongly between the four models, particularly
467 for ages > 160 Ma. In general, the four models use hybrid reference frames with present day to 70–100 Ma
468 described by a global moving hotspot reference frame and 100–230 Ma constrained by true-polar wander

469 corrected paleomagnetic data. The exception is M19-T, for which authors inverted for a reference frame from
470 0–80 Ma that minimized trench migration velocities and global net lithospheric rotation, and this model is
471 the only one which includes diffuse deformation along plate boundaries.

472 Though differences in plate motion do exist between the four models, important similarities are observed.
473 Absolute plate velocities during the ENAM breakup were variable but often fast (1.5–9 cm/yr) relative to
474 spreading (0.5–1.5 cm/yr half-rate), especially early in the breakup (180–200 Ma). From 170–200 Ma, APM
475 directions were significantly different from the spreading direction and often similar in azimuth to ENAM
476 anisotropy (up to $\pm 90^\circ$ rotated from spreading) for models S12-ESR, M16-AREPS, and M16-GPC. A $\sim 180^\circ$
477 reversal in plate direction from approximately north along the margin to south is accompanied by a drop
478 (and shortly followed by an increase) in plate velocity in all models except M19-T. This abrupt change in
479 plate direction occurs at ages ranging from 170–190 Ma, depending on the model. Although its origin is not
480 well understood, it roughly correlates in time with far-field plate reorganization processes such as opening of
481 the Gulf of Mexico as well as the subduction polarity reversal from west-dipping to east-dipping across the
482 Wrangellia Superterrane prior to its collision with western North America (*Shephard et al.*, 2013). As there
483 is a 180° ambiguity in the interpretation of flow direction for a given orientation of seismic anisotropy, both
484 orientations of margin-parallel plate motion are consistent with our observations.

485 At around the time of BSMA formation (~ 170 Ma), the spreading rate increased and plate velocities rotated
486 closer to the FSD, on average (Figure 17). Therefore, fabric east of the BSMA likely represents LPO
487 influenced both by spreading and APM related shearing. This is consistent with the $\sim 25\text{--}45^\circ$ clockwise
488 rotation in anisotropy east of the BSMA suggested by our 2-D inversions (Figure 9). Although the data do
489 not strictly require this rotation of anisotropy east of the BSMA (Figure 12), it cannot be ruled out. The
490 fast azimuths east of the BSMA and within the ENAM-CSE footprint (165–170 Ma) are not FSD-parallel,
491 indicating that the fabric was still modified by absolute plate motion even when spreading rates were on the
492 order of APM velocities. East of the ENAM-CSE at the younger FAIM experiment (115–130 Ma), APM
493 velocities were fast (2.5–6.5 cm/yr) relative to spreading (1–2.5 cm/yr half-rate) and approximately parallel
494 to the FSD, in agreement with the observed FSD-parallel anisotropy from *Gaherty et al.* (2004). Given
495 that both the ENAM and FAIM anisotropy observations correlate with APM at the time of spreading, our
496 preferred interpretation is that shear associated with APM dominates LPO formation when plate velocities
497 outpace the spreading rate.

498 4.5.2 Alternative olivine LPO fabric type

499 Anisotropy that is oriented perpendicular to the spreading direction could instead be indicative of an alter-
500 native olivine LPO fabric type. It is typically assumed that olivine with an A-type LPO dominates in the
501 upper mantle with fast [100] aligning parallel to the shear direction and slow [010] oriented perpendicular
502 to the shear plane (*Zhang and Karato, 1995*). However, recent laboratory experiments on olivine have doc-
503 umented several alternative fabric types with various orientations of the crystallographic axes with respect
504 to the shear plane (*Bystricky et al., 2000, Jung and Karato, 2001, Jung et al., 2006, Karato et al., 2008,*
505 *Katayama et al., 2004*).

506 A commonly invoked mechanism for explaining anisotropy perpendicular to shear is the so-called “a-c switch”
507 (*Holtzman et al., 2003, Qi et al., 2018*). Experiments on melt-bearing olivine show segregation of melt into
508 bands that leads to strain partitioning that produces LPO with [100] primarily oriented perpendicular to
509 the direction of shear and in the shear plane (*Holtzman et al., 2003*). Such a fabric could conceivably be
510 responsible for the margin-parallel anisotropy at ENAM, particularly if large volumes of melt were present.
511 However, in their more recent experiments on partially molten samples, *Qi et al. (2018)* emphasize that
512 melt tends to produce LPO with girdled [100] and [001] (i.e. fast axis dispersed within the shear plane),
513 which has only a slightly stronger [100] alignment perpendicular to shear but is significantly weaker than
514 melt-free olivine LPO. Instead, they predict that the a-c switch would most likely be observed seismically
515 as a transversely isotropic structure with very weak azimuthal anisotropy. As the Rayleigh-wave anisotropy
516 we observe is quite strong (2–4%), this mechanism is unlikely to explain our observations.

517 A second possibility is the presence of B-type LPO with fast [100] aligning perpendicular to shear and in the
518 shear plane (*Jung and Karato, 2001*). B-type LPO has been produced in the laboratory at low temperatures
519 (high stress) and moderate water content (*Jung and Karato, 2001*) and has been invoked at subduction zones
520 to explain trench-parallel anisotropy observed in the cold mantle wedge (*Karato et al., 2008, Nakajima and*
521 *Hasegawa, 2004*). However, given the relatively hot mantle potential temperatures estimated during ENAM
522 formation (1395–1420°C) (*Shuck et al., 2019*), B-type LPO is unlikely.

523 4.5.3 Contribution of the continental lid

524 The inferred gabbroic inclusions frozen within the thin continental lid, if preferentially oriented, could also
525 contribute to the observed margin-parallel anisotropy through a shape-preferred fabric (e.g. *Holtzman et al.,*
526 *2003, Kawakatsu et al., 2009, Schlue and Knopoff, 1976*). Shape-preferred models have been invoked to

527 explain margin-parallel structures in active rift environments (e.g. *Kendall et al.*, 2005), where the large
528 velocity contrast between melt-rich dikes and surrounding country rock can produce a significant apparent
529 anisotropy. For the thermal state of the ancient ENAM margin, the available velocity contrast between re-
530 tained gabbro and surrounding peridotite is much smaller. We calculate the equivalent shear-wave anisotropy
531 of vertical gabbro sheets embedded within a background mantle model via Backus averaging valid for seis-
532 mic wavelengths much greater than the gabbro channel width (*Backus*, 1962). We calculate shear-wave
533 azimuthal anisotropy (G) for a vertically oriented alternating layered structure with the relative proportions
534 of gabbro from Figure 16b and find anisotropy of $<1\%$ for the continental reference model. Anisotropy peaks
535 directly beneath the Moho because that is where gabbro is most abundant in the model. The Rayleigh-wave
536 anisotropy predicted from the depth-dependent G anisotropy model is $<0.5\%$ for the continental reference
537 model, which is far too weak to explain the 2–4% anisotropy we observe (Figure 18b). We also carry out the
538 calculation for gabbro in the oceanic reference model, which results in slightly stronger G anisotropy of up
539 to $\sim 2\%$, but the predicted Rayleigh-wave anisotropy is still much too weak ($<1\%$).

540 We conclude that while shape-preferred anisotropy due to aligned gabbro channels within the lithospheric lid
541 may slightly contribute to the margin-parallel anisotropy we observe, it is not strong enough to explain the
542 entire signal. Additionally, our inversion tests evaluating 2-D variations in anisotropy suggest a continuity in
543 structure across the BSMA with at most a subtle rotation, arguing against a purely continental-lithosphere
544 origin.

545 Given the strong relic margin-parallel LPO inferred from shear-wave splitting studies of the eastern U.S.
546 with splitting that largely parallels the structural grain of the Appalachians (e.g. *Barruol et al.*, 1997a,b,
547 *Long et al.*, 2016, *Wagner et al.*, 2012), it may be tempting to attribute margin-parallel anisotropy west of
548 the BSMA to a preserved accreted terrain. However, predominantly null shear-wave splitting observed in
549 North Carolina directly onshore the ENAM-CSE can be interpreted as negligible (or vertical) lithospheric
550 LPO (*Lynner and Bodmer*, 2017). Furthermore, Rayleigh-wave anisotropy of the onshore region shows
551 weak and laterally variable anisotropy in the lithosphere (*Wagner et al.*, 2018), in contrast to the regionally
552 consistent anisotropy we observe. Olivine LPO formed during collisional orogenic processes would likely be
553 disrupted and altered by later extensional deformation during rifting (*Barruol et al.*, 1997a), and therefore,
554 it is unlikely that such a strong, regionally coherent LPO would remain intact at the ENAM-CSE.

555 We favor the interpretation that the lithospheric olivine LPO retains a record of strong margin-parallel
556 APM during early ultra-slow spreading of the Atlantic (Figure 19). This explanation satisfies anisotropy
557 observations at both the ENAM-CSE and FAIM experiments separated in age by ~ 50 Myr. It may also

558 explain why we do not observe a significant rotation in anisotropy to FSD-parallel immediately east of the
559 BSMA. It is also likely to produce anisotropy down to at least 60 km depth, which corresponds to the
560 full depth sensitivity of the 17–32 s observations. We cannot rule out an additional small contribution to
561 anisotropy from the thin continental lid between the ECMA and BSMA produced either by preferentially
562 oriented gabbro channels or a relic LPO, but that alone is unable to account for the magnitude and coherence
563 of Rayleigh-wave anisotropy that we observe across the region. Instead, LPO fabric frozen into the lithosphere
564 at slow spreading environments, such as in the Atlantic, is likely to retain a complex deformation signal that
565 records the relative balance between absolute plate motion and seafloor spreading.

566 5 Conclusions

567 We present a shear velocity model of the uppermost mantle lithosphere together with observations of
568 Rayleigh-wave anisotropy offshore the Eastern U.S. that provide new constraints on the late stages of rifting
569 and onset of seafloor spreading at the ENAM (Figure 19). The shear velocity model contains a proto-oceanic
570 domain defined by oceanic crust overlying a 15–20 km thick slow (4.4–4.55 km/s) lithospheric lid interpreted
571 as continental mantle with up to $\sim 15\%$ gabbro extending from the margin ~ 200 km east to the BSMA.
572 East of the BSMA, shallow mantle velocities are fast (> 4.6 km/s) and indicative of more typical oceanic
573 lithosphere, suggesting that the BSMA marks the final breakup of the ENAM from West Africa and onset
574 of normal seafloor spreading ~ 170 Ma, as previously hypothesized (*Bécel et al.*, 2020, *Shuck et al.*, 2019).

575 We observe Rayleigh-wave anisotropy in the lithosphere with a fast direction parallel to the margin, cor-
576 relating approximately with the plate motion direction at the time, rather than the direction of ultra-slow
577 spreading ($\lesssim 2$ cm/yr half-rate). Nearly 800 km southeast of ENAM at ~ 50 Myr younger seafloor, the FAIM
578 experiment showed FSD-parallel anisotropy (*Gaherty et al.*, 2004) that also correlates with the fossil-APM
579 direction. We propose that lithosphere LPO formed at slow-spreading MORs, such as the Atlantic, primarily
580 records mantle shear imparted by absolute plate motion rather than by classic corner flow.

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587 **Data Availability Statement**

588 All broadband waveform data was accessed through the IRIS Data Management Center under network
589 code YO using the open-source Python package ObsPy (<https://docs.obspy.org/>). The script used for
590 retrieving daily waveform data can be found here: https://github.com/jbrussell/fetch_NOISE.

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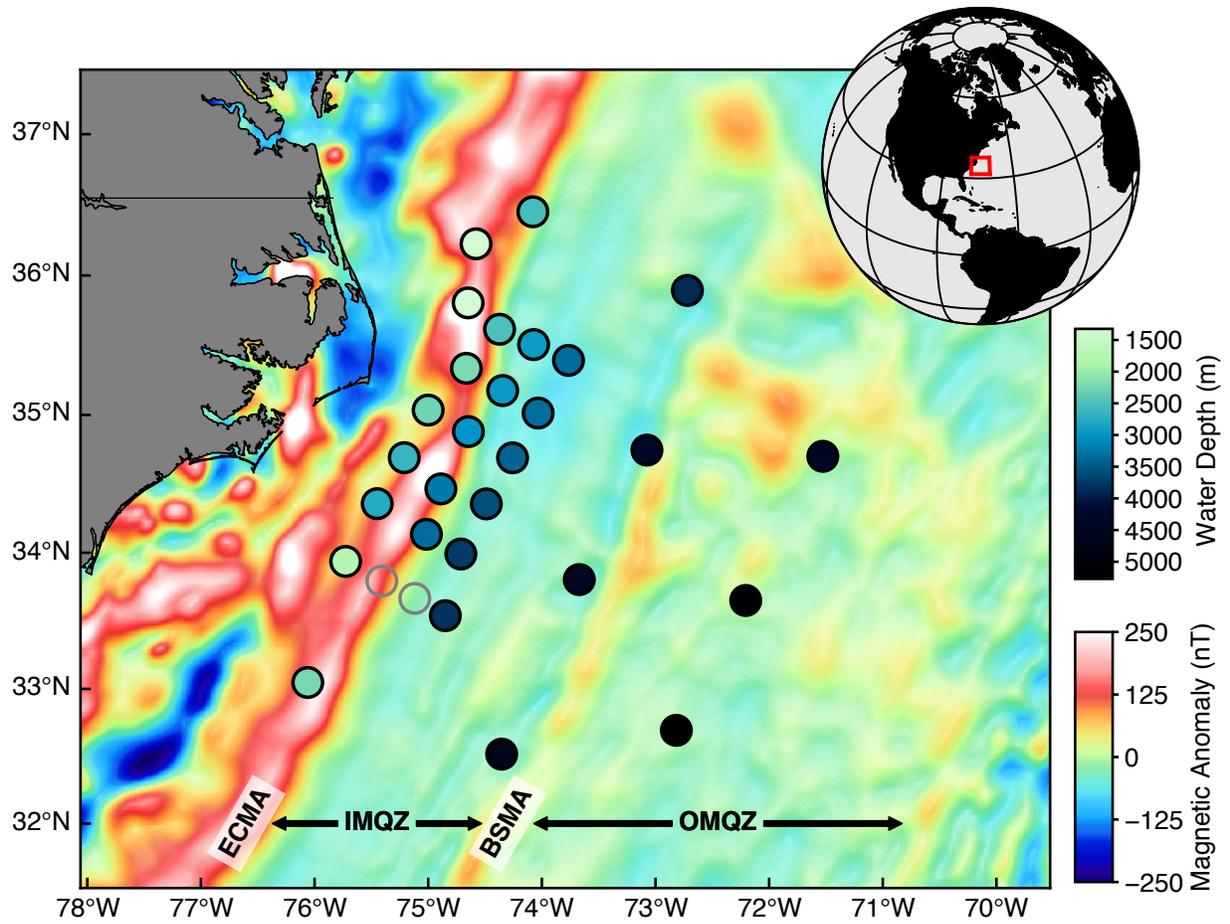


Figure 1: Magnetic anomalies of the ENAM-CSE region with broadband OBS shown as circles and colored by water depth. Open circles denote stations with poor data quality. East Coast Magnetic Anomaly (ECMA); Blake Spur Magnetic Anomaly (BSMA); Inner Magnetic Quiet Zone (IMQZ); Outer Magnetic Quiet Zone (OMQZ)

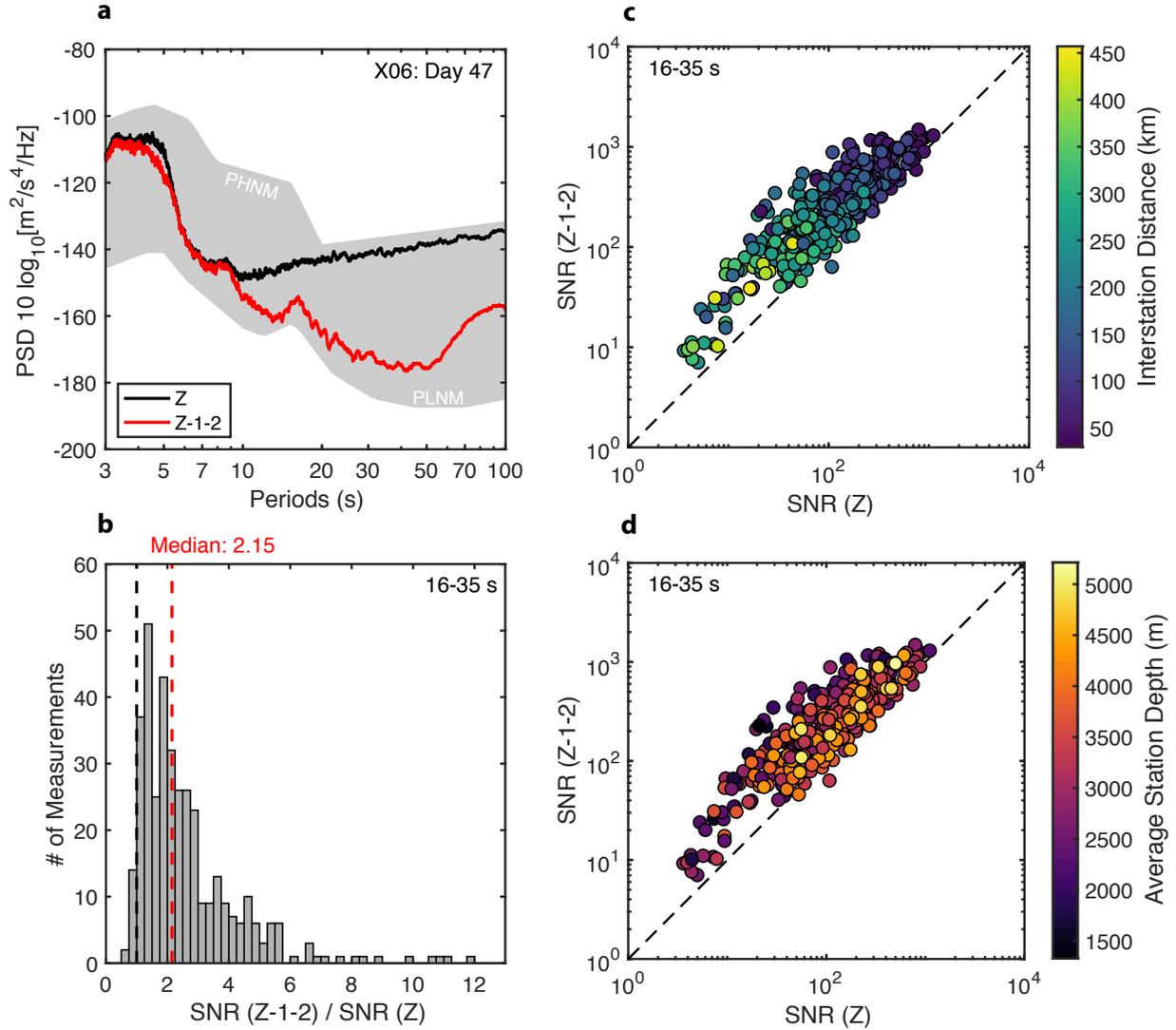


Figure 2: Ambient noise empirical Green's function (EGF) signal-to-noise ratio (SNR) improvement from daily tilt corrections. (a) Example 24 hour smoothed spectra for station X06 on day 47 of the deployment. The raw vertical channel (Z) and tilt-corrected vertical (Z-1-2) are shown. The grey region is bounded below by the Peterson low noise model (PLNM) and above by the high noise model (PHNM) (Peterson, 1993). The primary microseism peak at ~ 16 s period is visible only after the tilt correction is applied. (b) Histogram of tilt-corrected SNR values relative to raw values for all EGFs filtered from 16–35 s period. The red dashed line shows the median value of 2.15 and the black dashed line marks a value of 1 (i.e., no improvement). (c) Comparison of SNR values for each EGF, where the one-to-one line is dashed in black and points are colored by interstation distance. (d) Same as c) but colored by average station water depth. $SNR = RMS(signal)^2 / RMS(noise)^2$, where *signal* is defined as any arrival with group velocity greater than 1.5 km/s.

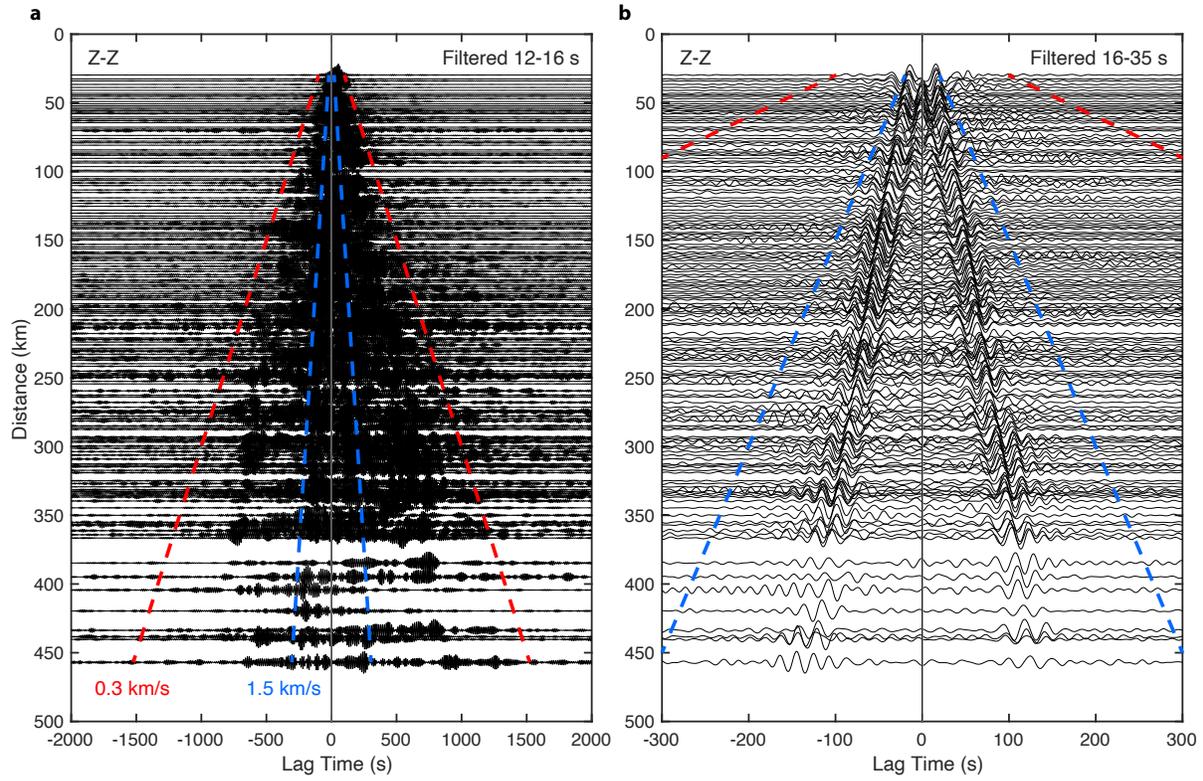


Figure 3: Vertical component ambient noise empirical Green's functions (EGFs) showing Rayleigh wave energy filtered from (a) 12–16 s and (b) 16–35 s period. Dashed blue and red lines indicate group velocities of 1.5 km/s and 0.3 km/s, respectively.

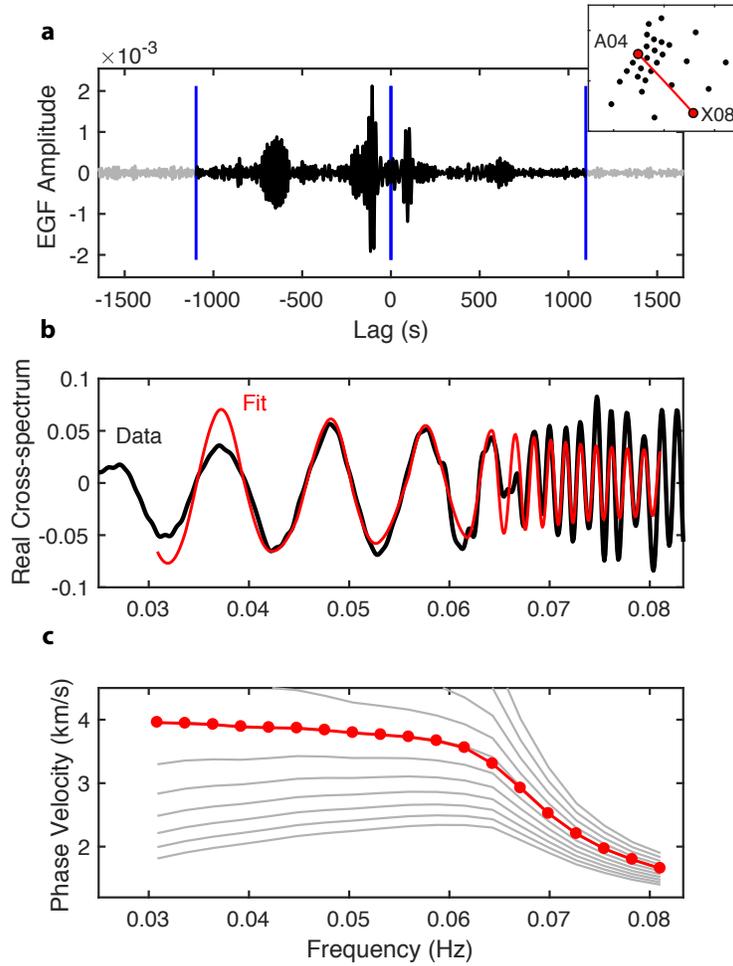


Figure 4: Demonstration of the cross-spectral zero-crossing analysis for phase velocity extraction from station pair A04–X08. (a) EGF cross-correlation in the time domain with station A04 and X08 indicated in red in the inset map separated by 330 km distance. The vertical blue lines mark the 0.3 km/s group velocity window applied prior to dispersion analysis. (b) Real part of the EGF cross-spectrum formed by taking the Fourier transform of a). The data are shown in black and the fit determined from the zero-crossing analysis in red. (c) Interstation dispersion curve extracted from the zero crossings of b). Grey lines show possible dispersion curves. The red points mark the final interpolated dispersion measurements selected based on a nominal mantle velocity of 3.9 km/s.

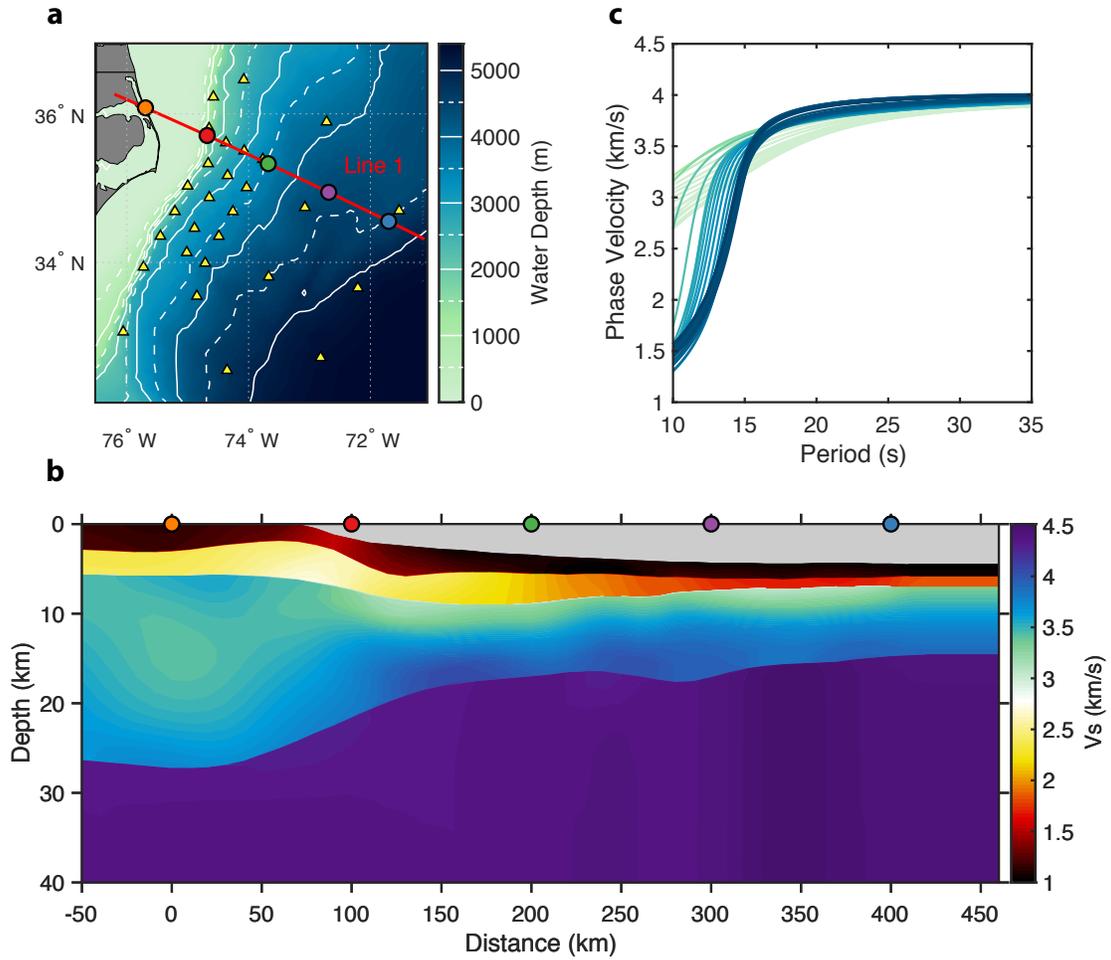


Figure 5: (a) Map of the ENAM region showing Line 1 from *Shuck et al. (2019)* in red. Broadband OBS are shown by yellow triangles. White contours show water depth with 500 m change between dashed and solid lines. Colored circles correspond to those in b). (b) Shear velocity V_S along Line 1 converted from V_P assuming $V_P/V_S = 1.8$ in the crust and mantle and 2.22 in the sediments. (c) Line 1 phase velocities calculated from b) and colored by water depth as in a).

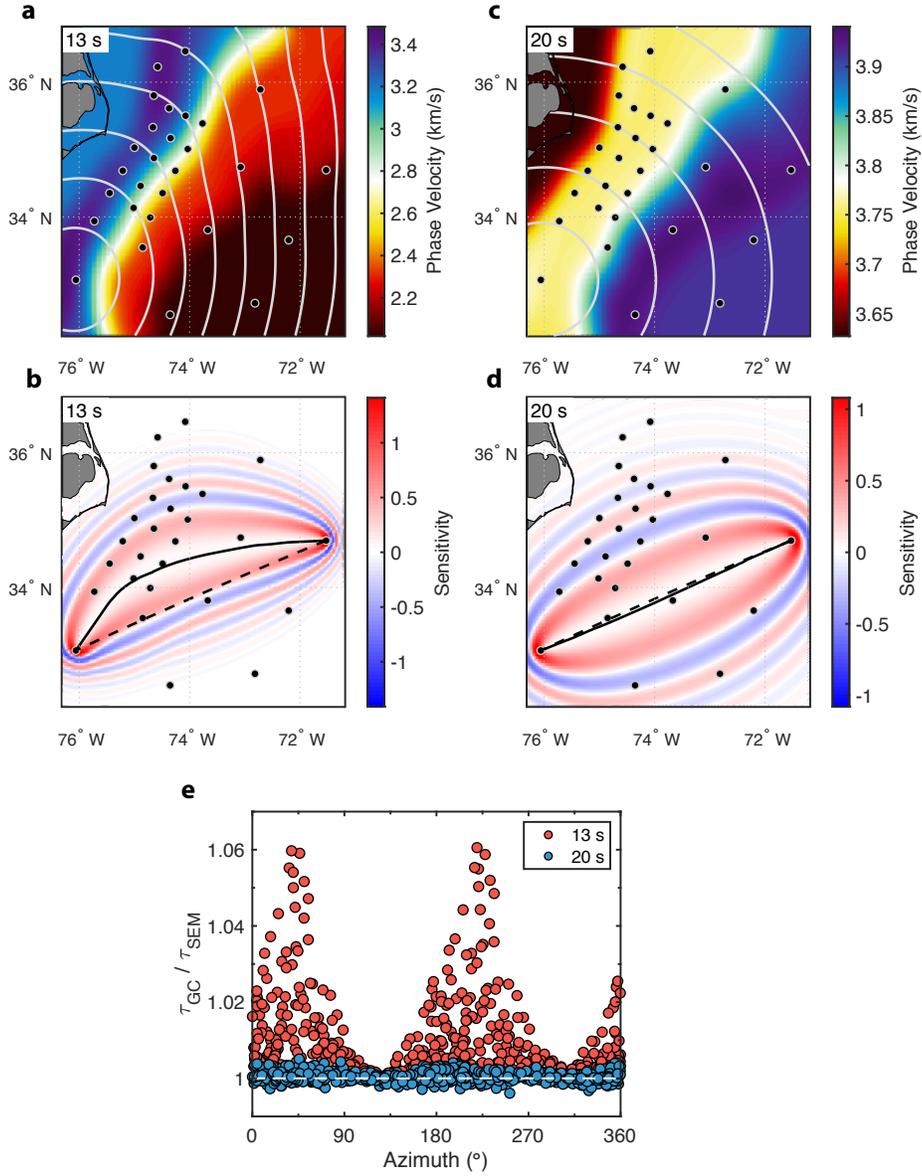


Figure 6: Finite-frequency sensitivity kernels and importance of off-great-circle propagation. (a) Synthetic phase velocity map at 13 s period after Gaussian smoothing ($\sigma = 0.2^\circ$). Grey contours show S-wave phase delay at 25 second intervals due to an impulse at station X10 modeled using the spectral element method (SEM) (i.e., $\tau(\mathbf{x})$ in eq.(4)). Note the strong wavefield distortion due to large lateral velocity gradients associated with changes in water depth. (b) Finite-frequency sensitivity kernel for station pair X10-X04. The dashed and solid black lines show the great circle (GC) and SEM ray paths, respectively. (c,d) same as a), b) but for 20 s period. The SEM ray path is nearly coincident with the GC. (e) Apparent azimuthal anisotropy of GC phase delays (τ_{GC}) relative to SEM phase delays (τ_{SEM}) for all station pairs with $r > 100$ km at 13 s (red) and 20 s (blue).

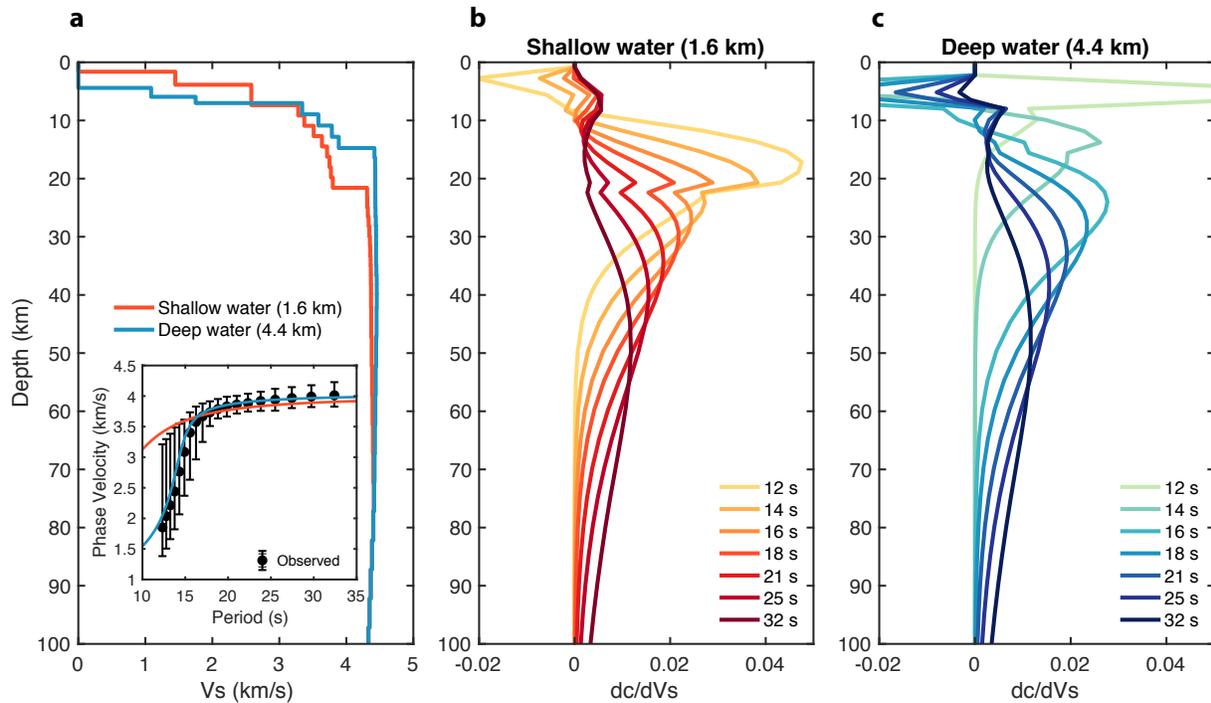


Figure 7: Shear velocity sensitivity kernels for endmember shallow- and deep-water structure along Line 1. (a) V_S profiles for 1.6 km *shallow* water in red (100 km along Line 1, see Figure 5) and 4.4 km *deep* water in blue (400 km along Line 1). The inset shows phase velocity predictions for the two models (colored lines) and the median (black circles) and range (bars) of all interstation phase velocity measurements. (b,c) Phase velocity sensitivities to shear velocity perturbations for shallow- and deep-water structure.

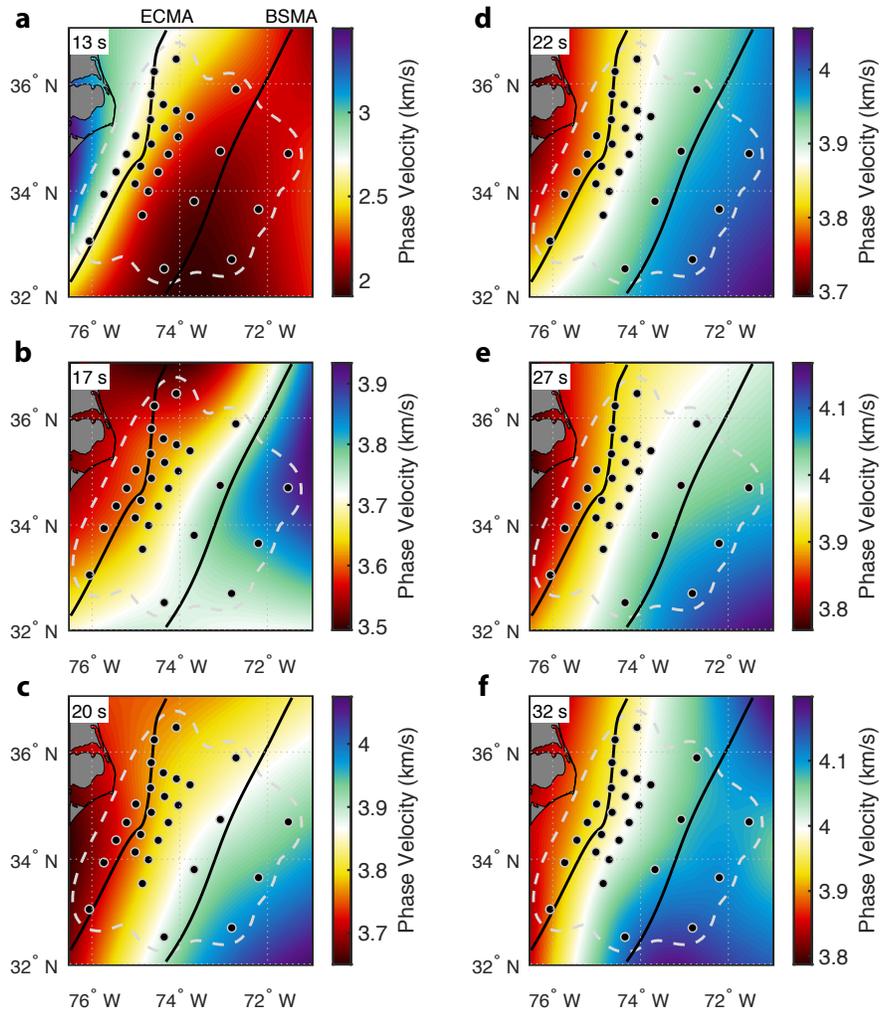


Figure 8: (a–f) Isotropic phase velocity maps for six periods ranging from 13–32 s. Black lines mark the East Coast Magnetic Anomaly (ECMA) and Blake Spur Magnetic Anomaly (BSMA). The dashed grey contour indicates the limit of resolution based on a value of 0.1 on the diagonals of the frequency-averaged resolution matrix.

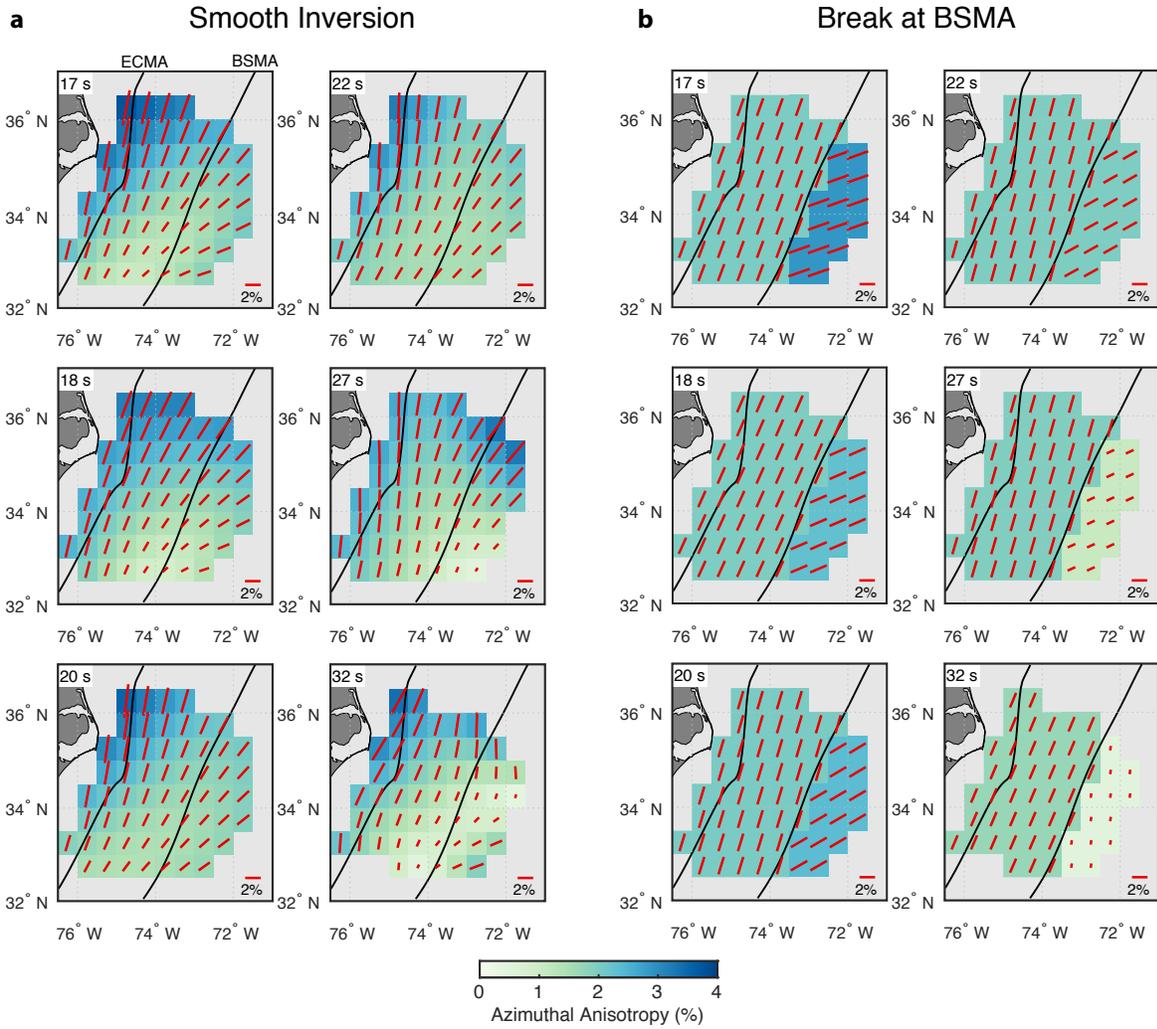


Figure 9: 2-D azimuthal anisotropy maps for periods ranging from 17–32 s for (a) a smoothly varying inversion and (b) an inversion that minimizes the first spatial derivative with a break at the Blake Spur Magnetic Anomaly (BSMA). Red sticks show anisotropy fast azimuths and their lengths scale with anisotropy magnitude. Black lines indicate the prominent magnetic anomalies.

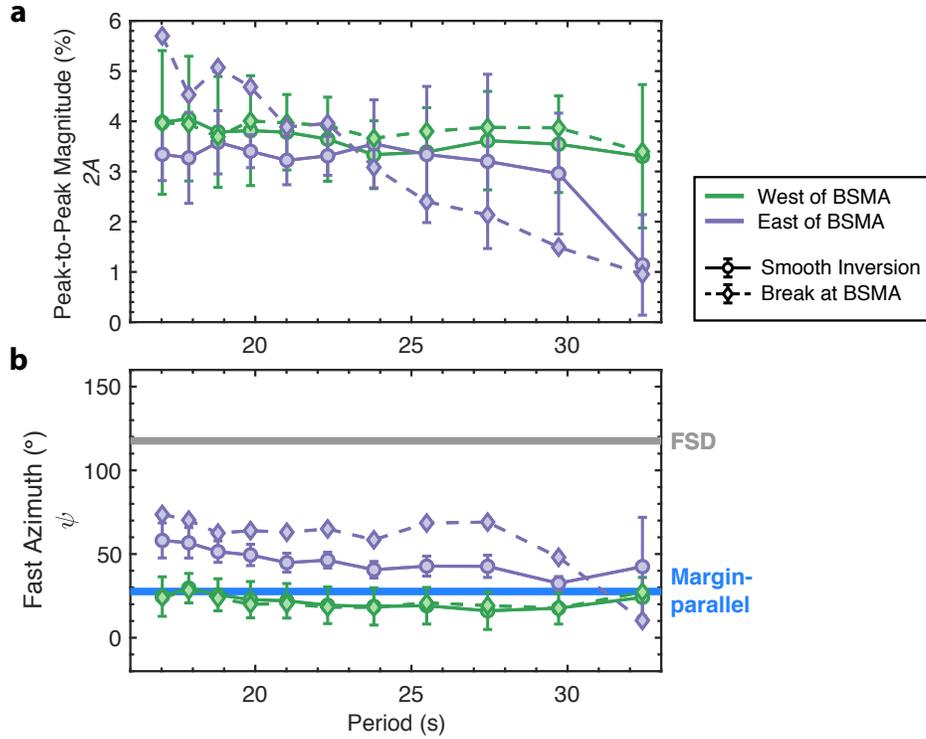


Figure 10: (a) Anisotropy peak-to-peak strength and (b) fast propagation azimuth for the two model solutions in Figure 9 for regional averages representing west of the BSMA (green) and east of the BSMA (purple). Symbols show the mean and 2σ variations for the corresponding regions. Thin solid and dashed lines correspond to Figure 9a and 9b, respectively. Thick grey and blue lines mark the fossil-spreading direction (FSD) and margin-parallel direction.

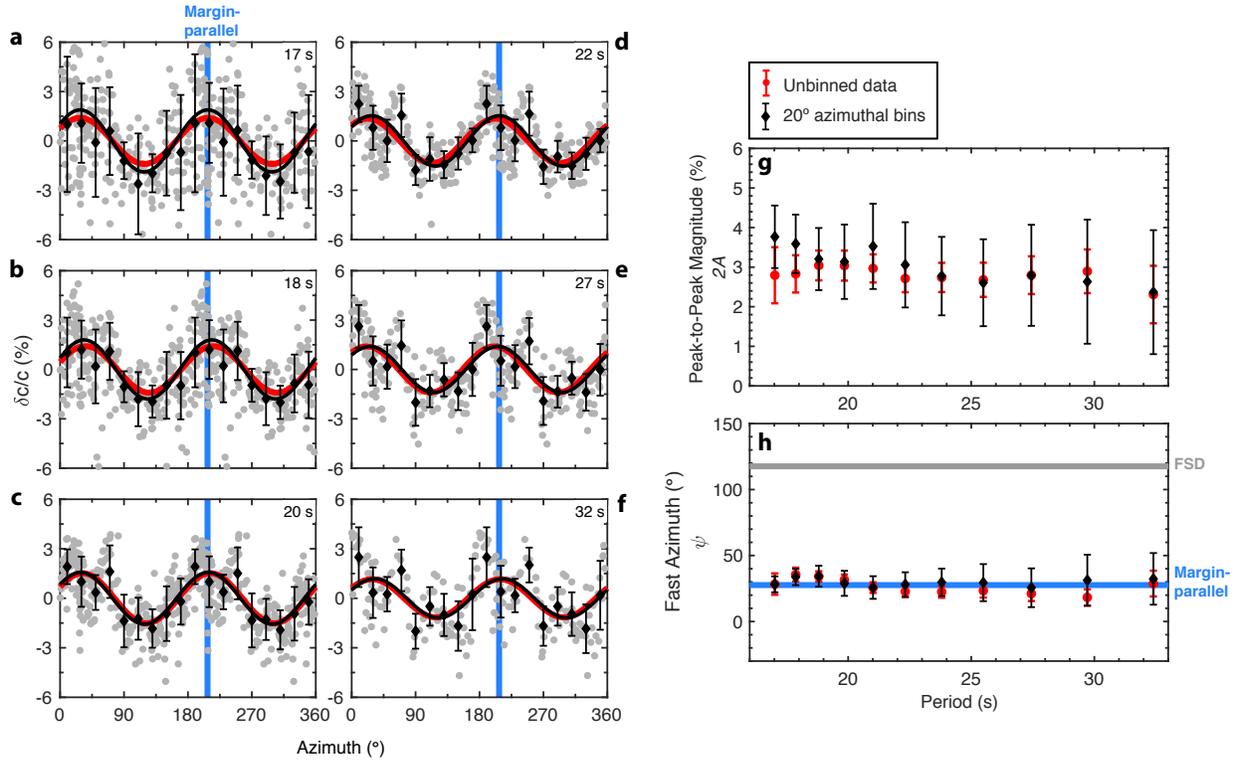


Figure 11: (a–f) Array-averaged 1-D azimuthal anisotropy ranging from 17–32 s period. Grey circles show interstation phase velocity deviations from the path-averaged isotropic values, and black diamonds show the mean and standard deviation for 20° azimuthal bins. The 2θ fit to the individual measurements is shown in red and the fit to the binned data is in black. The margin-parallel direction is indicated by the vertical blue dashed lines. (g) Anisotropy peak-to-peak strength and (h) fast propagation azimuth for the unbinned data in red and binned data in black. Error bars show 2- σ uncertainties. Thick grey and blue dashed lines mark the fossil-spreading direction (FSD) and margin-parallel direction, respectively.

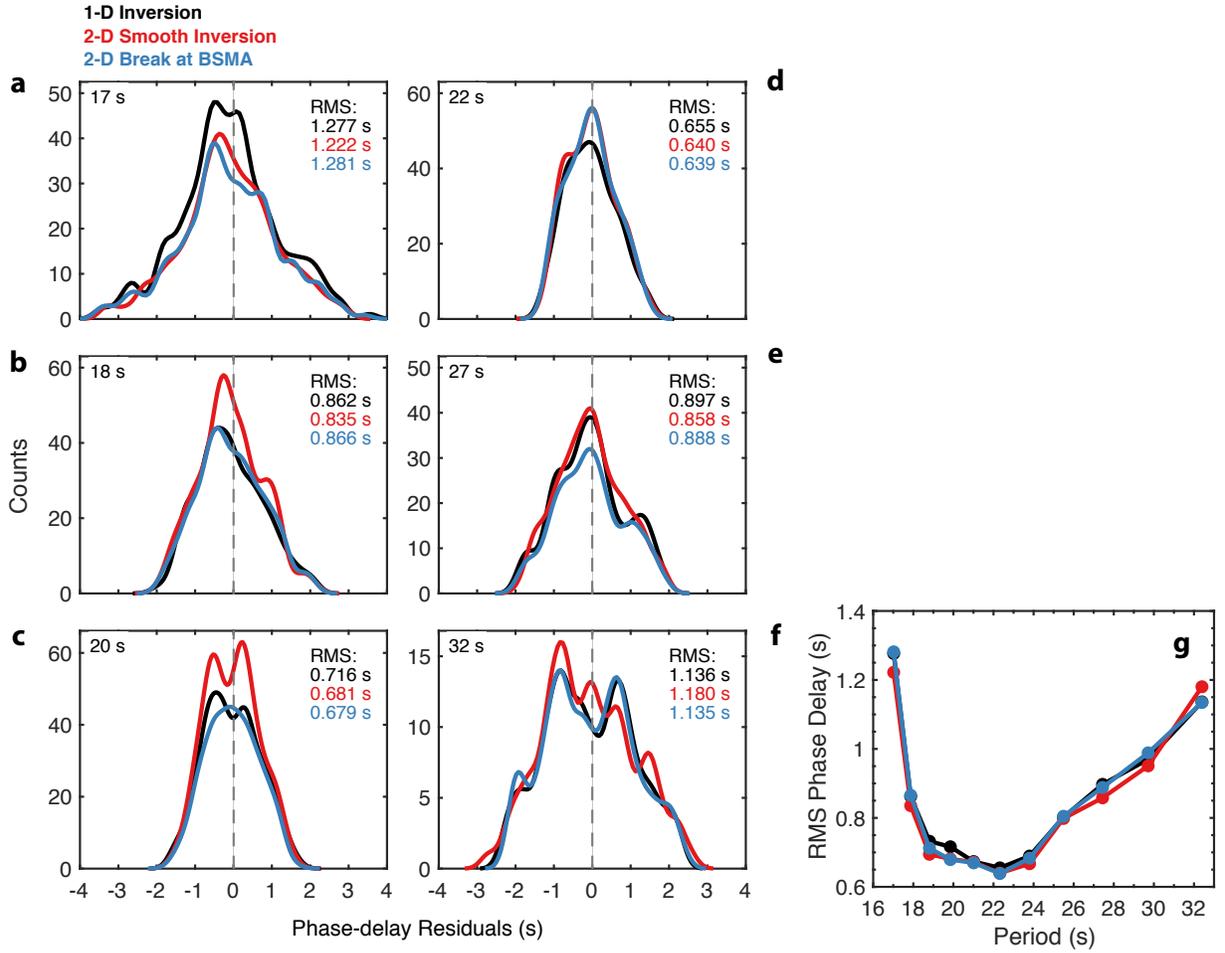


Figure 12: Comparison of phase-delay residuals and root-mean-square (RMS) misfits for three inversion types. (a–f) Smoothed histograms of phase-delay residuals for the 1-D inversion from Figure 11 (black), the 2-D smooth inversion from Figure 9a (red), and the 2-D inversion with the break at the BSMA Figure 9b (blue) for periods ranging from 17–32 s. (g) RMS phase delays for each of the three inversions.

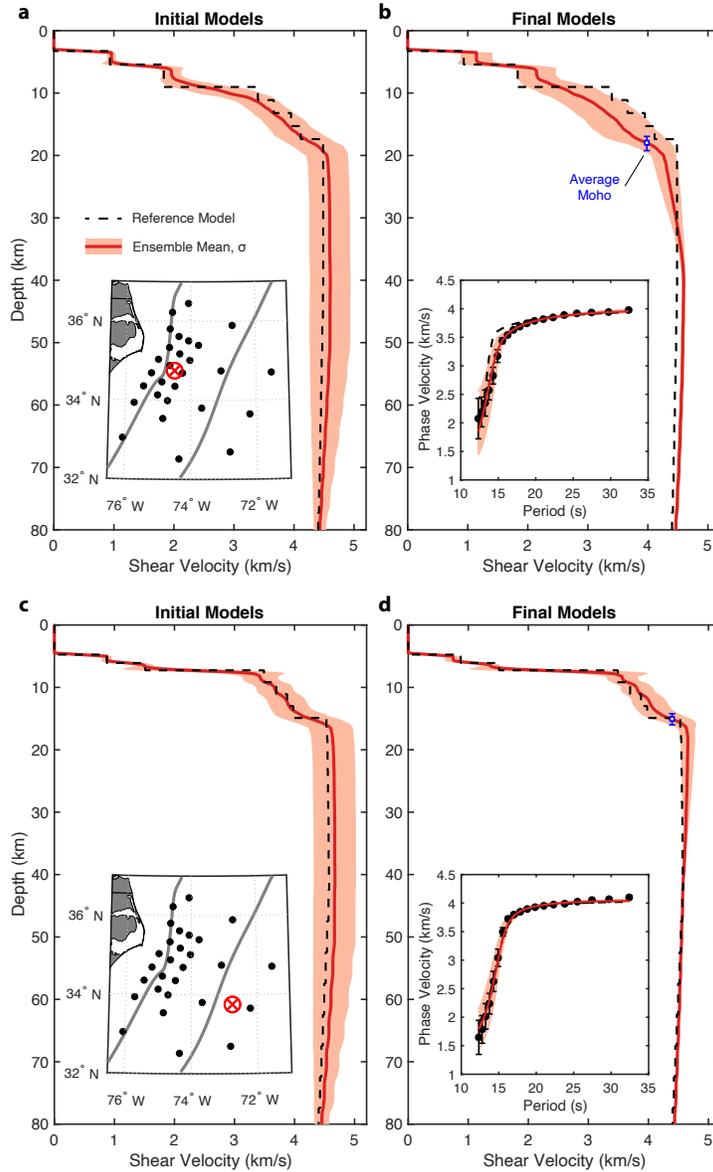


Figure 13: Representative 1-D V_S profiles from the Monte Carlo inversion for locations on the western and eastern edges of the array. (a) Starting model ensemble mean (red) and one standard deviation (σ , light shading). Reference model shown by black dashed line. The profile location is indicated by the red crosshair in the inset map. (b) Final model ensemble mean and standard deviation for all models with a data misfit of $\chi^2 \leq 1.75$. The mean and standard deviation Moho depth of acceptable models is shown by the blue symbol. The inset shows the dispersion data and 1- σ uncertainties in black, mean and range of model fits in red, and reference model prediction (black dashed). (c,d) Same as a), b) but for a location on the eastern edge of the array.

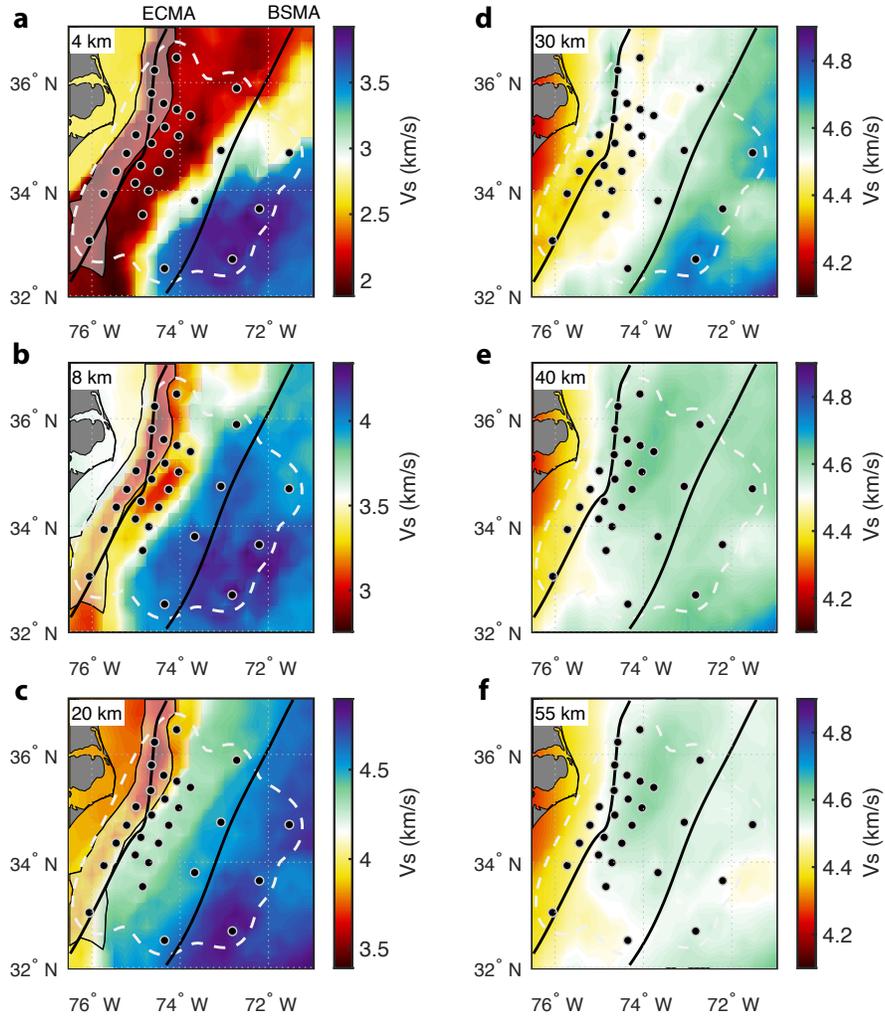


Figure 14: Horizontal slices through the mean 3-D shear velocity model. (a,b) Crustal slices at 4 km (upper crust) and 8 km (lower crust) below the seafloor (or free surface). (c-f) Mantle slices at 20, 30, 40, and 55 km depth. The white dashed line marks the limit of resolution. Regions of the model near the shelf where water depth changes rapidly (>2 km/ $^\circ$) are masked in grey, as these regions may contain biased slow velocity estimates (see main text). Major magnetic anomalies, ECMA and BSMA, are indicated by thick black lines.

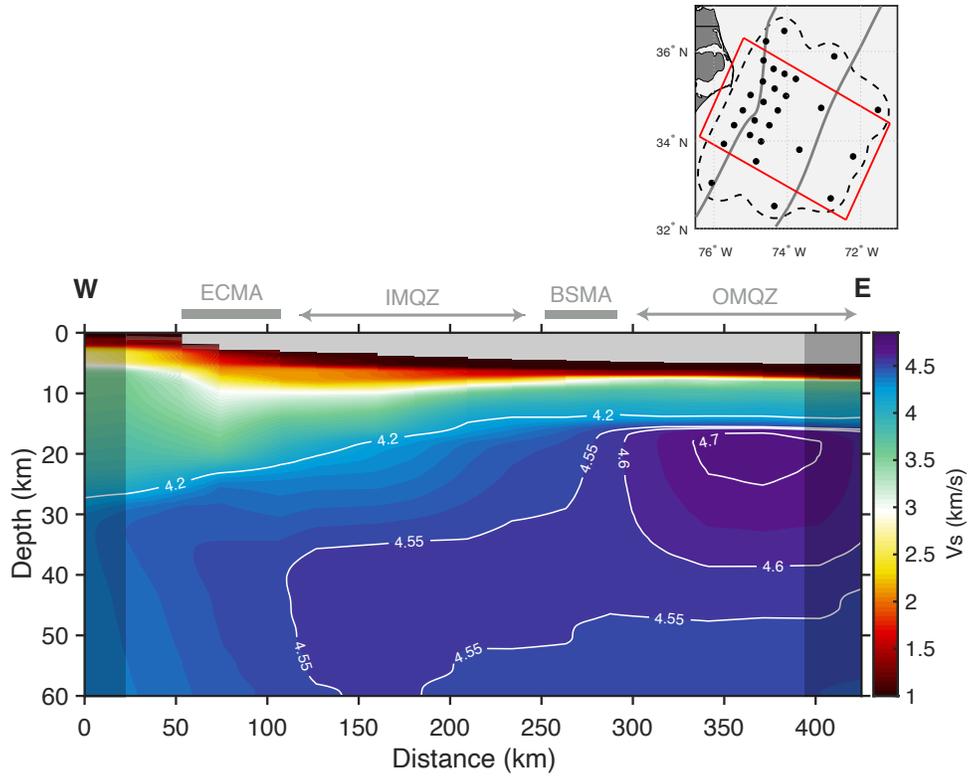


Figure 15: Average margin-perpendicular V_S structure. The region included in the average is shown by the red box in the upper map. The extent of the ECMA and BSMA are indicated by the grey bars. IMQZ = Inner magnetic quiet zone; OMQZ = Outer magnetic quiet zone.

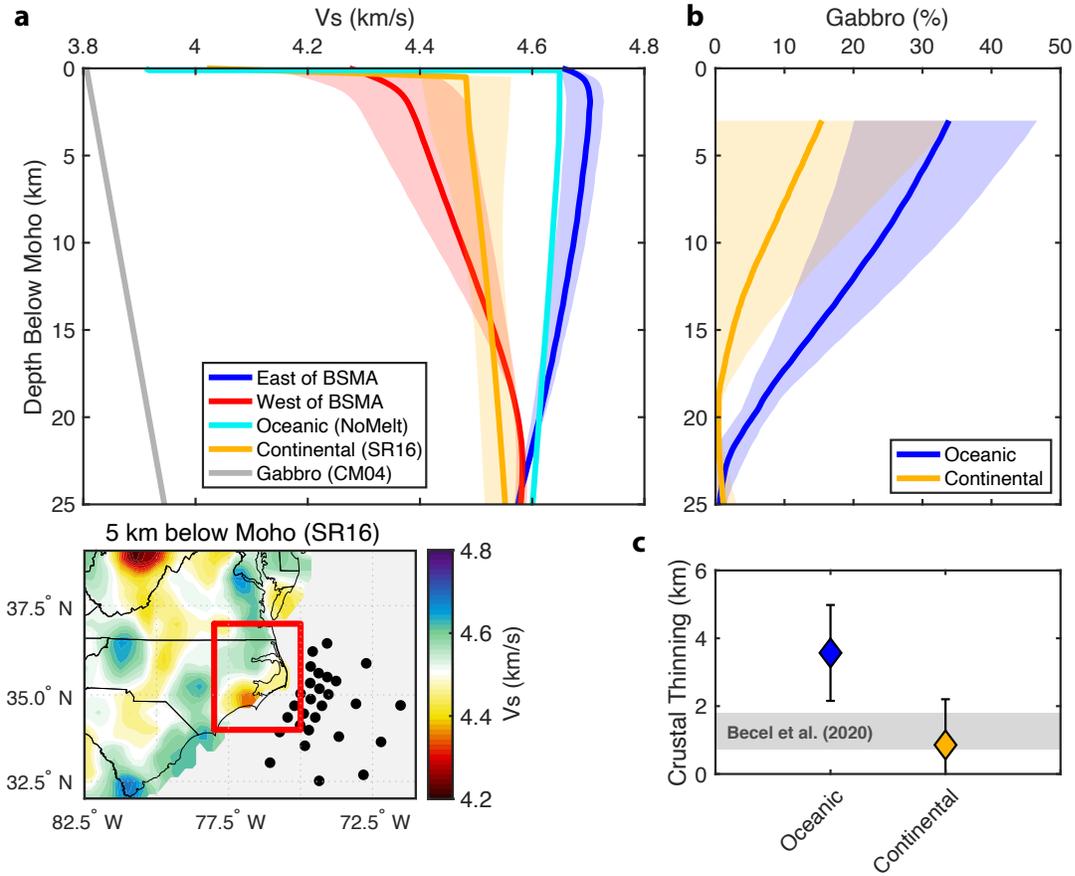


Figure 16: Comparing average V_S across the BSMA and evaluating the possible influence of gabbro inclusions. (a) Average shear velocity profiles for west (red) and east (blue) of the BSMA. Reference continental (*Shen and Ritzwoller, 2016*) (SR16) and oceanic (*Russell et al., 2019*) (NoMelt) models are shown in orange and cyan, respectively. The continental reference model is determined by averaging V_S within the red boxed region of SR16 shown in the lower left map. Gabbro V_S estimated after *Carlson and Miller (2004)* (CM04) assuming $V_P/V_S = 1.8$ is shown in grey (see text for details). (b) Percentage of gabbro added to the continental (orange) and oceanic (blue) reference profiles in order to match V_S west of the BSMA in a). (c) Equivalent crustal thinning west of the BSMA calculated by integrating b) assuming the crust is composed of 100% gabbro. The grey band represents the observed change in crustal thickness across the BSMA from *Bécel et al. (2020)*, *Shuck et al. (2019)*. Crustal thinning is consistent with up to $\sim 15\%$ gabbro added to continental mantle, but is inconsistent with oceanic mantle.

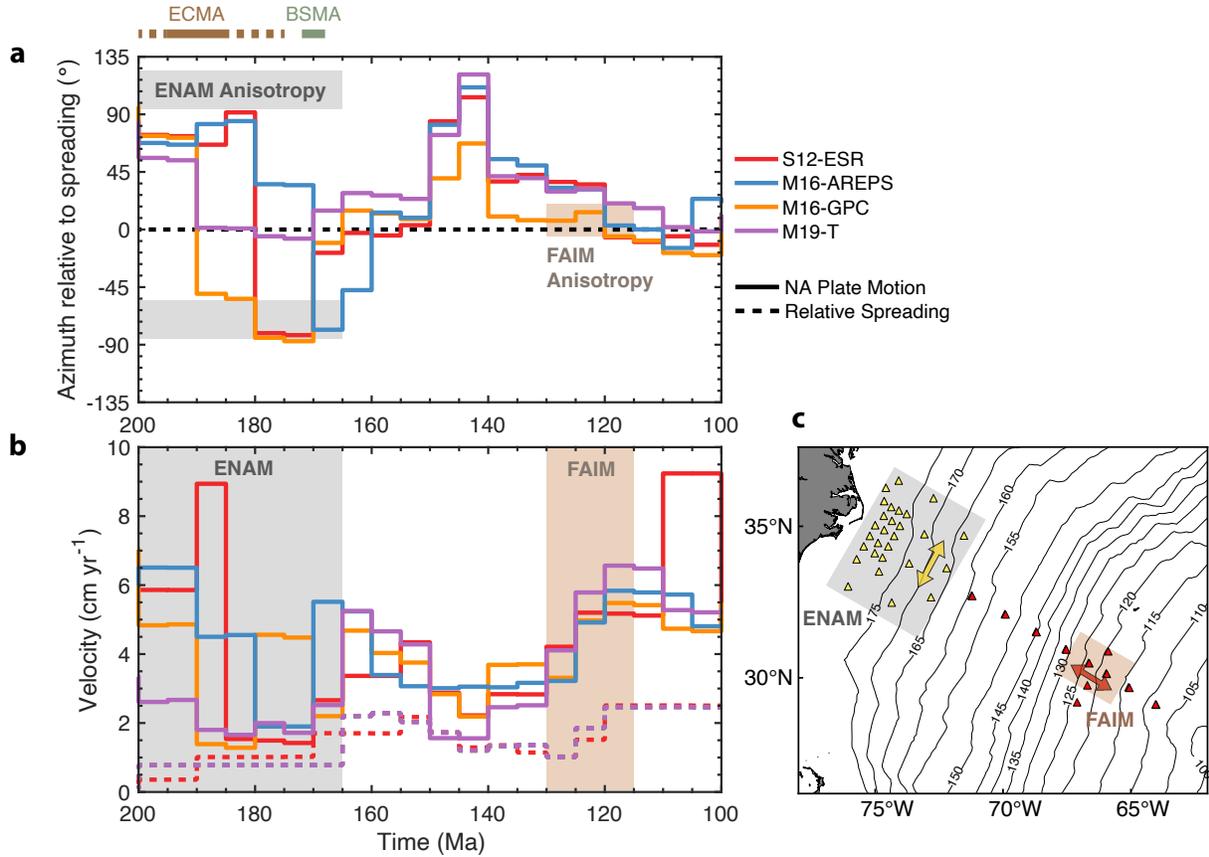


Figure 17: Comparison of anisotropy observations to paleo plate motions from 100–200 Ma from plate reconstruction models. Solid lines indicate absolute plate motion of North American (NA), while dashed lines indicate relative spreading motion of NA with respect to Africa. (a) Azimuth of NA plate motion relative to the spreading direction. Line colors correspond to four different plate reconstruction models (S12-ESR (*Seton et al.*, 2012); M16-AREPS (*Müller et al.*, 2016); M16-GPC (*Matthews et al.*, 2016); M19-T (*Müller et al.*, 2019)). Grey and tan regions mark the anisotropy fast azimuths at ENAM from this study and the Far-offset Active-source Imaging of the Mantle (FAIM) experiment (*Gaherty et al.*, 2004), respectively. Approximate timing of the ECMA and BSMA emplacements are indicated along the top. (b) NA plate speed (solid) and half-spreading rate of NA with respect to Africa (dashed; M16-GPC, M16-AREPS, and M19-T overlap one another). Grey and tan regions indicate the approximate range of seafloor ages at ENAM and FAIM, respectively. (c) Map of ENAM (yellow) and FAIM (red) OBS with shading that indicates approximate seafloor locations of anisotropy observations. Seafloor age contours from *Müller et al.* (2008) are shown in black in increments of 5 Myr.

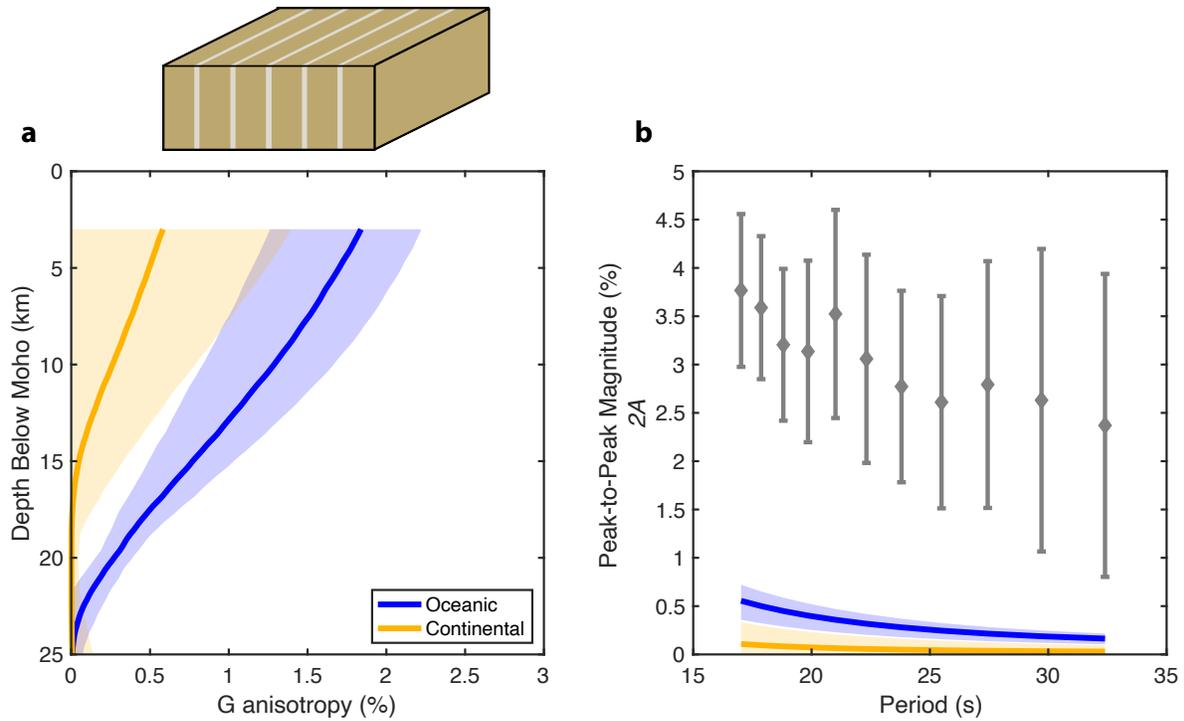


Figure 18: Quantification of shape-preferred anisotropy due to vertical gabbro channels. (a) Shear-wave azimuthal anisotropy magnitude (G) for a vertically layered structure consisting of alternating layers of gabbro and continental/oceanic mantle in the proportions shown in Figure 16b. The calculation uses *Backus* (1962) to estimate the horizontal layered long-wavelength equivalent elastic tensor, which is then rotated 90° to simulate vertical layers. G anisotropy is calculated from the elastic tensor at each depth following *Montagner and Nataf* (1986). Colors are as in Figure 16. Estimated V_P anisotropy is shown in Figure S8. (b) Predicted azimuthal anisotropy of Rayleigh-wave phase velocity for the profiles shown in a) using sensitivity kernels from Figure 7c. Grey symbols show the binned 1-D anisotropy observations from Figure 11g.

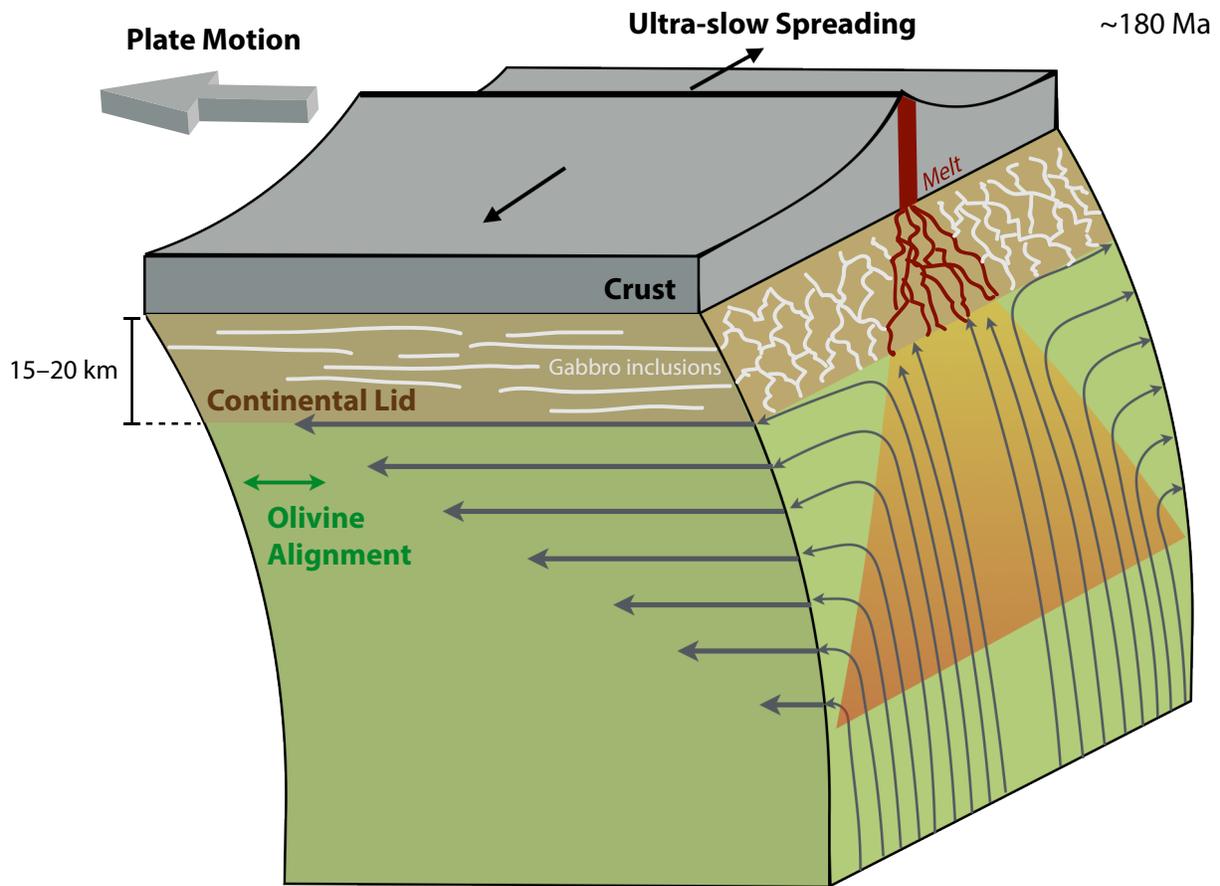


Figure 19: Schematic demonstrating our preferred interpretation of the structure and processes that occurred ~180 Ma prior to the completion of Pangea breakup. The 15–20 km thick continental lid truncated the top of the melting region at the time. As melt percolated through the lid forming the proto-oceanic crust, up to ~15% gabbro was crystallized within melt channels. Due to the slow spreading rate, corner flow at the ridge was weak relative to plate motion, resulting in olivine LPO oriented sub-parallel to the plate-motion direction in the underlying asthenosphere, which was subsequently frozen into the cooling oceanic mantle lithosphere. The combination of margin-parallel LPO in the oceanic lithosphere and significantly weaker shape-preferred orientation of the frozen gabbro channels in the lid likely contribute to the observed margin-parallel anisotropy.

SUPPLEMENTARY FIGURES

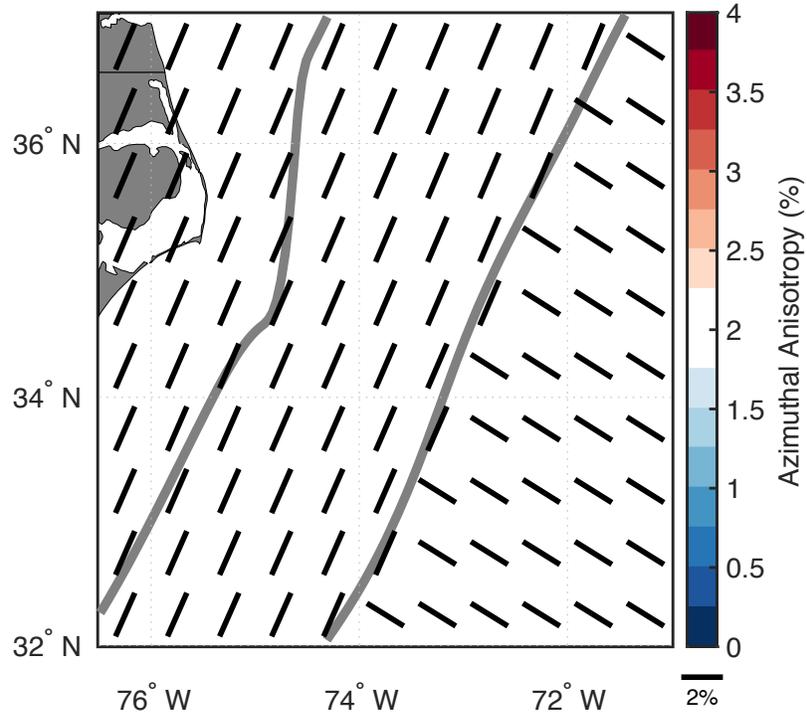


Figure S1: Input anisotropy model used in the synthetic tests in Figures S2–S4. Black sticks show the anisotropy directions and their lengths scale with anisotropy magnitude. Anisotropy strength is held constant at 2% across the model. West of the BSMA anisotropy is margin-parallel, and east of the BSMA it abruptly rotates 90° to FSD-parallel.

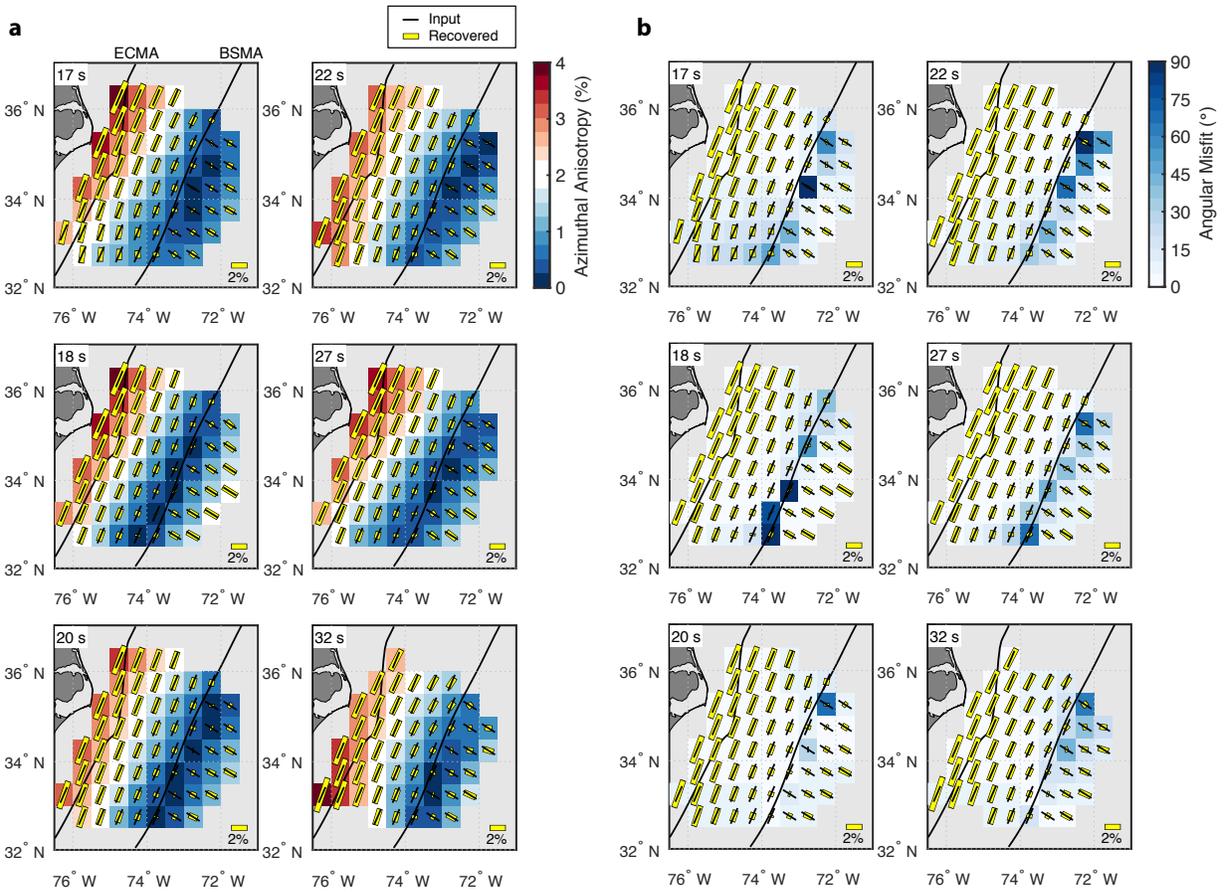


Figure S2: Synthetic anisotropy recovery test results for the smooth anisotropy model parameterization. The input model is shown in Figure S1. (a) Maps of recovered anisotropy magnitudes, where deviations from white indicate overshoot (red) and undershoot (blue) relative to the input model. (b) Absolute angular misfit between fast azimuths of the input and recovered model. Deviations from white indicate poor fast azimuth recovery. Black sticks represent in the input model and yellow bars indicate the recovered model at each frequency.

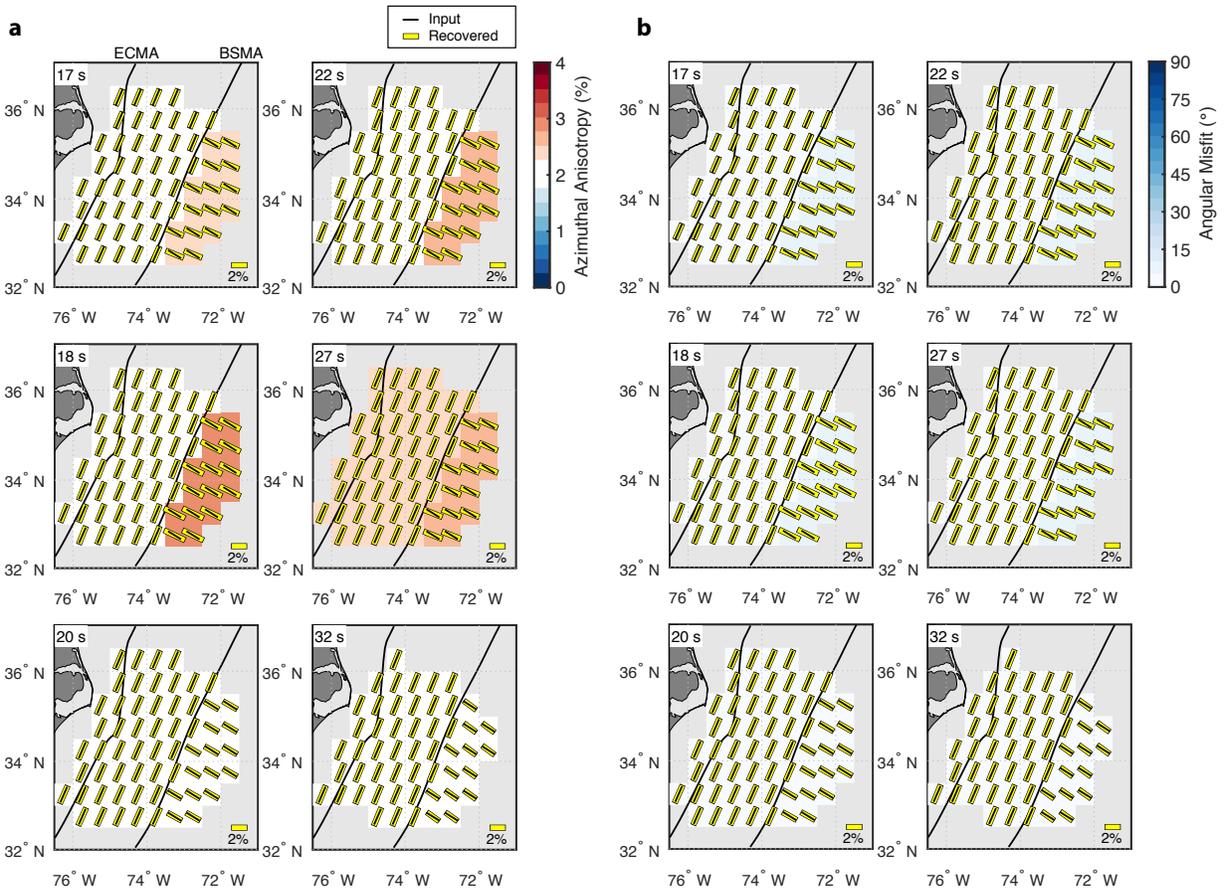


Figure S3: Same as Figure S2 but for the model parameterization with a break at the BSMA.

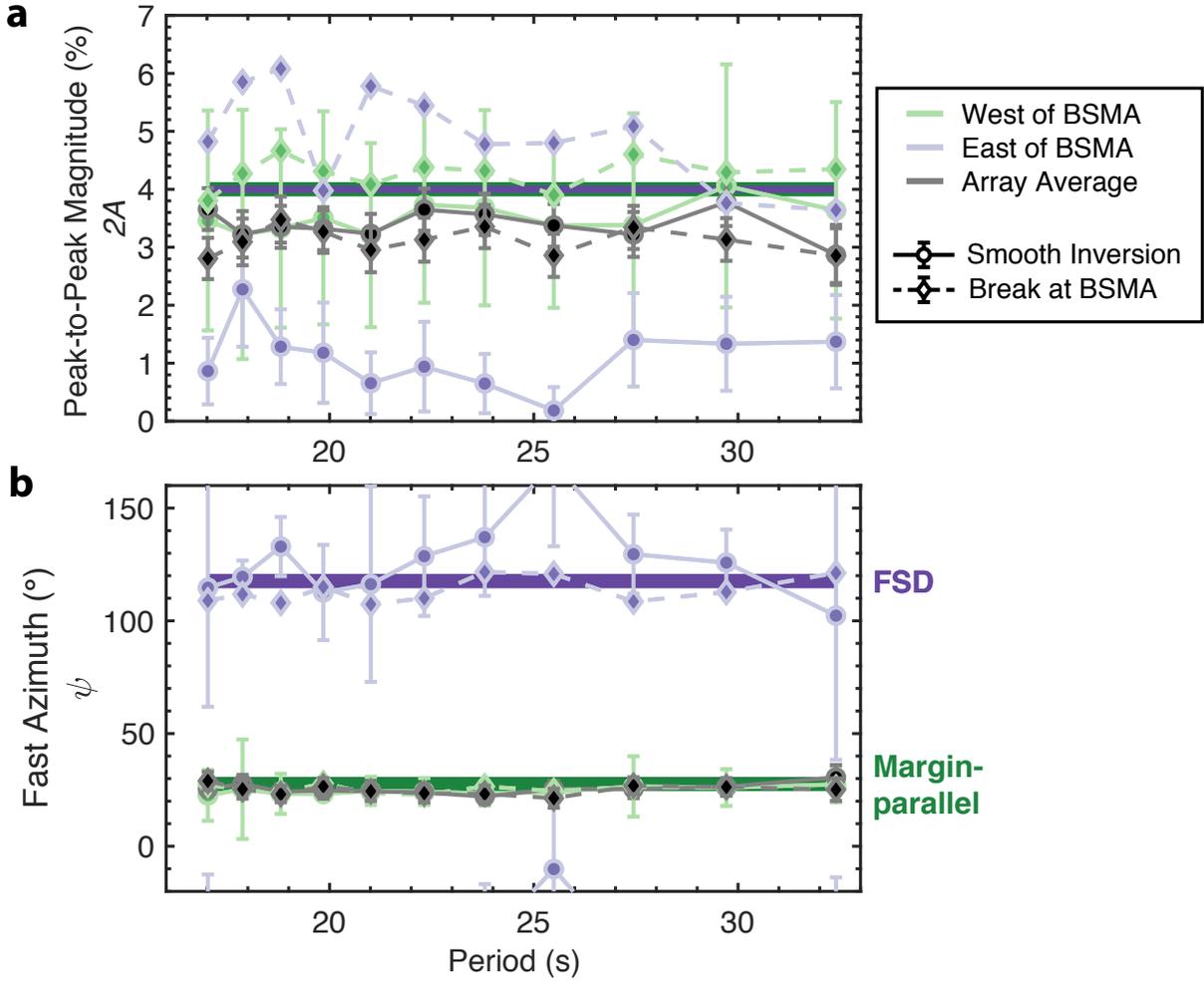


Figure S4: Summary of synthetic recovery tests. (a) Peak-to-peak anisotropy magnitude and (b) fast azimuths recovered from 17–32 s period. Thick bars indicate the input model values. Results for the smooth parameterization and break at BSMA are shown by solid and dashed lines, respectively. Regional averages for the recovered models are indicated by color: west of the BSMA (green), east of the BSMA (purple), and array average (grey).

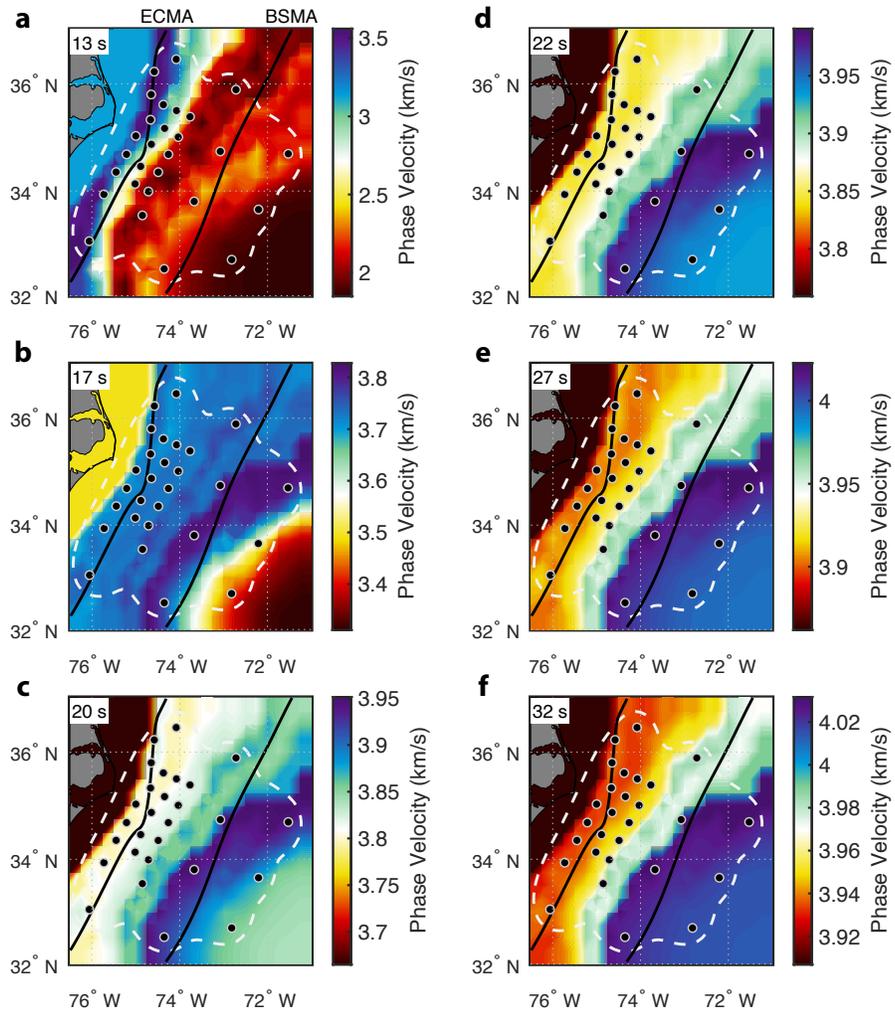


Figure S5: Starting model phase velocity maps for the same frequencies as in Figure 8 in the main text.

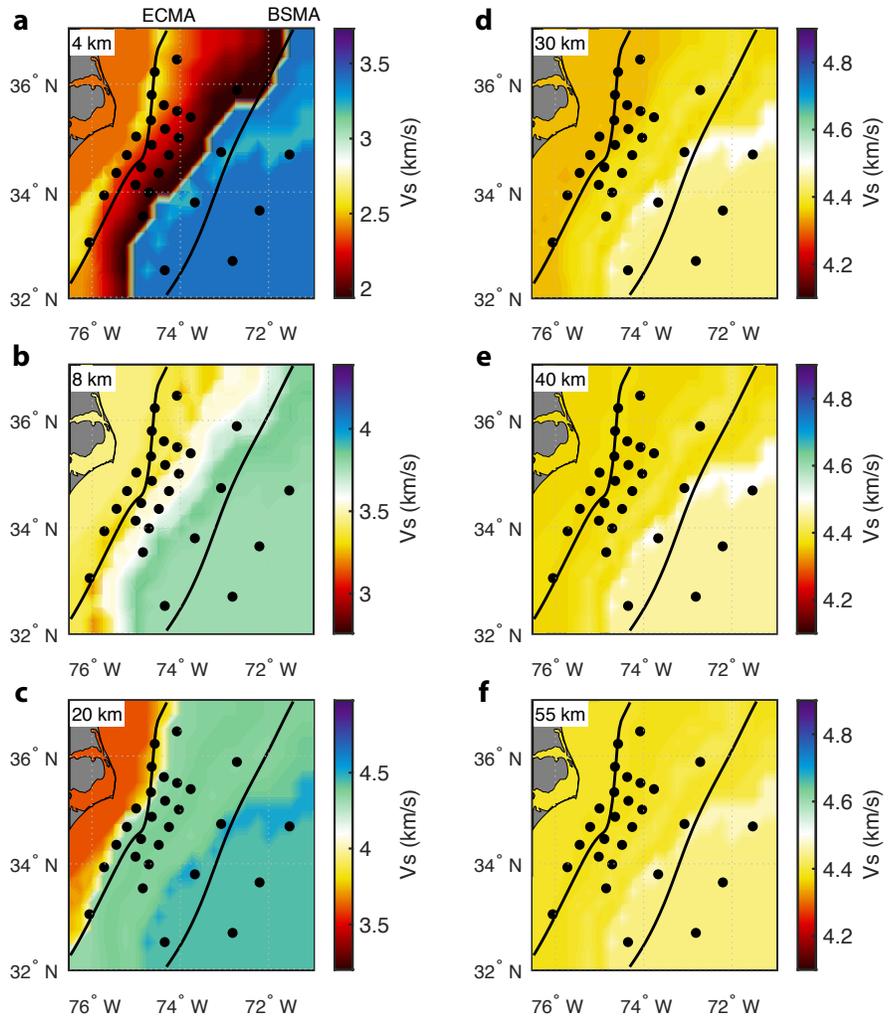


Figure S6: Horizontal slices through the reference V_S model, as in Figure 14 in the main text.

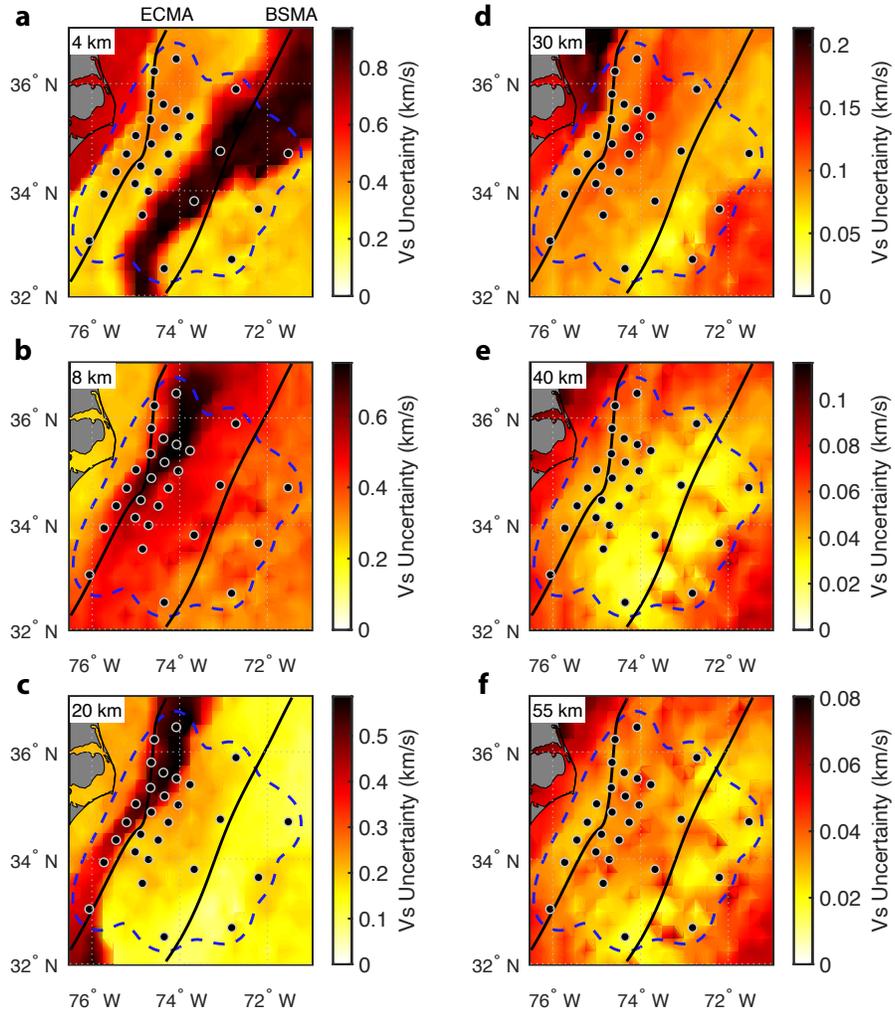


Figure S7: One standard deviation of the final V_S model ensemble with data misfit $\chi^2 < 1.75$ for the depth slices shown in Figure 14 in the main text.

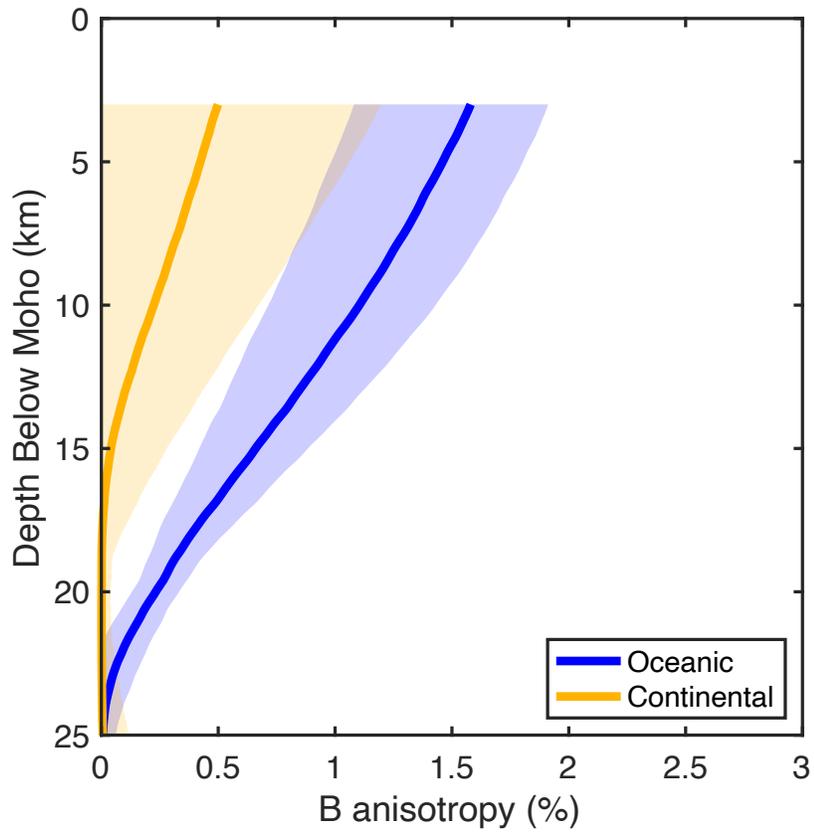


Figure S8: V_P azimuthal anisotropy magnitude (B) calculated for vertical gabbro channels analogous to Figure 18a.